



BECQUEREL  
PROJECT

Проект  
БЕККЕРЕЛЬ

Beryllium (Boron)

Clustering

Quest in

Relativistic Multifragmentation

<http://becquerel.jinr.ru>

# Электромагнитные взаимодействия тяжелых ионов

П.И. Зарубин (ОИЯИ)

## Electromagnetic dissociation of relativistic $^{18}\text{O}$ nuclei

D. L. Olson and B. L. Berman

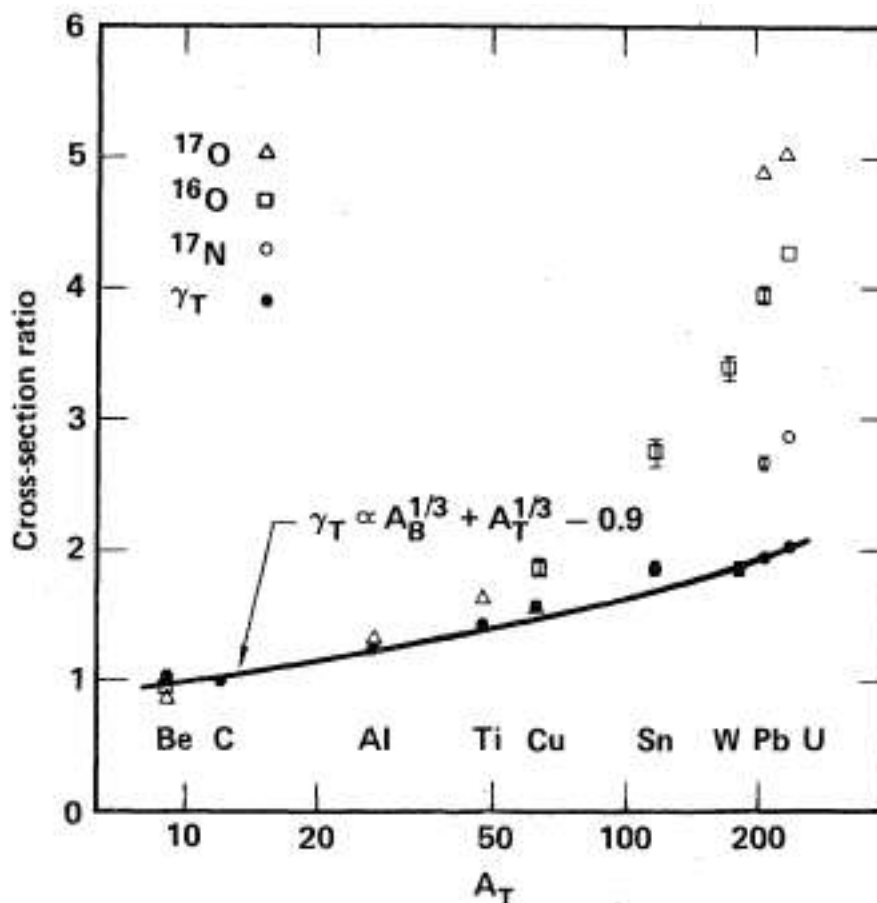
*Lawrence Livermore Laboratory, University of California, Livermore, California 94550*

D. E. Greiner, H. H. Heckman, P. J. Lindstrom, and G. D. Westfall

*Lawrence Berkeley Laboratory, University of California, Berkeley, California*

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*University of California Space Sciences Laboratory, Berkeley, California 94720*





# 10.7 A GeV Au

PHYSICAL REVIEW C 72, 048801 (2005)

## Multifragmentation reactions and properties of stellar matter at subnuclear densities

A. S. Botvina<sup>1</sup> and I. N. Mishustin<sup>2,3</sup>

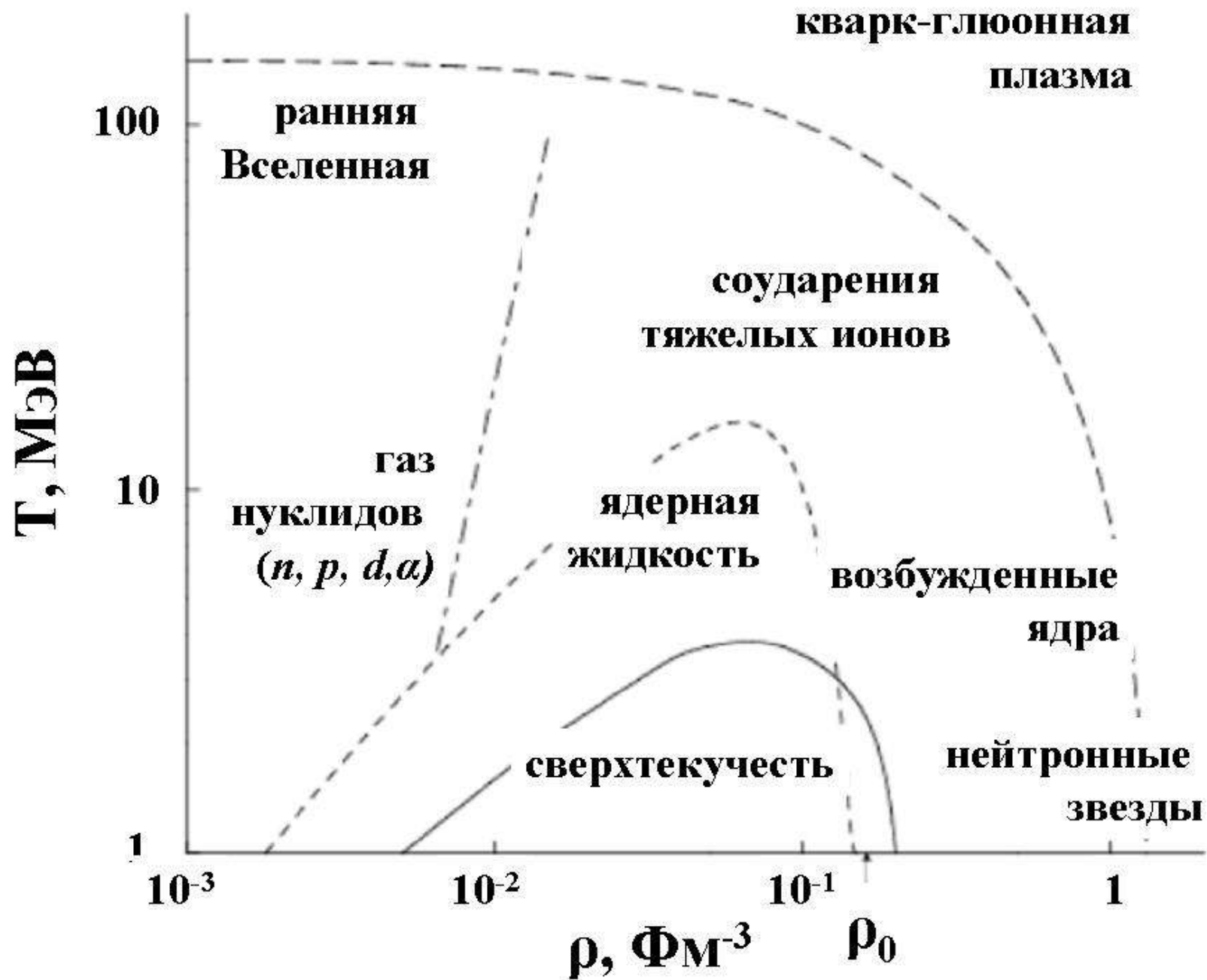
<sup>1</sup>*Institute for Nuclear Research, Russian Academy of Sciences, RU-117312 Moscow, Russia*

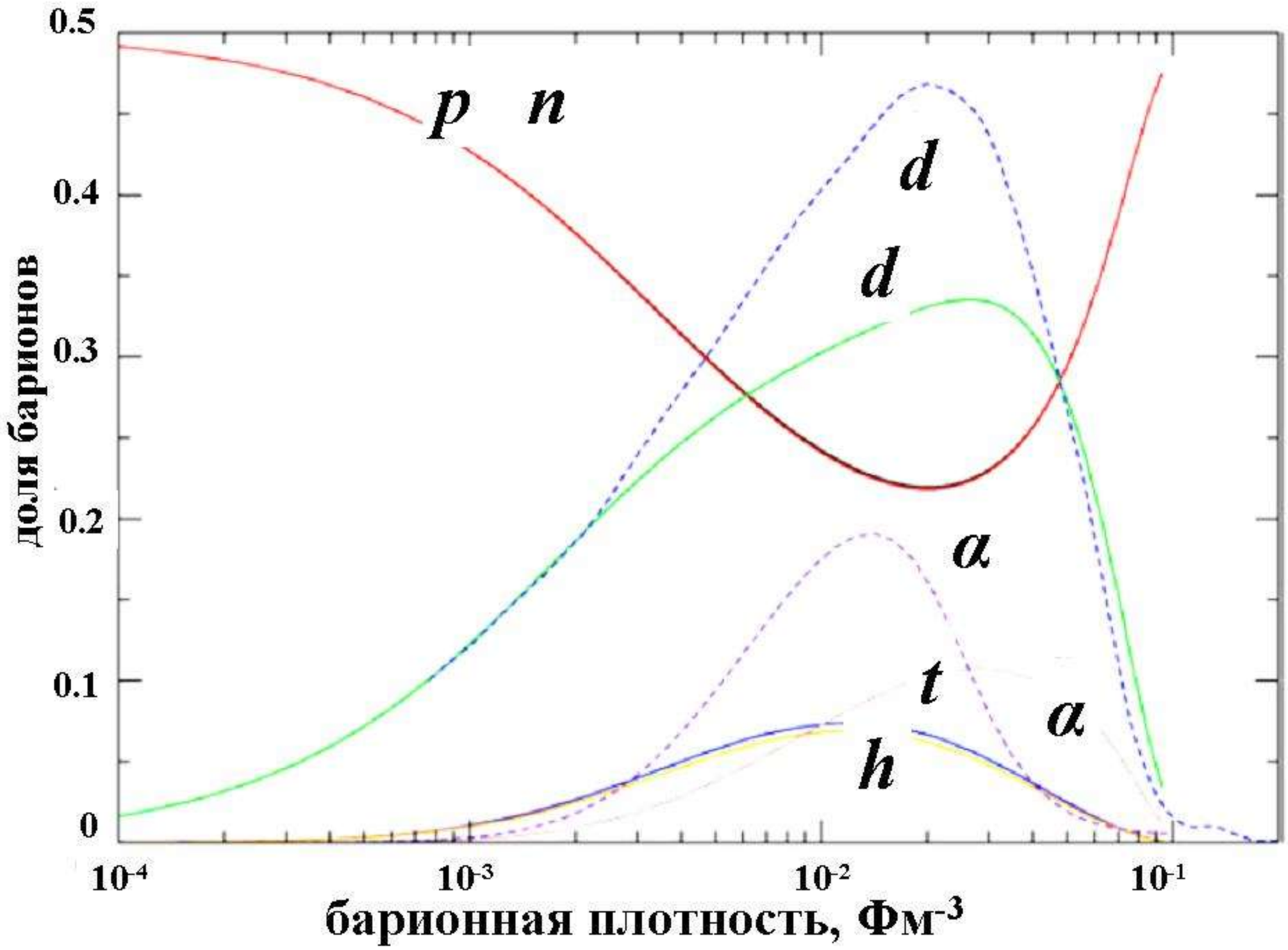
<sup>2</sup>*Frankfurt Institute for Advanced Studies, J.W. Goethe University, D-60438 Frankfurt am Main, Germany*

<sup>3</sup>*Kurchatov Institute, Russian Research Center, RU-123182 Moscow, Russia*

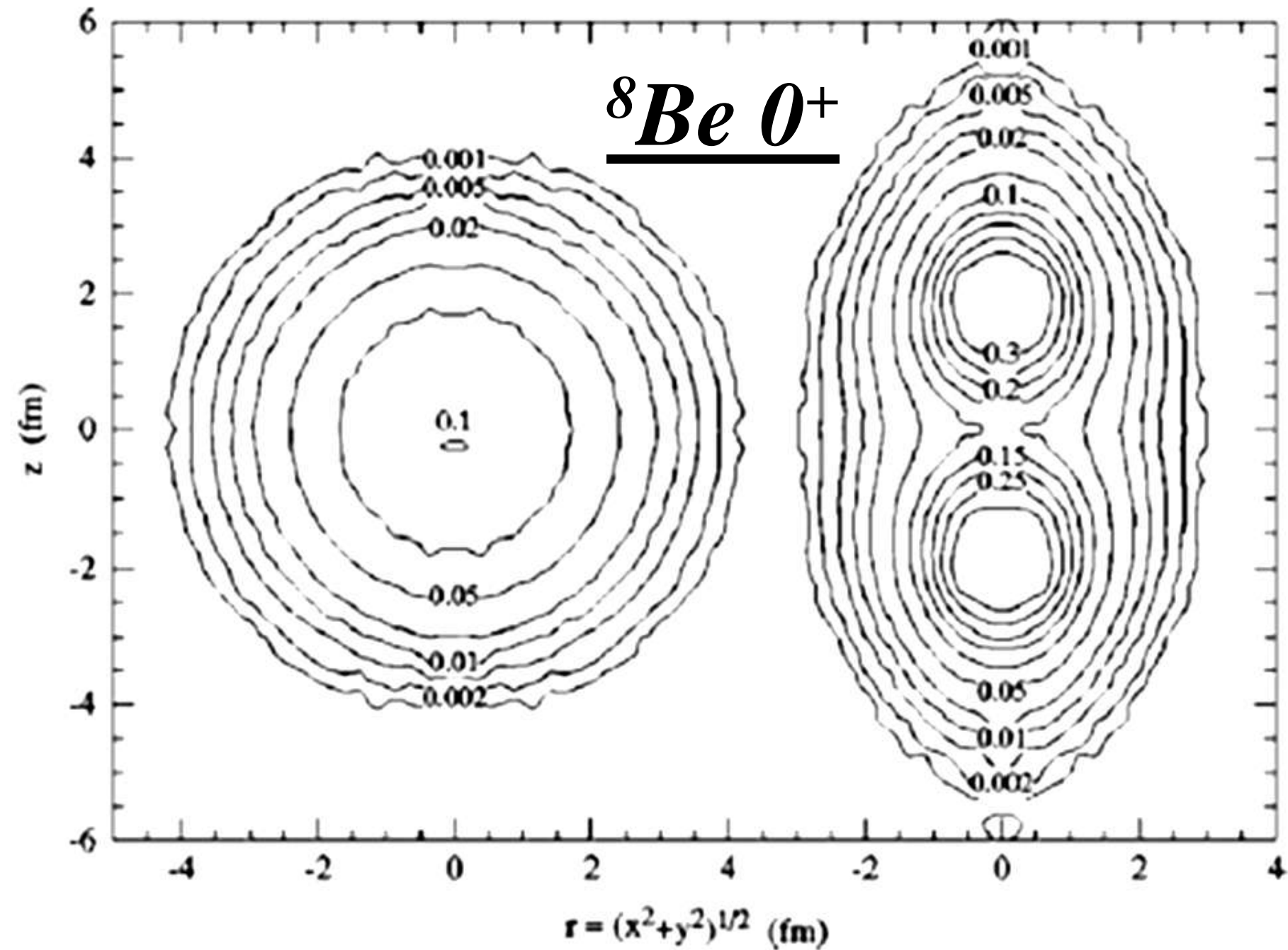
(Received 20 June 2005; published 24 October 2005)

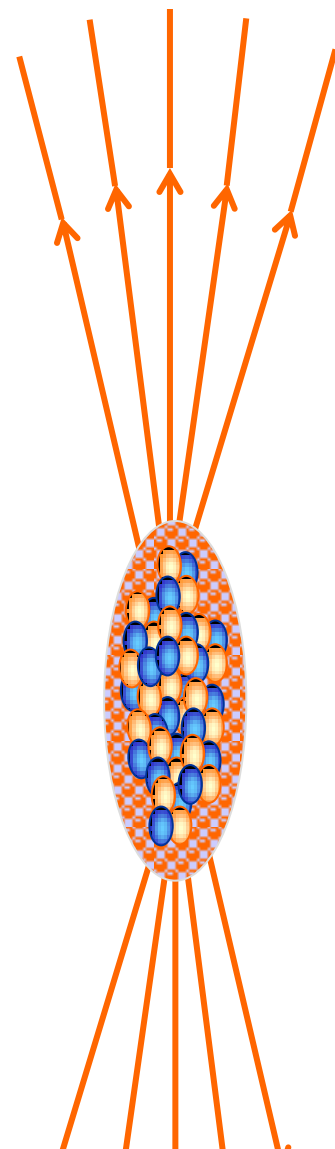
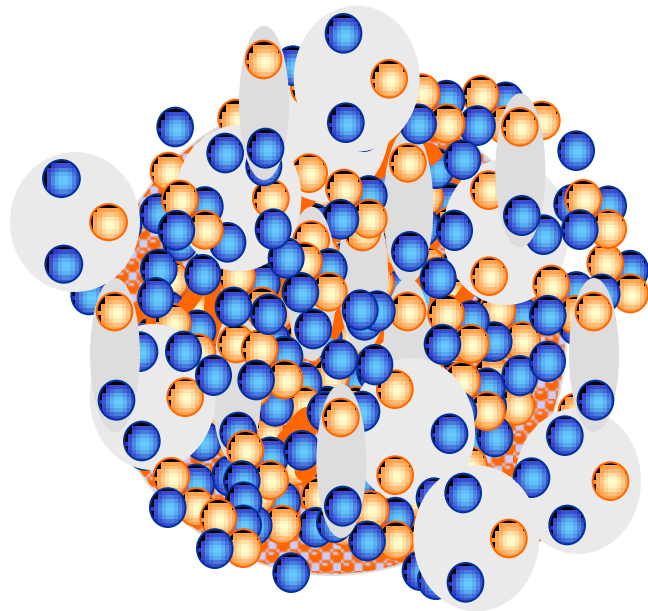
We point out the similarity of thermodynamic conditions reached in nuclear multifragmentation and in supernova explosions. We show that a statistical approach previously applied for nuclear multifragmentation reactions can also be used to describe the electroneutral stellar matter. Then properties of hot unstable nuclei extracted from the analysis of multifragmentation data can be used to determine a realistic nuclear composition of hot supernova matter.

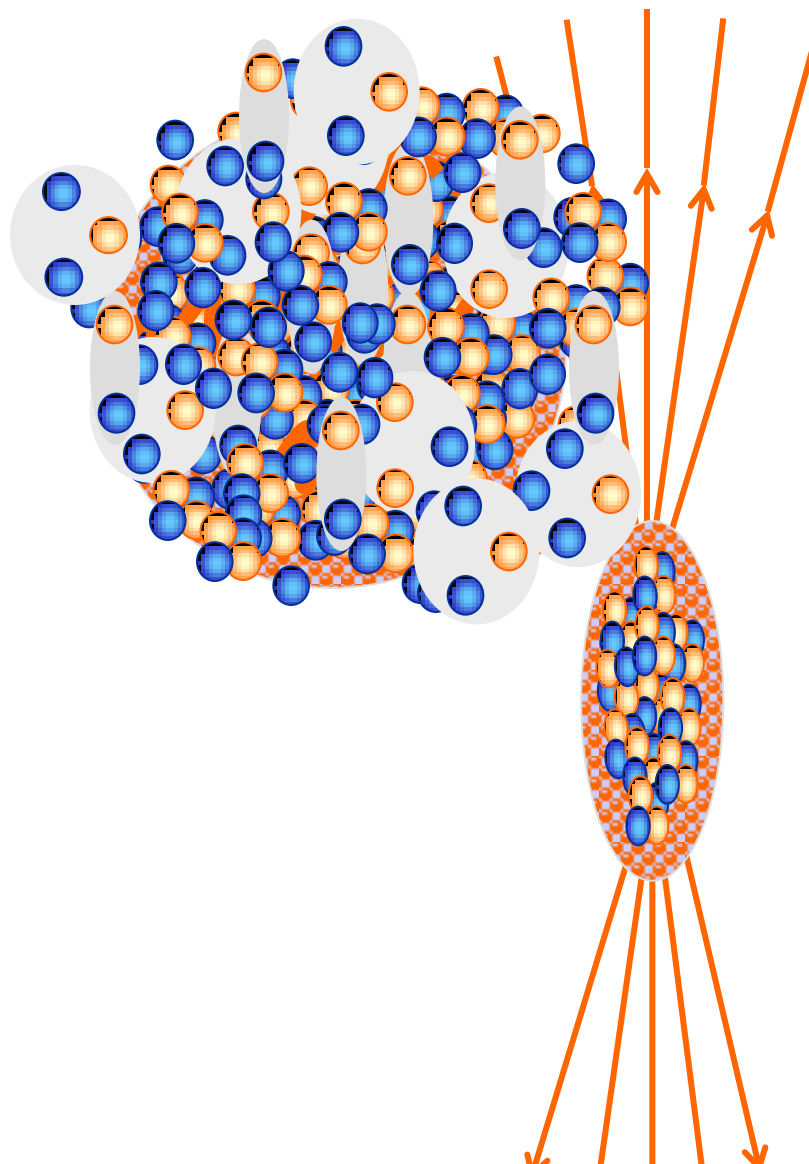




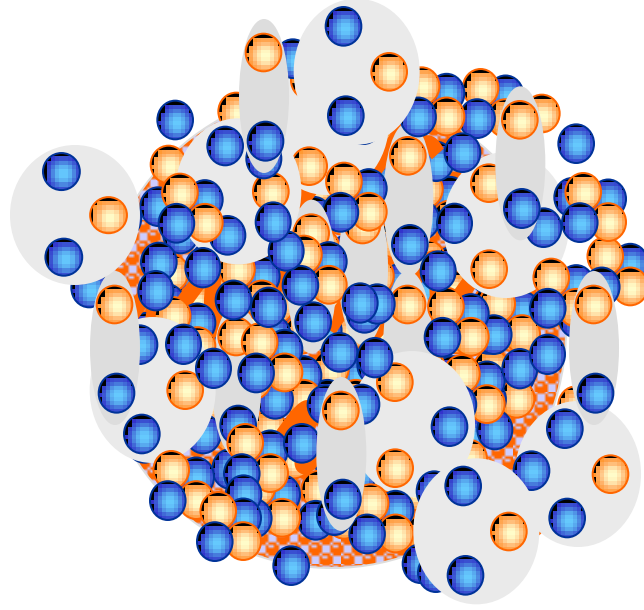
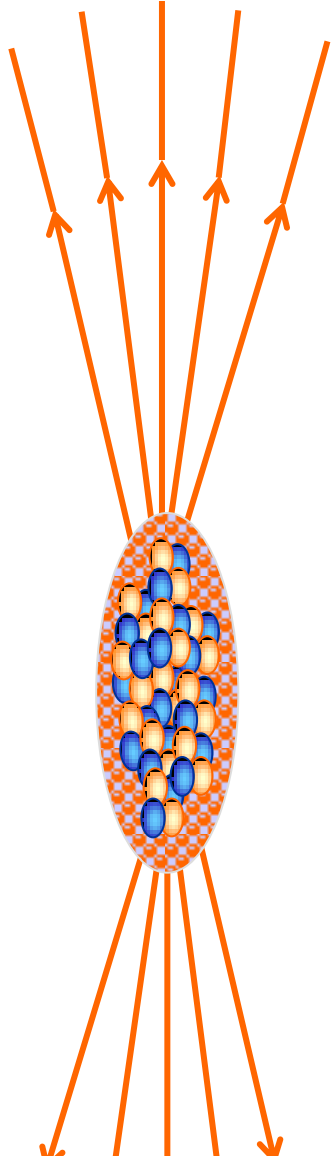
${}^8\text{Be } 0^+$

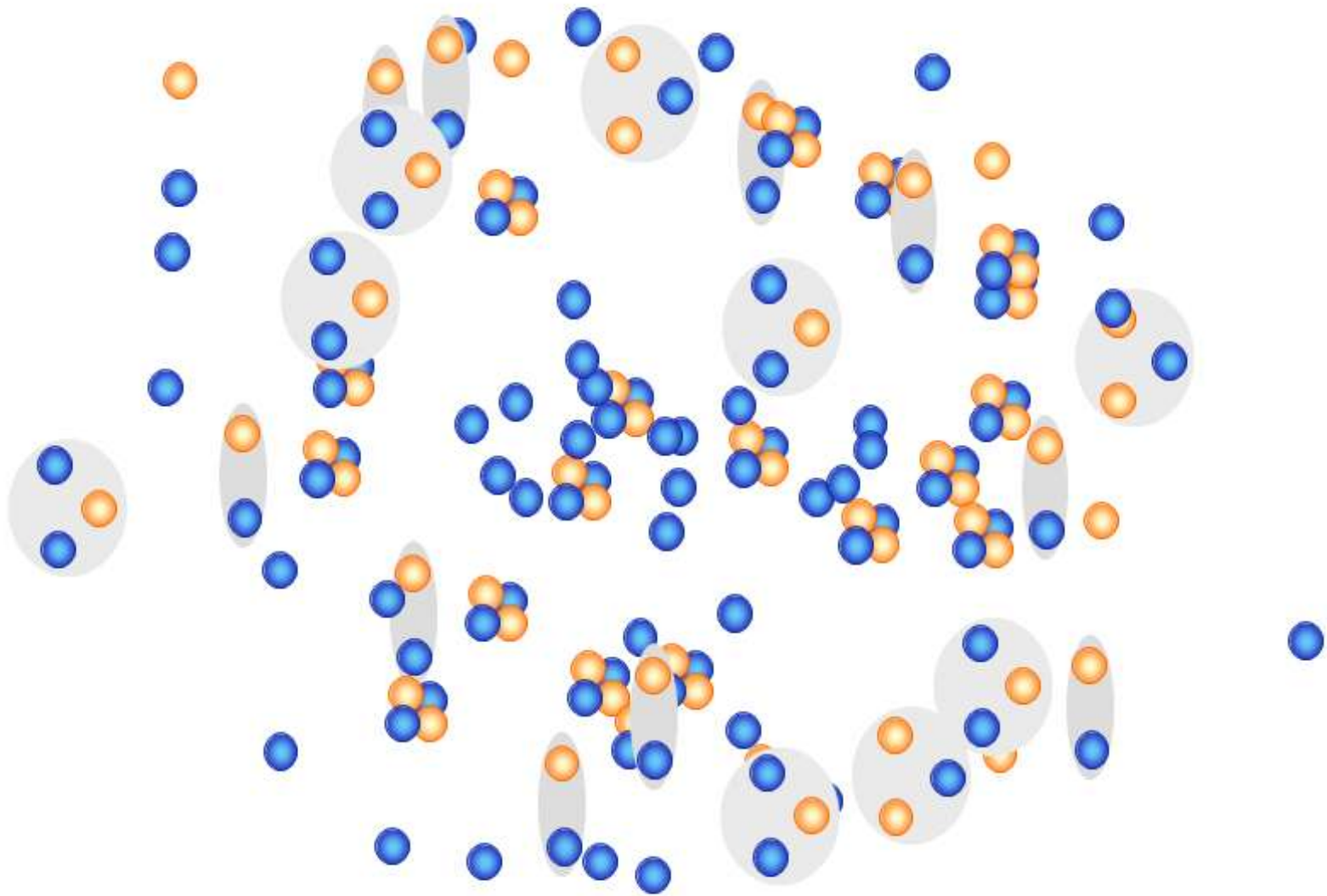












## Electromagnetic dissociation of relativistic heavy ions

W. J. Llope and P. Braun-Munzinger

*Department of Physics, State University of New York at Stony Brook, Stony Brook, New York 11794*

In particular, electromagnetic excitation of modes based on the nuclear giant dipole resonance (GDR) may lead to very exotic final states<sup>1,2</sup> in which neutrons oscillate against protons with a very large amplitude. The existence and decay mechanisms of such states is unknown at present. However, this electromagnetic process efficiently excites collective states so that little or no temperature is produced during the very short time scale (of order  $1 \text{ fm}/c$ ) of the collision. One may thus hope to use this type of reaction to search for fragile, weakly bound exotic states such as multineutron clusters which might be formed in the decay of the possibly strongly excited multi-GDR states.

## Electromagnetic dissociation of relativistic heavy ions

W. J. Llope and P. Braun-Munzinger

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(Received 12 January 1990)

A framework is developed for the quantitative analysis of the electromagnetic dissociation of relativistic nuclei. This includes treatment of multiple excitations of the giant dipole resonance, coupled with calculations of the fragmentation probabilities in the framework of the statistical model.

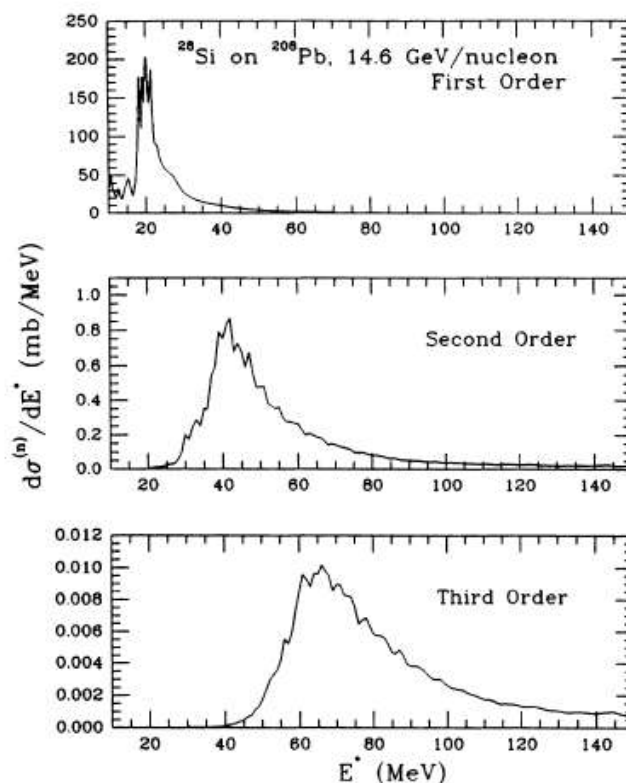
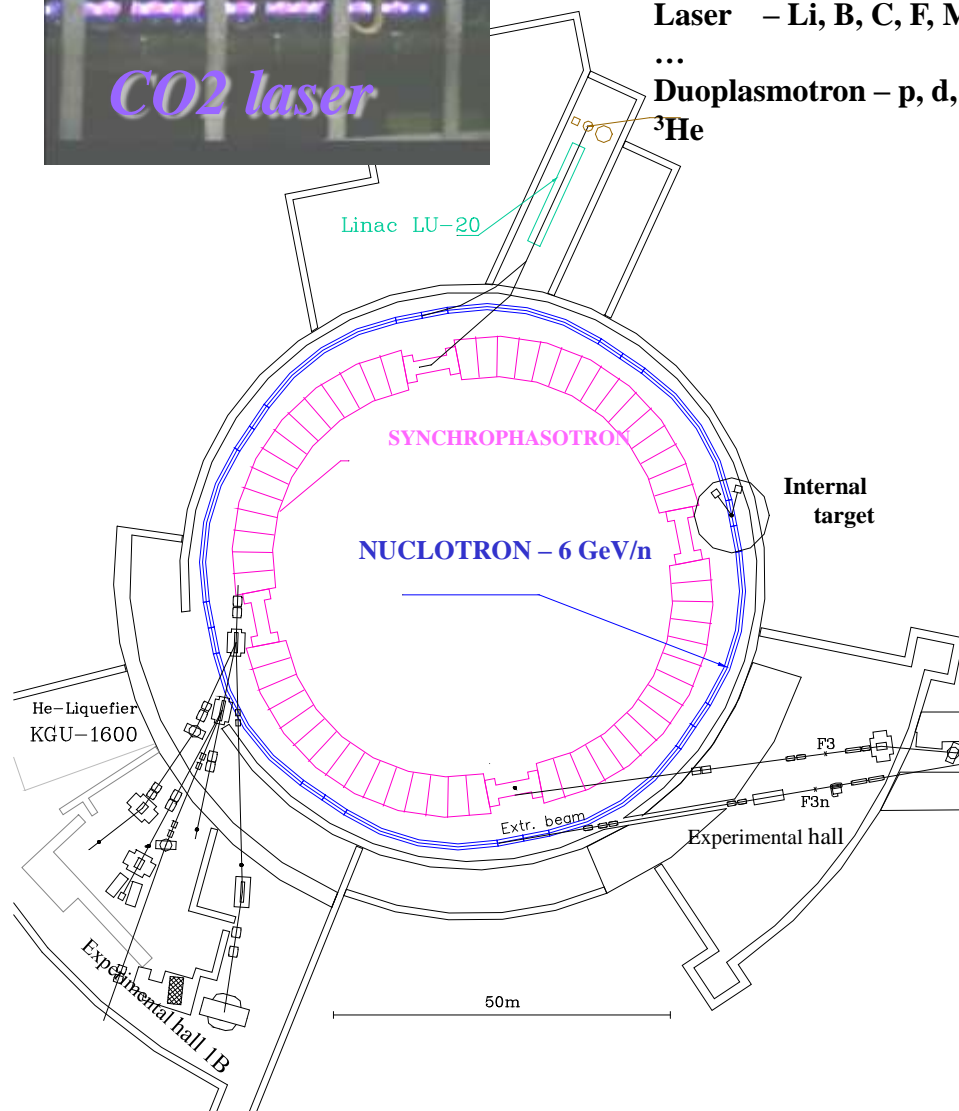


FIG. 3. The total differential Coulomb excitation cross sections for  $^{28}\text{Si}$  on  $^{208}\text{Pb}$  at  $E_{\text{lab}}/A=14.6$  GeV for the first-, second-, and third-order processes.

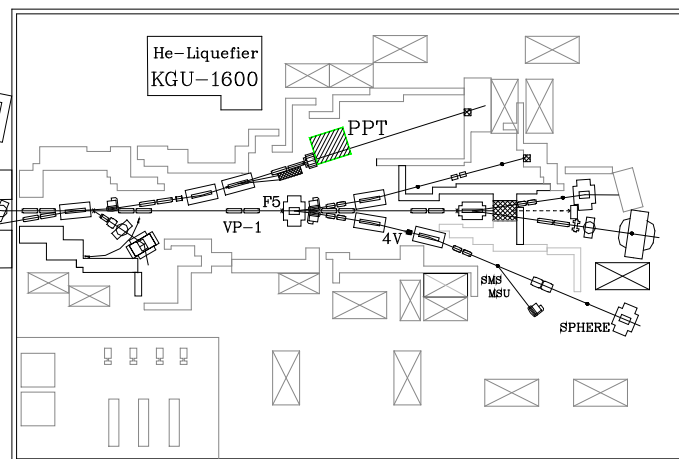
# Dubna: Relativistic Nuclei

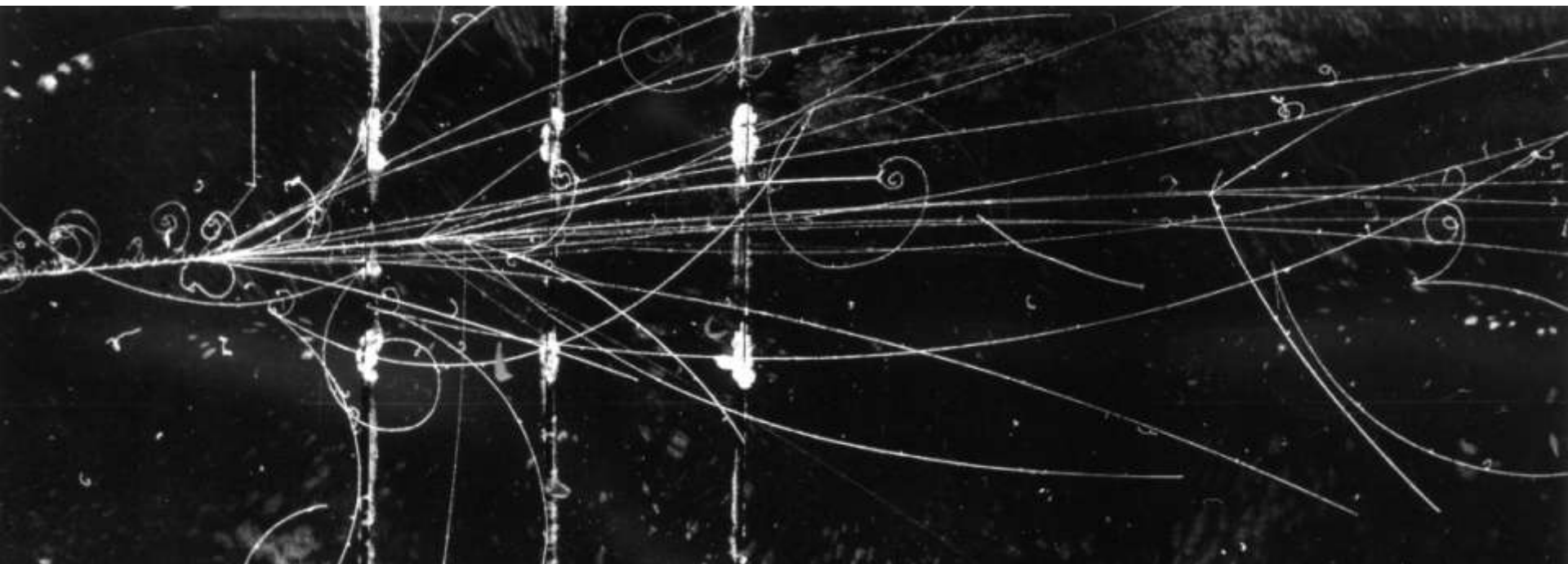


- Polaris –  $d^\uparrow$
- EBIS – N, Ar, Fe ...
- Laser – Li, B, C, F, Mg
- ...
- Duoplasmatron – p, d,  $\alpha$ ,  $^3\text{He}$



**Experimental hall 205**

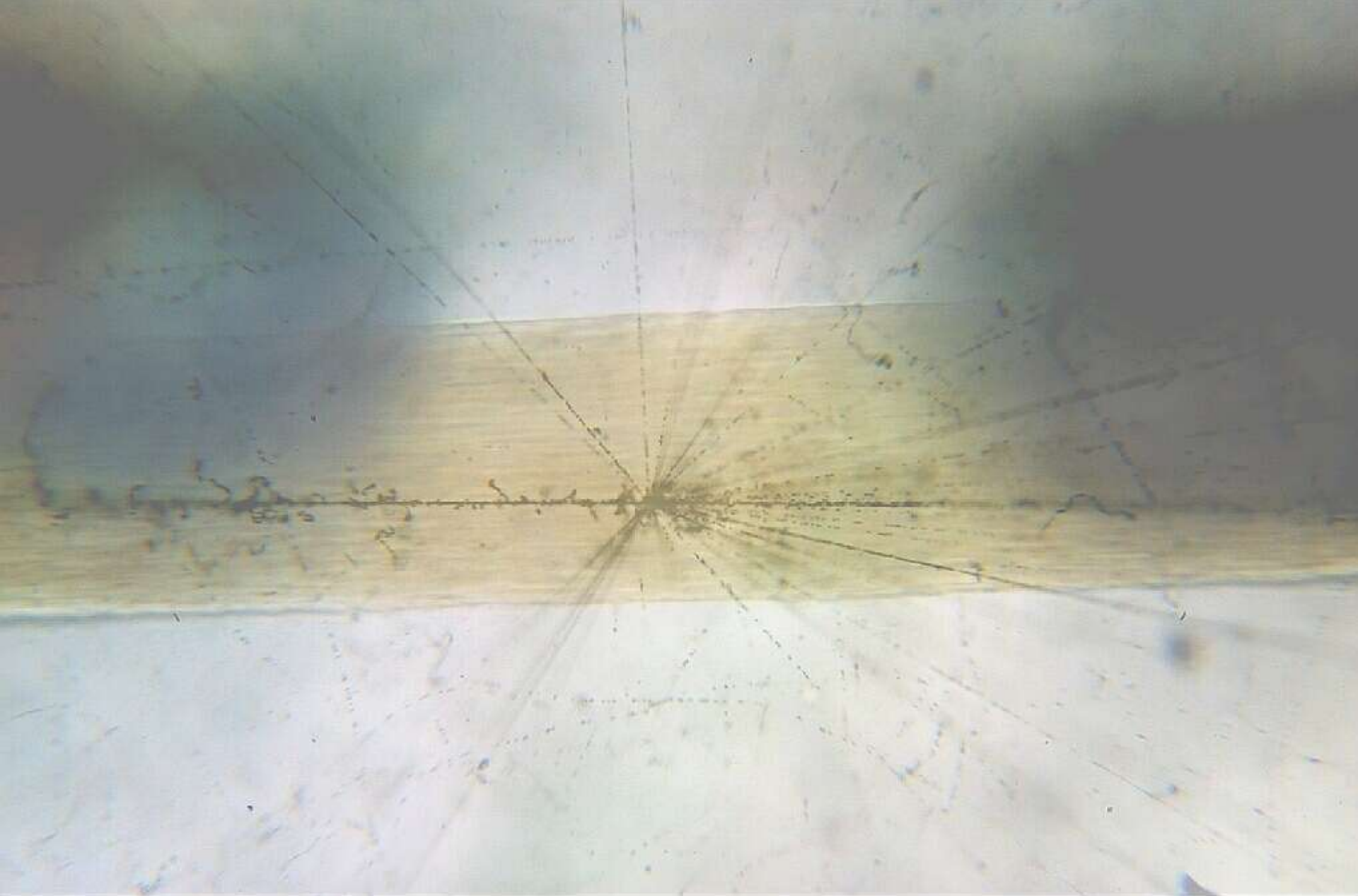




**Фотография взаимодействия ядра  $^{12}\text{C}$  с импульсом 4.5A ГэВ/с в пропановой пузырьковой камере (ЛВЭ ОИЯИ, 1974 г.).**



*0.5  $\mu\text{m}$  resolution, identification of charges and H&He isotopes*



**Photo of human hair superposed on nuclear star induced by relativistic sulphur nuclei in nuclear track emulsion**



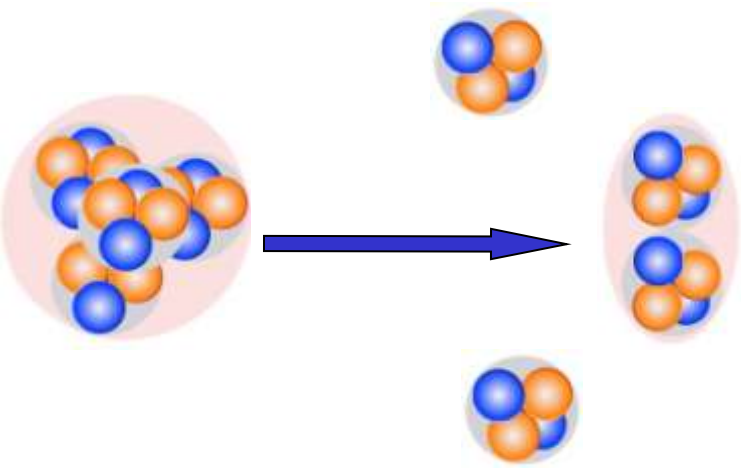
**ELEMENTARY PARTICLES AND FIELDS**  
**Experiment**

**Coherent Dissociation  $^{12}\text{C} \rightarrow 3\alpha$  in Lead-Enriched Emulsion  
at 4.5 GeV/c per Nucleon**

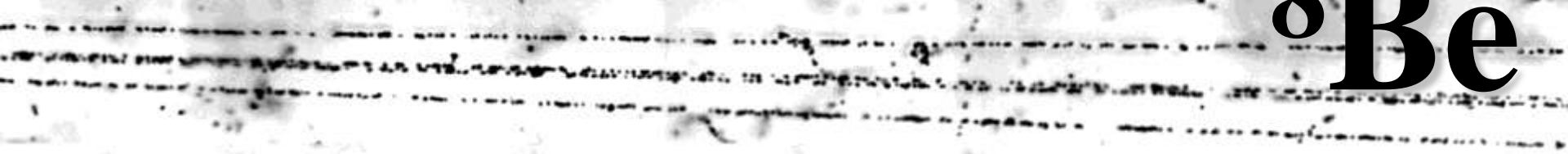
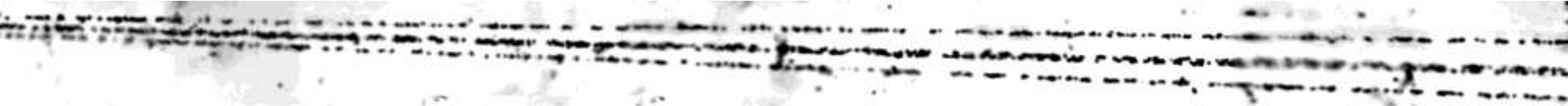
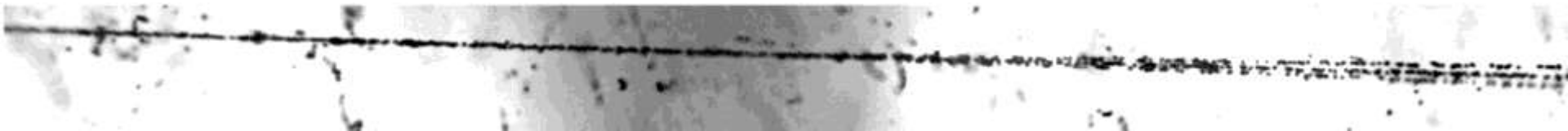
**V. V. Belaga, A. A. Benjaza<sup>1)</sup>, V. V. Rusakova, J. A. Salamov<sup>2)</sup>, and G. M. Chernov**

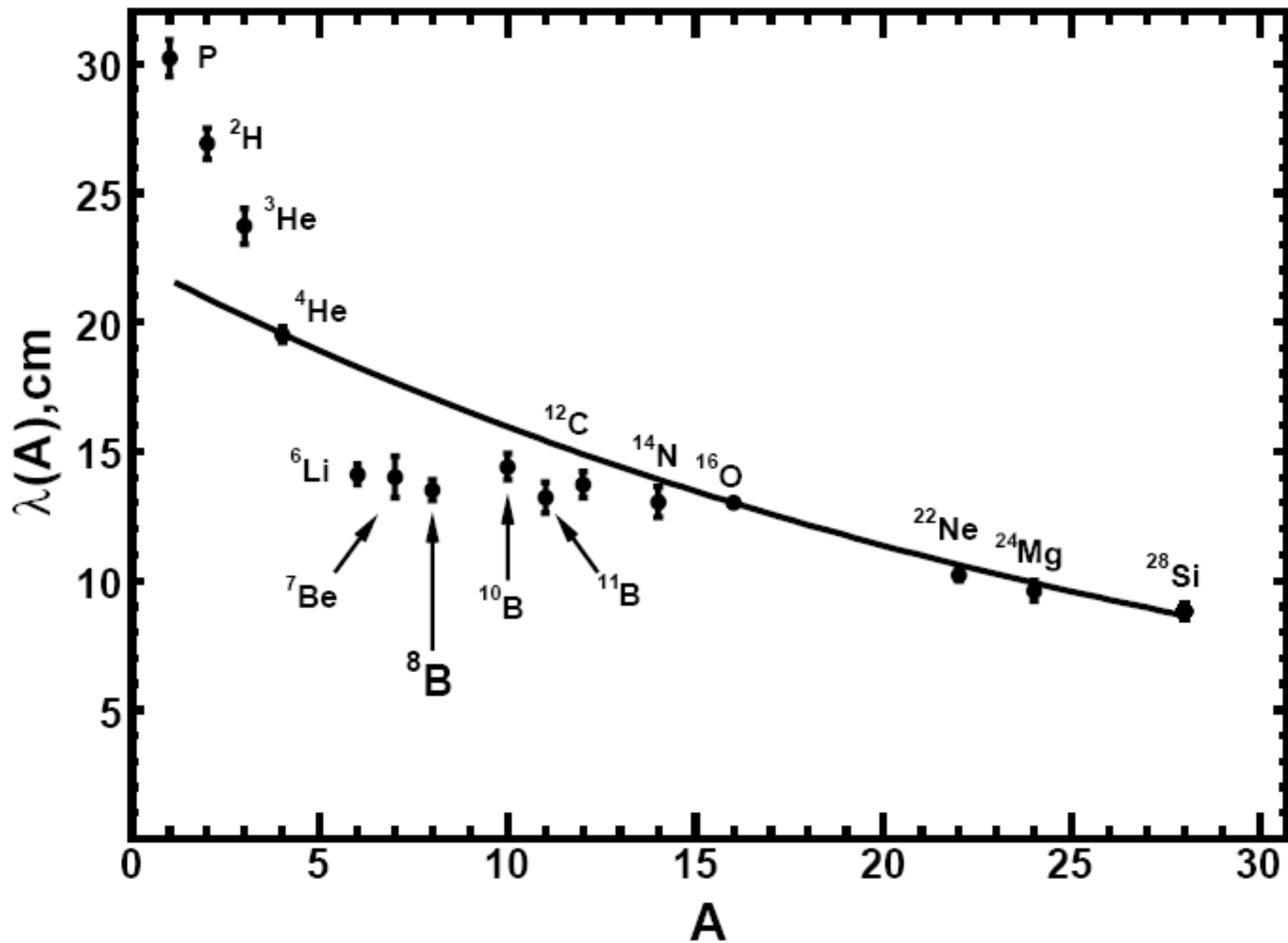
**$\approx 1\%$  of Inelastic Interactions**





4.5A GeV/c  $^{16}\text{O}$





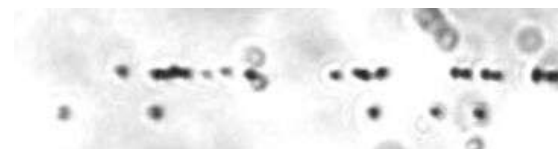
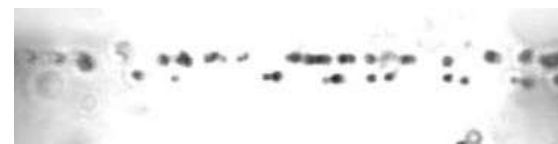
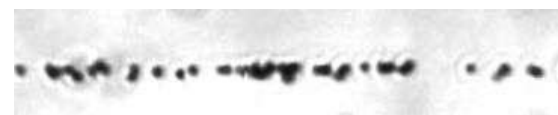
ELEMENTARY PARTICLES AND FIELDS  
Experiment

Interactions of Relativistic  ${}^6\text{Li}$  Nuclei  
with Photoemulsion Nuclei

M. I. Adamovich, V. G. Bogdanov<sup>1)</sup>, I. A. Konorov, V. G. Larionova,  
N. G. Peresadko, V. A. Plyushchev<sup>1)</sup>, Z. I. Solovyeva<sup>1)†</sup>, and S. P. Kharlamov

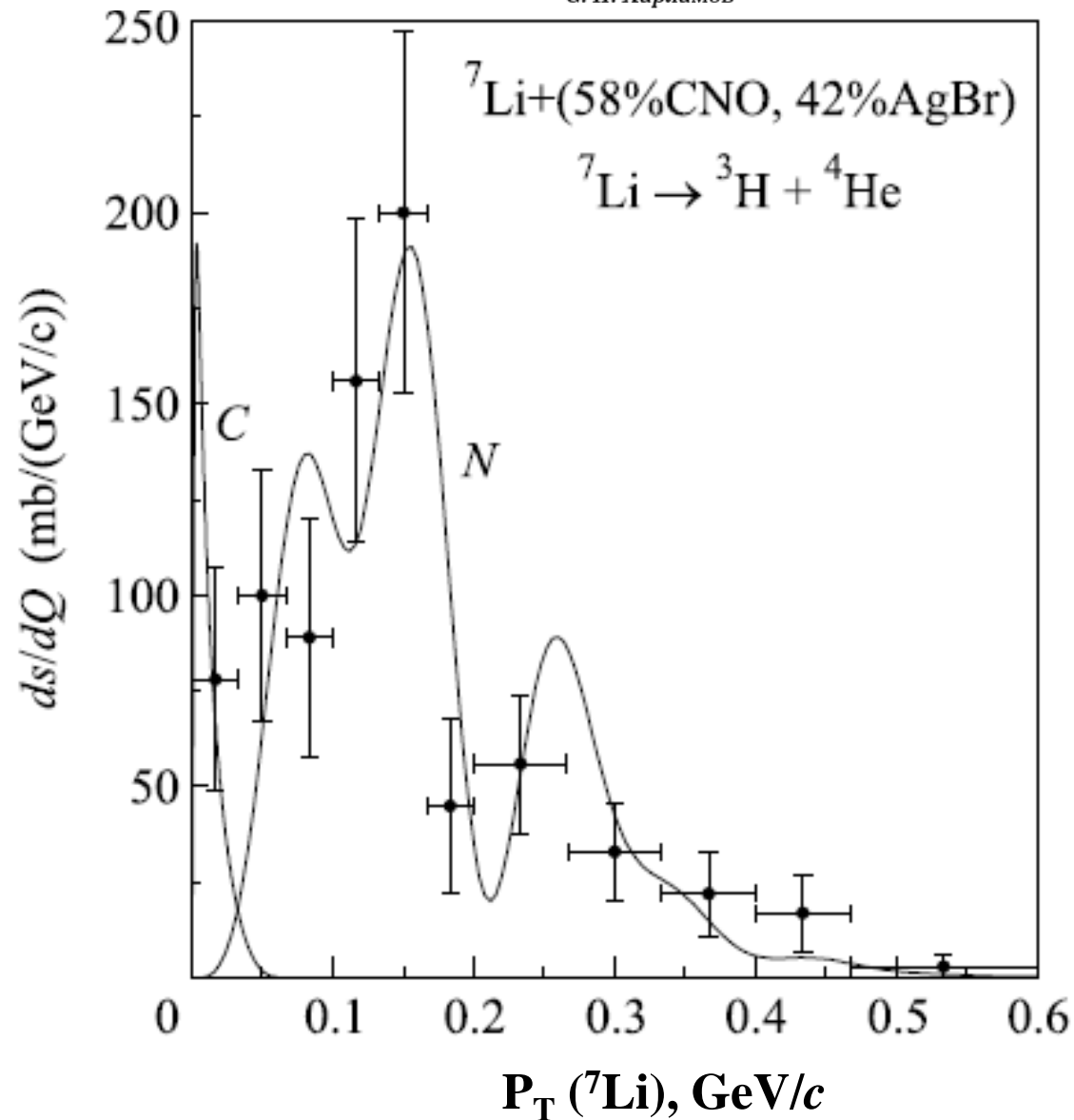
Number of events of  ${}^6\text{Li}$  coherent dissociation

Dissociation channel	Number of events	
	without the excitation of the target nucleus ( $N_h = 0$ )	with the excitation of the target nucleus ( $N_h \neq 0$ )
${}^4\text{He} + d$	23	24
${}^3\text{He} + t$	4	1
$t + d + p$	4	3
$d + d + d$	0	2

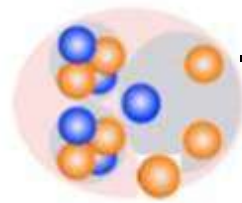


# Роль ядерного и электромагнитного взаимодействий в когерентной диссоциации релятивистского ядра ${}^7\text{Li}$ по каналу ${}^3\text{H} + {}^4\text{He}$

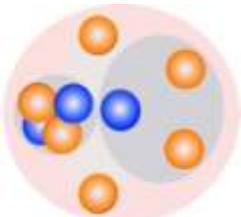
Н. Г. Пересадыко, В. Н. Фетисов<sup>1)</sup>, Ю. А. Александров, С. Г. Герасимов, В. А. Дронов, В. Г. Ларионова, Е. И. Тамм, С. П. Харламов



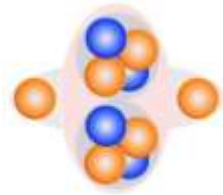
$^{12}\text{N}$  11 ms



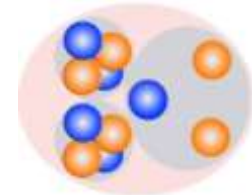
$^9\text{C}$  0.13 s



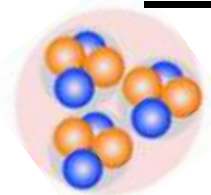
$^{10}\text{C}$  19 s



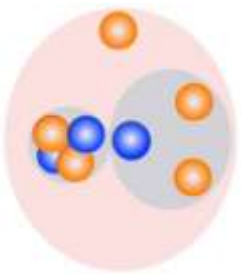
$^{11}\text{C}$  20 m



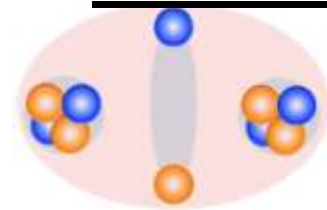
$^{12}\text{C}$  99%



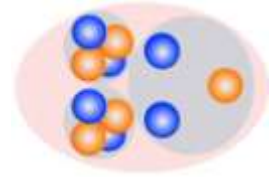
$^8\text{B}$  0.8 s



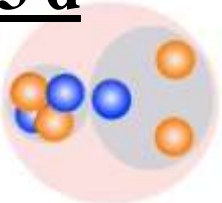
$^{10}\text{B}$  19.8%



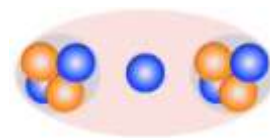
$^{11}\text{B}$  80.2%



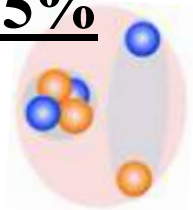
$^7\text{Be}$  53 d



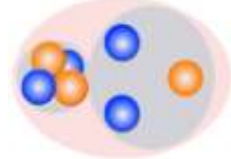
$^9\text{Be}$  100%

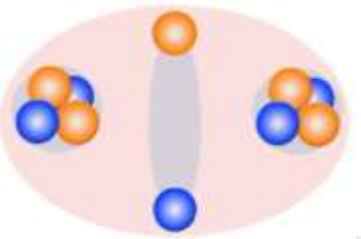


$^6\text{Li}$  7.5%

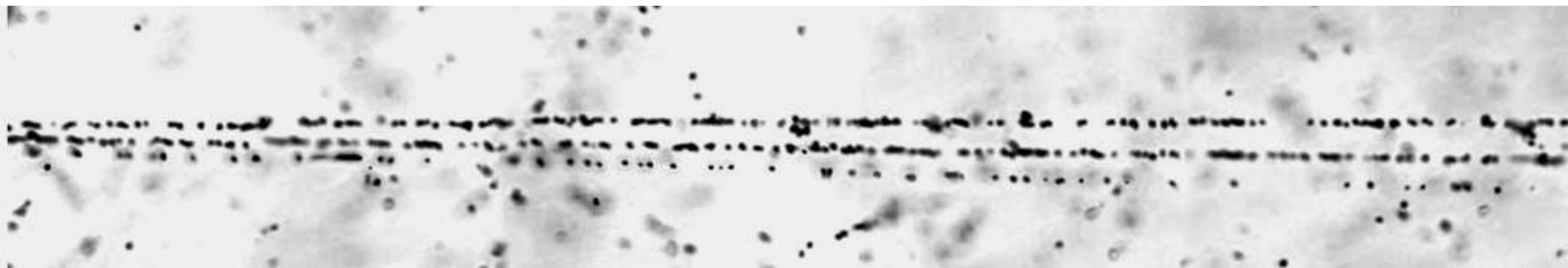


$^7\text{Li}$  92.5%





# 1A GeV <sup>10</sup>B

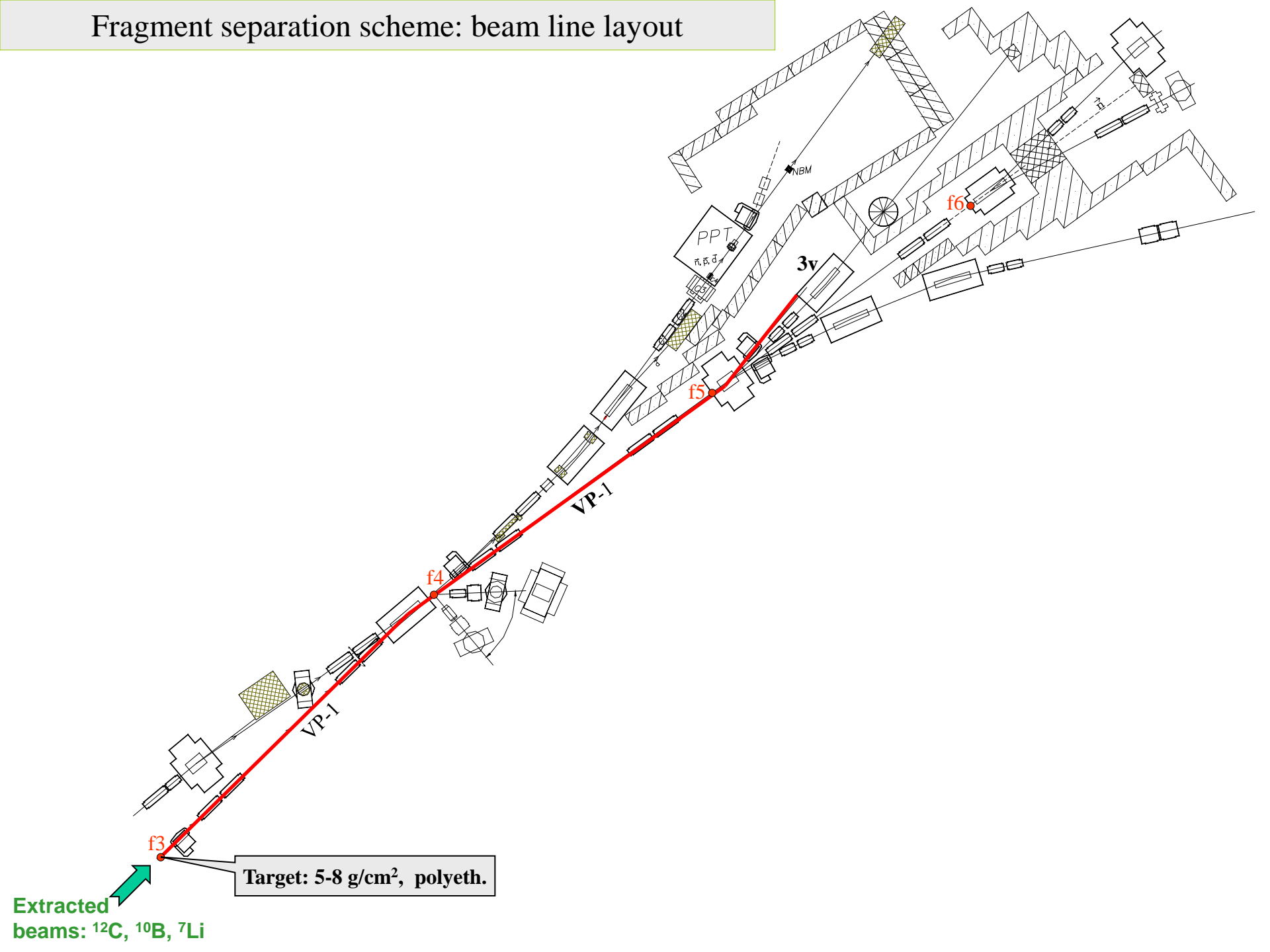


**2He+H (6 MeV) - 73%**

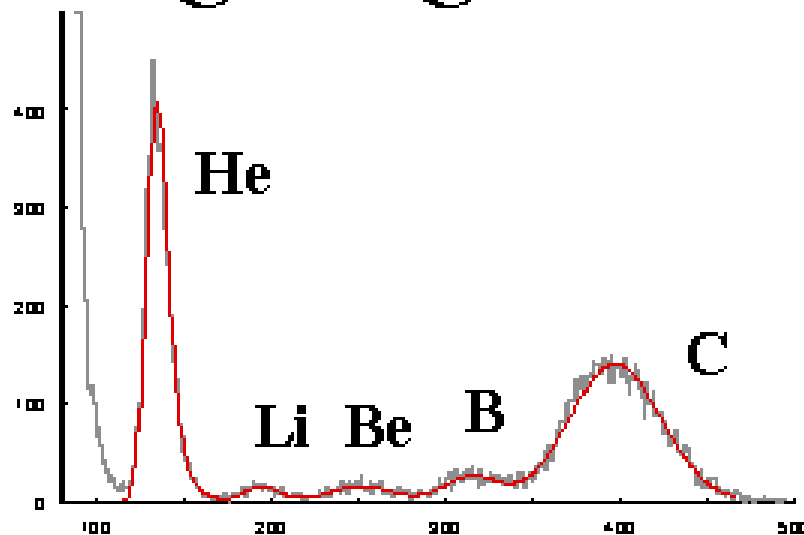
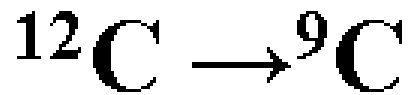
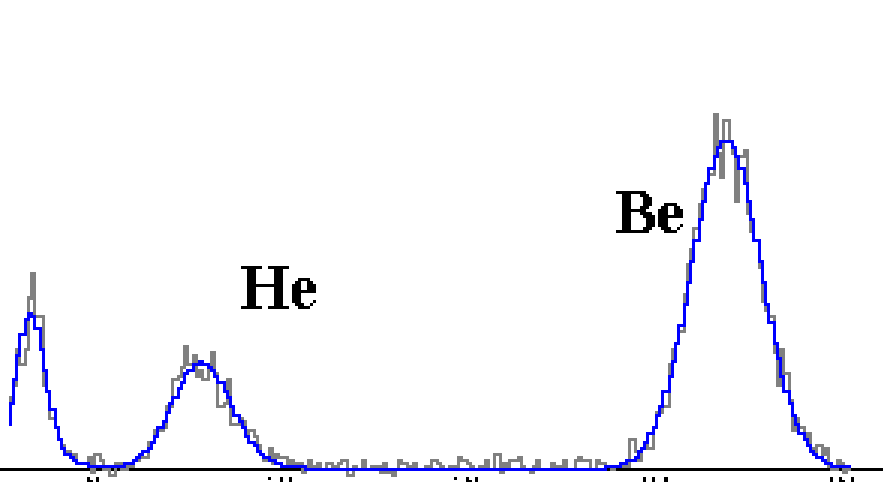
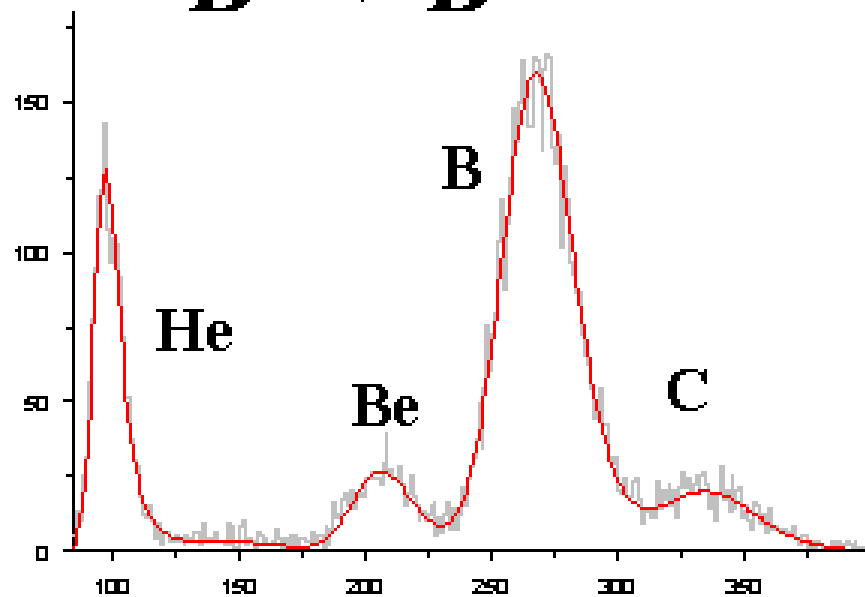
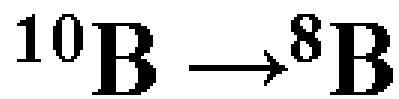
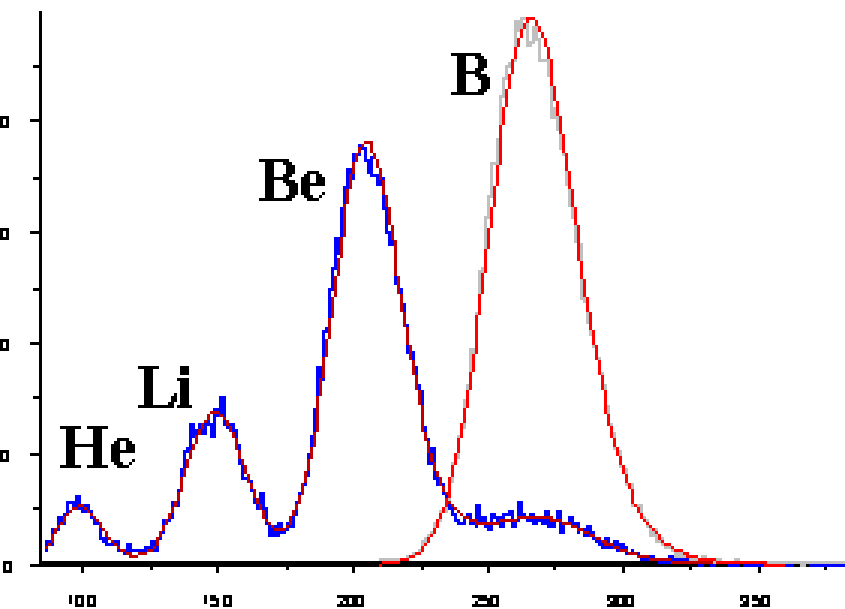
**Li+He (4.5 MeV) - 12%**

**<sup>9</sup>Be+p (6.6 MeV) - 2%**

# Fragment separation scheme: beam line layout







## ELEMENTARY PARTICLES AND FIELDS

### Experiment

# Electromagnetic Dissociation of Relativistic $^8\text{B}$ Nuclei in Nuclear Track Emulsion

R. Stanoeva<sup>1),2)</sup>, D. A. Artemenkov<sup>1)</sup>, V. Bradnova<sup>1)</sup>, S. Vokál<sup>1),3)</sup>,  
L. A. Goncharova<sup>4)</sup>, P. I. Zarubin<sup>1)\*</sup>, I. G. Zarubina<sup>1)</sup>, N. A. Kachalova<sup>1)</sup>,  
A. D. Kovalenko<sup>1)</sup>, D. O. Krivenkov<sup>1)</sup>, A. I. Malakhov<sup>1)</sup>, G. I. Orlova<sup>4)</sup>,  
N. G. Peresadko<sup>4)</sup>, N. G. Polukhina<sup>4)</sup>, P. A. Rukoyatkin<sup>1)</sup>, V. V. Rusakova<sup>1)</sup>,  
M. Haiduc<sup>5)</sup>, S. P. Kharlamov<sup>4)</sup>, M. M. Chernyavsky<sup>4)</sup>, and T. V. Shchedrina<sup>1)</sup>

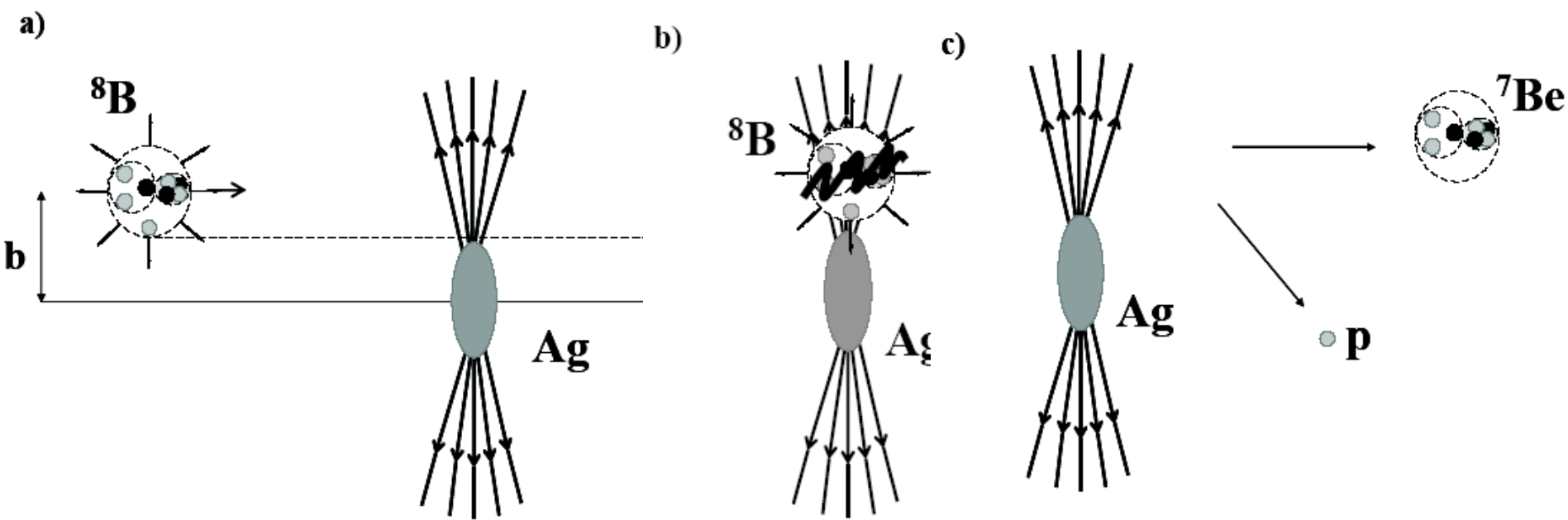
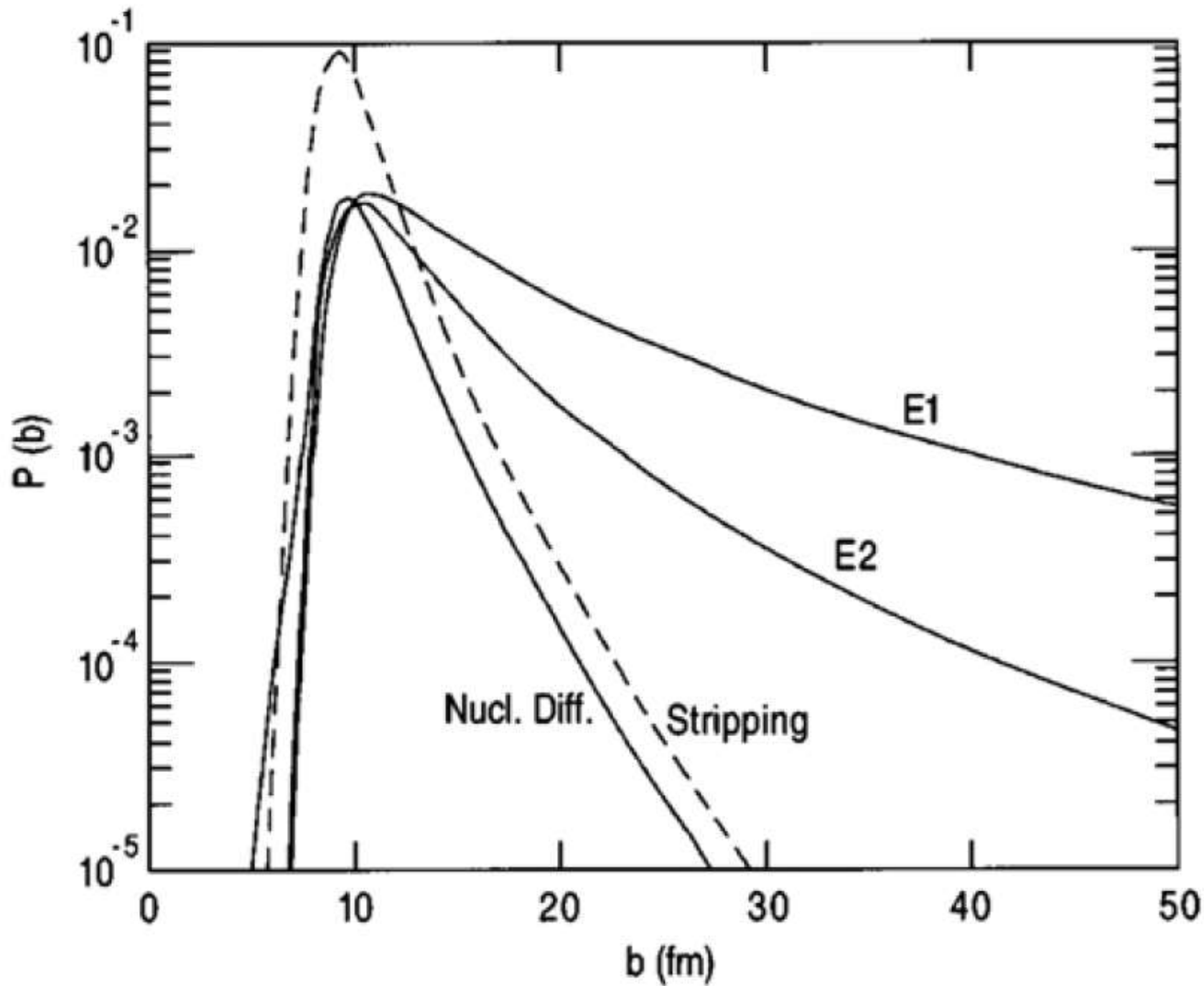


Fig. 1. Scheme of the electromagnetic dissociation of a relativistic  $^8\text{B}$  nucleus in the field of a silver nucleus: (a) approach of the nuclei at an impact parameter  $b$ ; (b) quasireal-photon absorption by the  $^8\text{B}$  nucleus; and (c) dissociation into two fragments,  $p$  and  $^7\text{Be}$ .



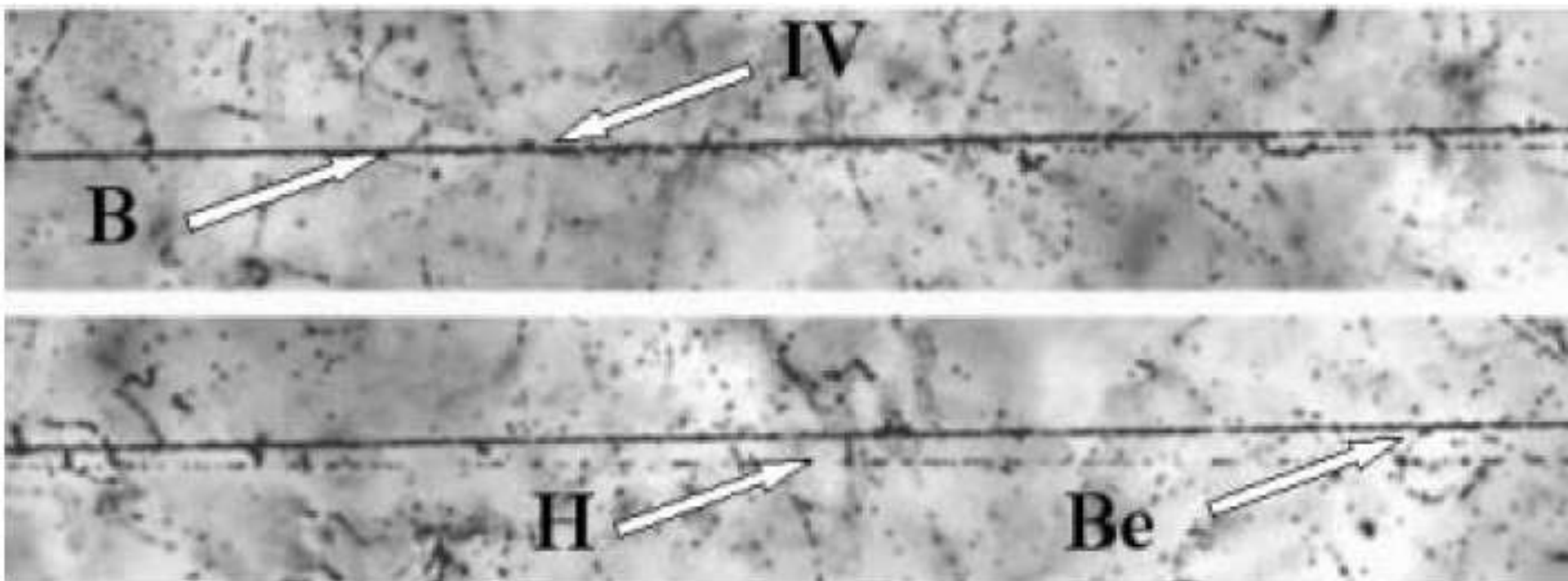
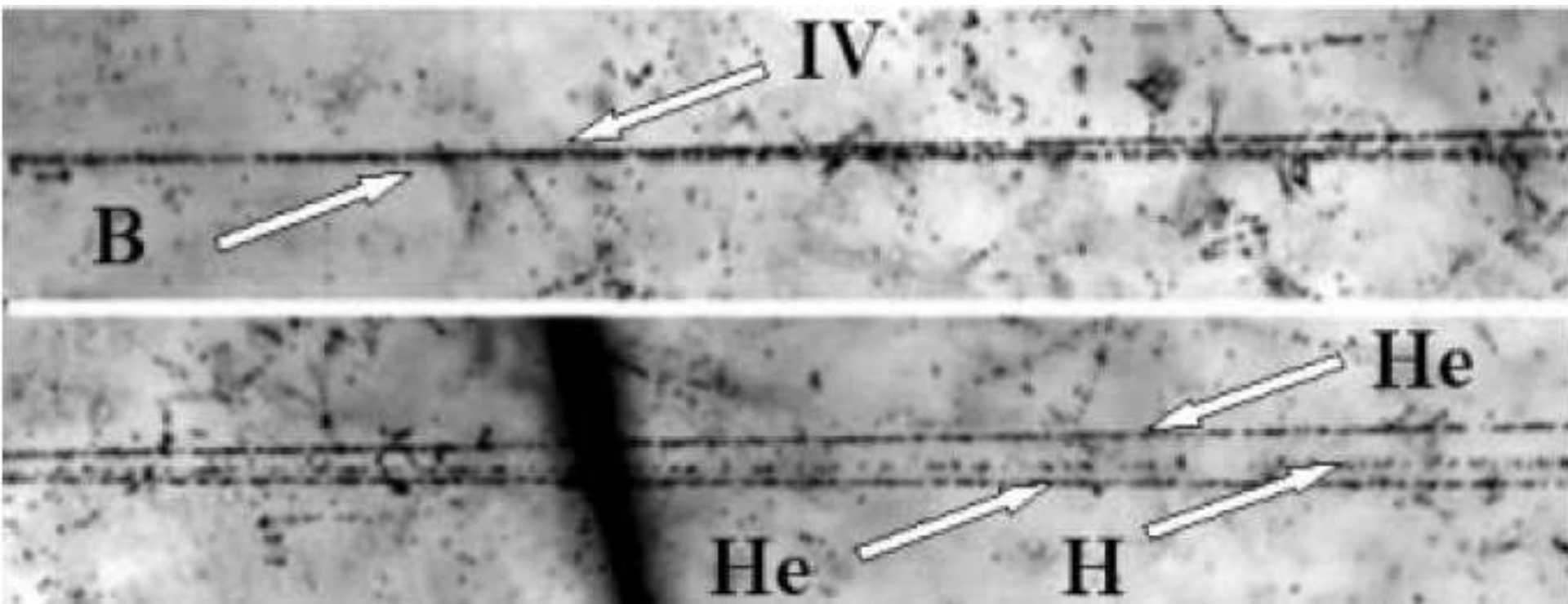
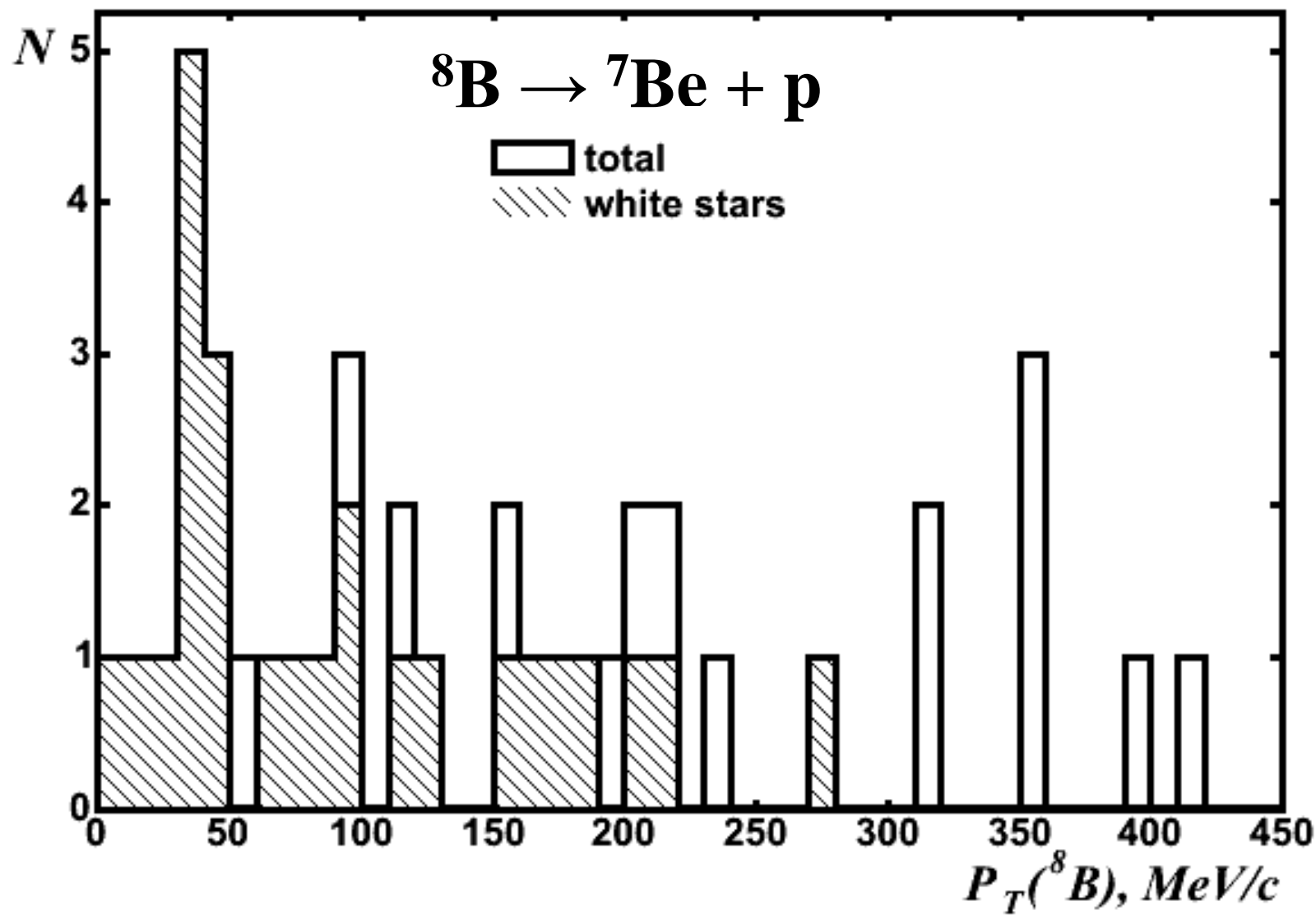


FIG. 1: Example of peripheral interaction of a 1.2 A GeV  ${}^8\text{B} \rightarrow {}^7\text{Be} + \text{p}$  in a nuclear track emulsion (“white” star). The interaction vertex (indicated as **IV**) and nuclear fragment tracks (**H** and **Be**) in a narrow angular cone are seen on the upper and bottom microphotograph.



	$Q_{\min}$ ( $^{10}\text{B}$ ), MeV	$N_{ws}$ ( $^{10}\text{B}$ )	% ( $^{10}\text{B}$ )	$Q_{\min}$ ( $^8\text{B}$ ), MeV	$N_{ws}$ ( $^8\text{B}$ )	% ( $^8\text{B}$ )
2He+H	6.0	30	73	1.724	14	27
He+3H	25	5	12	8.6	12	23
Be+H	6.6	1	2	0.138	25	48
B		-	-		1	2
Li+He	4.5	5	13	3.7	-	-



## Systematic study of $^8\text{B}$ breakup cross sections

Henning Esbensen

*Physics Division, Argonne National Laboratory, Argonne, Illinois 60439*

Kai Hencken

*Institut für Physik, Universität Basel, 4056 Basel, Switzerland*

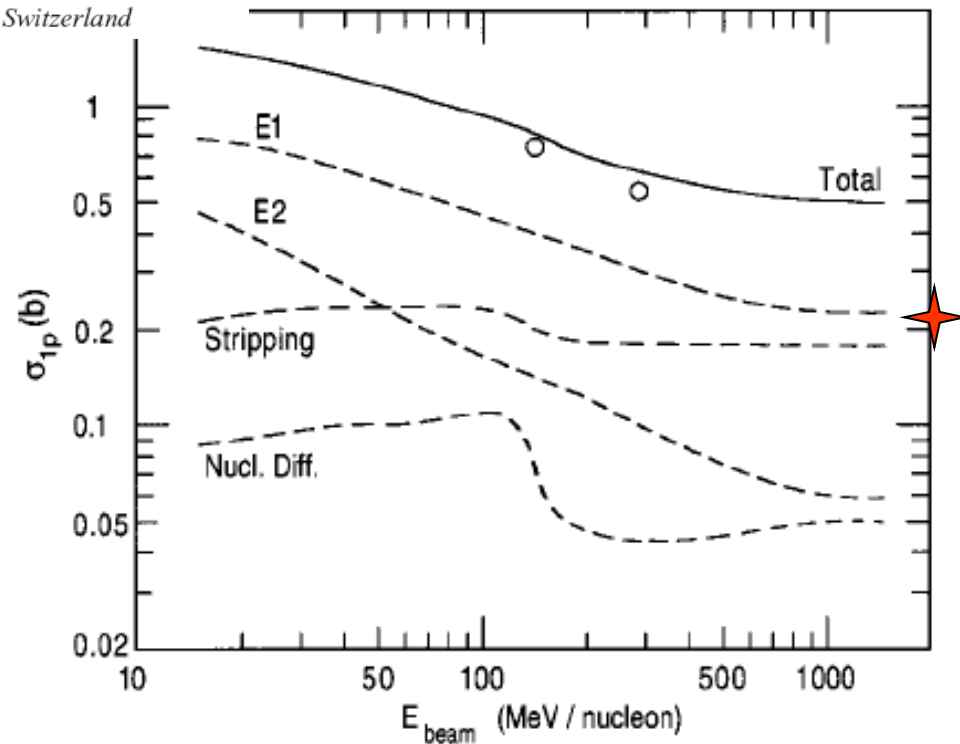


FIG. 8. One-proton removal cross section (solid curve) for  $^8\text{B}$  on a Pb target, as a function of the beam energy. The dashed curves show the contributions from stripping, nuclear induced diffraction, and the  $E1$  and  $E2$  Coulomb dissociation. The two data points are from Ref. [8].



Detailed study of relativistic  ${}^9\text{Be} \rightarrow 2\alpha$   
fragmentation in peripheral collisions  
in a nuclear track emulsion\*

D. A. Artemenkov\*\*, D. O. Krivenkov, T. V. Shchedrina, R. Stanoeva,  
P. I. Zarubin

“white” stars



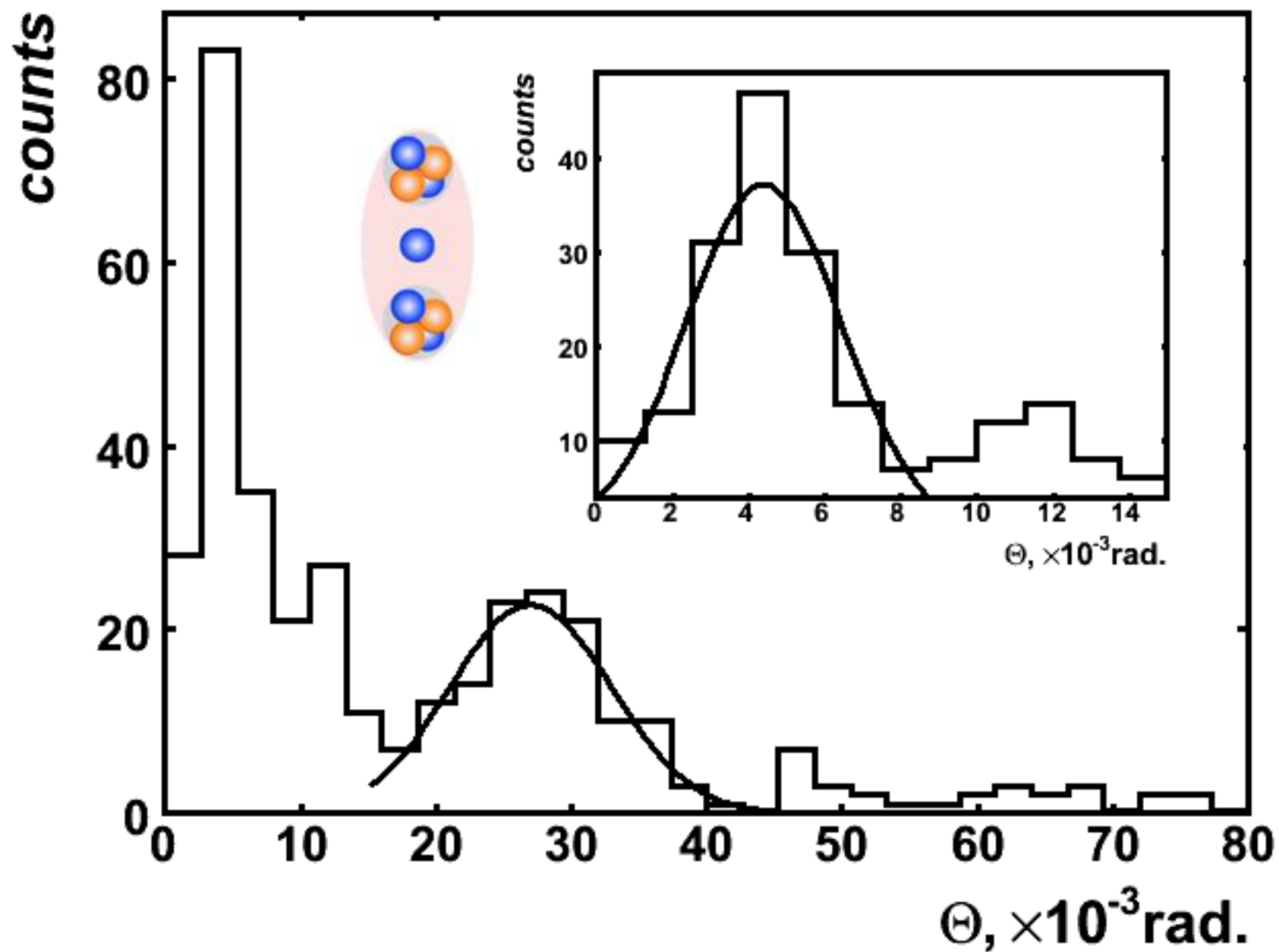
+1.7 MeV

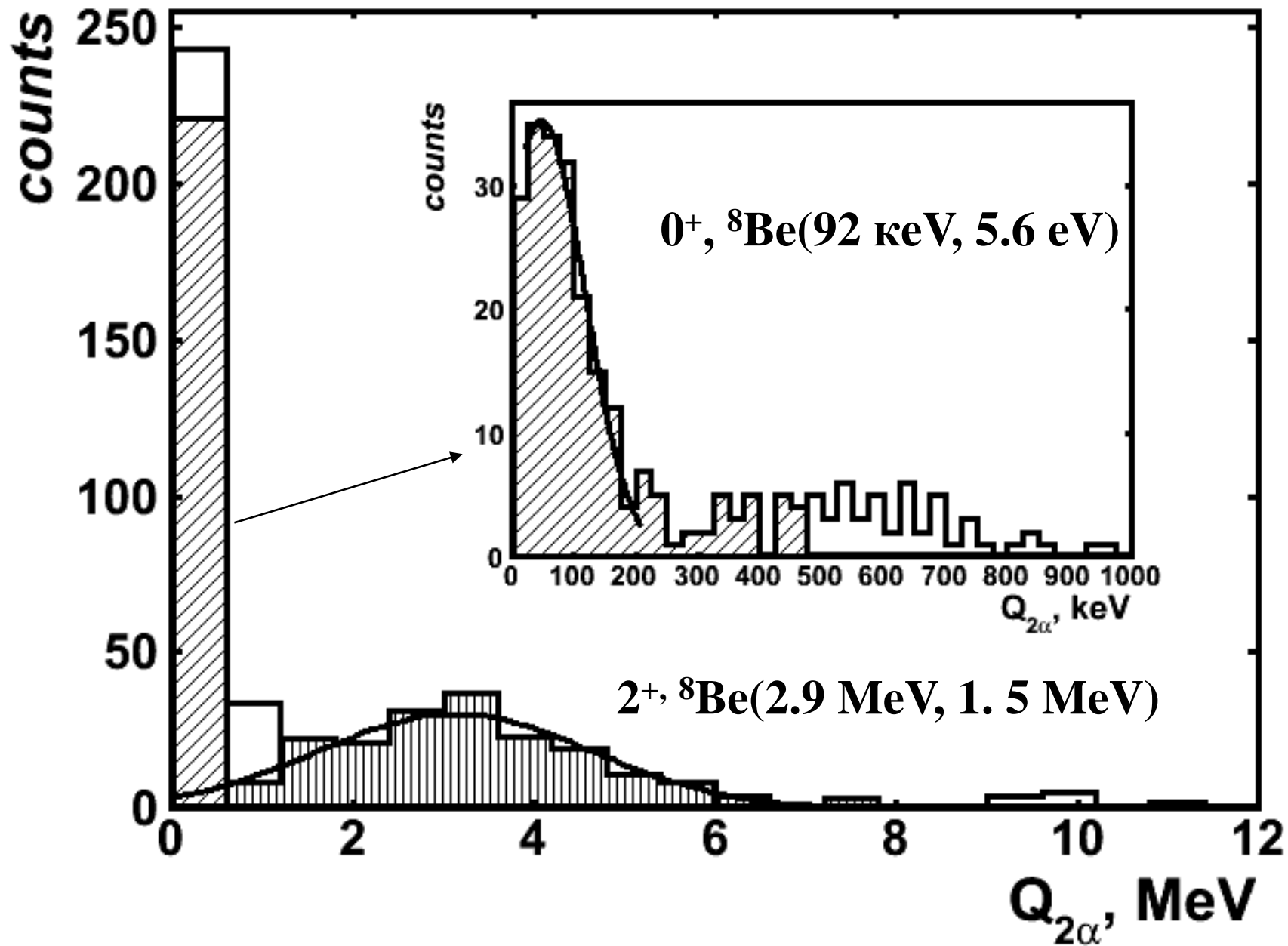
stars with target proton like recoil (g-particle)



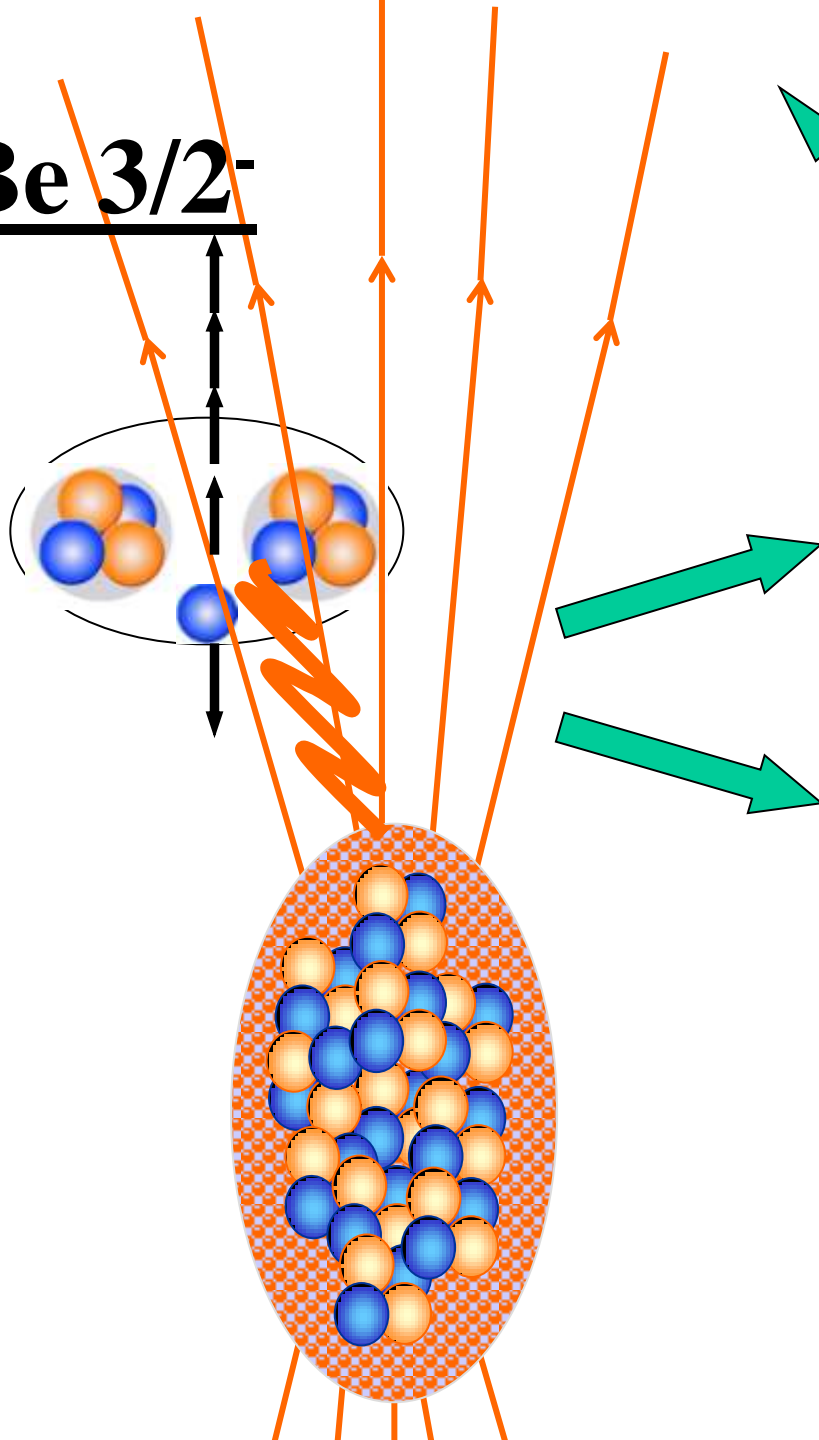
stars with heavy fragment of target nucleus (b-particle)



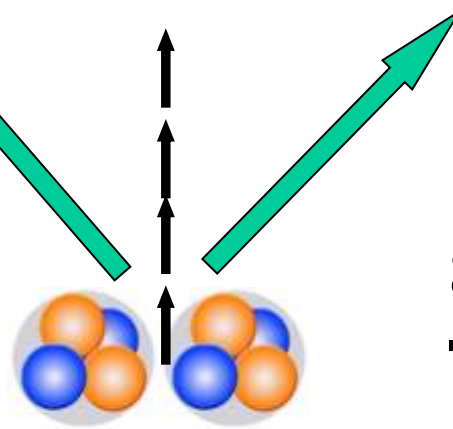




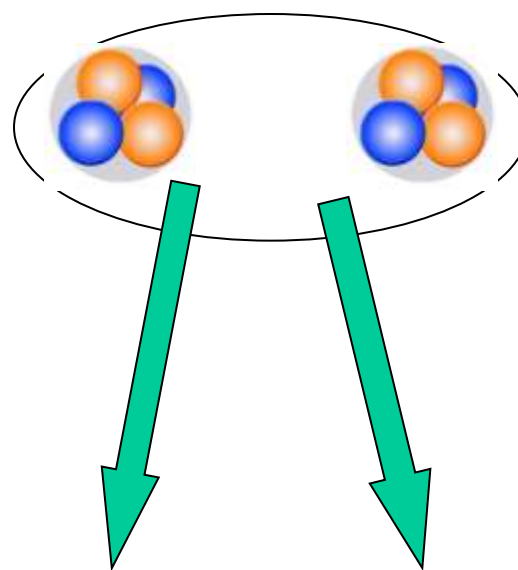
${}^9\text{Be } 3/2^-$



${}^8\text{Be } 2^+$



${}^8\text{Be } 0^+$



$$|{}^9\text{Be}(3/2^-)\rangle = \varpi_{0^+} | [{}^8\text{Be}(0^+) \otimes n_{p(3/2)} ]_{3/2^-} \rangle + \varpi_{2^+} | [{}^8\text{Be}(2^+) \otimes n_{p(3/2)} ]_{3/2^-} \rangle$$

где  $\omega_{0^+} = 0.535$  и  $\omega_{2^+} = 0.465$  – весовые коэффициенты.

1. Y. L. Parfenova and Ch. Leclercq-Willain, «Hyperfine anomaly in Be isotopes and neutron spatial distribution: A three-cluster model for  ${}^9\text{Be}$ », Phys. Rev. C 72, 054304 (2005).

2. Y. L. Parfenova and Ch. Leclercq-Willain, «Hyperfine anomaly in Be isotopes in the cluster model and the neutron spatial distribution», Phys. Rev. C 72, 024312(2005).

Доли событий  $\Theta_{0^+}$  и  $\Theta_{2^+}$  составляют  $0.56 \pm 0.04$  и  $0.44 \pm 0.04$ . Эти величины демонстрируют соответствие весам  $0^+$  и  $2^+$  состояний, принятым в теоретических работах [1,2]. Они указывают присутствие этих состояний как компонент основного состояния ядра  ${}^9\text{Be}$ .

# Резюме

Впервые выполнено исследование структуры  $2\alpha + n$  ядра  ${}^9\text{Be}$  методом релятивистской фрагментации. В основном, фрагментация  ${}^9\text{Be} \rightarrow 2\alpha$  (81 %) протекает через состояния  $0^+$  и  $2^+$  ядра  ${}^8\text{Be}$  с близкими вероятностями.

Для когерентной диссоциации  ${}^9\text{Be} \rightarrow 2\alpha$  на тяжелых ядрах наблюдается уменьшение среднего значения поперечного импульса  $P_{Tsum}$   $\alpha$ -пары по сравнению с взаимодействиями на протонах мишени.

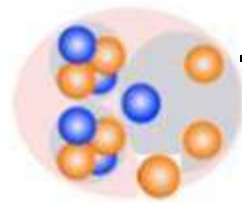
Для основной части событий когерентной диссоциации распределение величин суммарного поперечного импульса  $P_{Tsum}$  пар  $\alpha$ -частиц, может быть объяснено в рамках статистической модели недостающим поперечным импульсом нейтрона – фрагмента ядра  ${}^9\text{Be}$ .

Отсутствует значимое различие распределений по  $P_{Tsum}$  для событий когерентной диссоциации через состояния  $0^+$  и  $2^+$  ядра  ${}^8\text{Be}$ . Это обстоятельство указывает на схожий механизм возникновения  $\alpha$ -пар и их одновременное присутствие этих состояний с близкими весами в основном состоянии ядра  ${}^9\text{Be}$ .

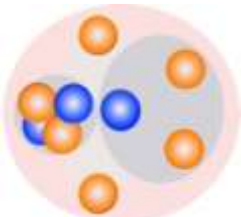
В целом, представленные данные можно рассматривать как доказательство того, что в структуре ядра  ${}^9\text{Be}$  с высокой вероятностью имеется кор в виде двух состояний ядра  ${}^8\text{Be}$  и внешнего нейтрона.

Полученные результаты согласуются с теоретическими работами по описанию структуры ядра  ${}^9\text{Be}$ , предполагающими присутствие в его основном состоянии состояния  $0^+$  и  $2^+$  ядра  ${}^8\text{Be}$  приблизительно с одинаковыми весами.

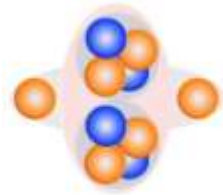
$^{12}\text{N}$  11 ms



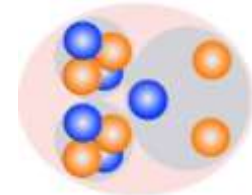
$^9\text{C}$  0.13 s



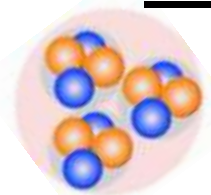
$^{10}\text{C}$  19 s



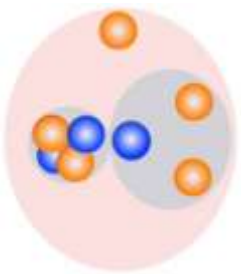
$^{11}\text{C}$  20 m



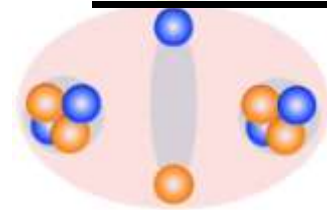
$^{12}\text{C}$  99%



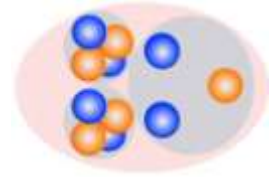
$^8\text{B}$  0.8 s



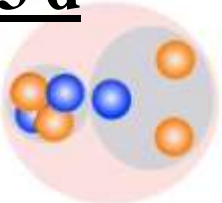
$^{10}\text{B}$  19.8%



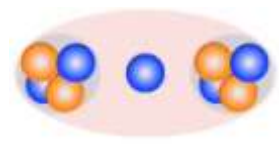
$^{11}\text{B}$  80.2%



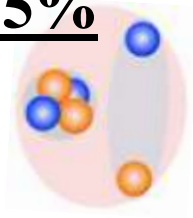
$^7\text{Be}$  53 d



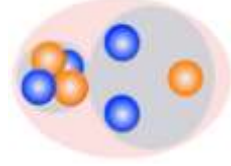
$^9\text{Be}$  100%

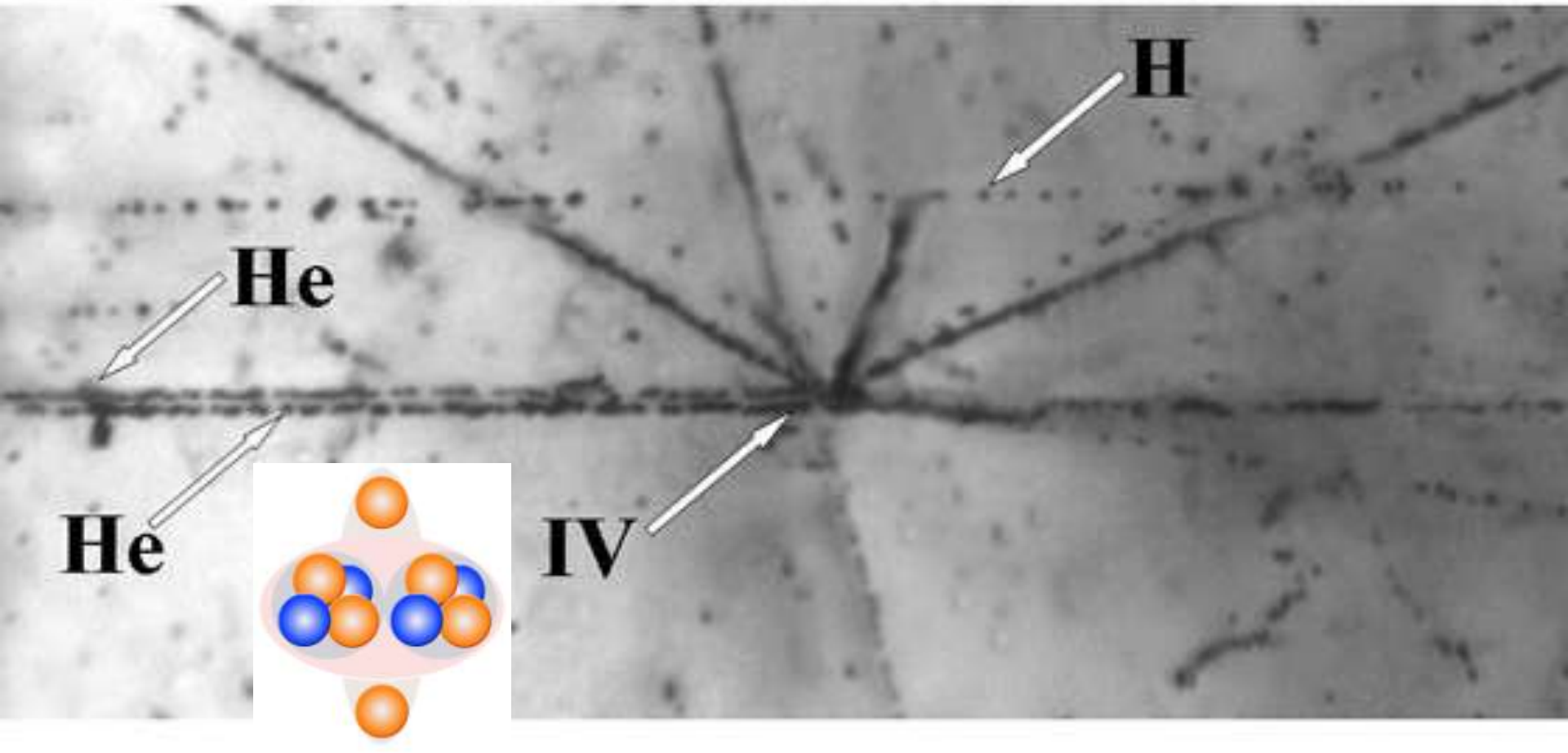
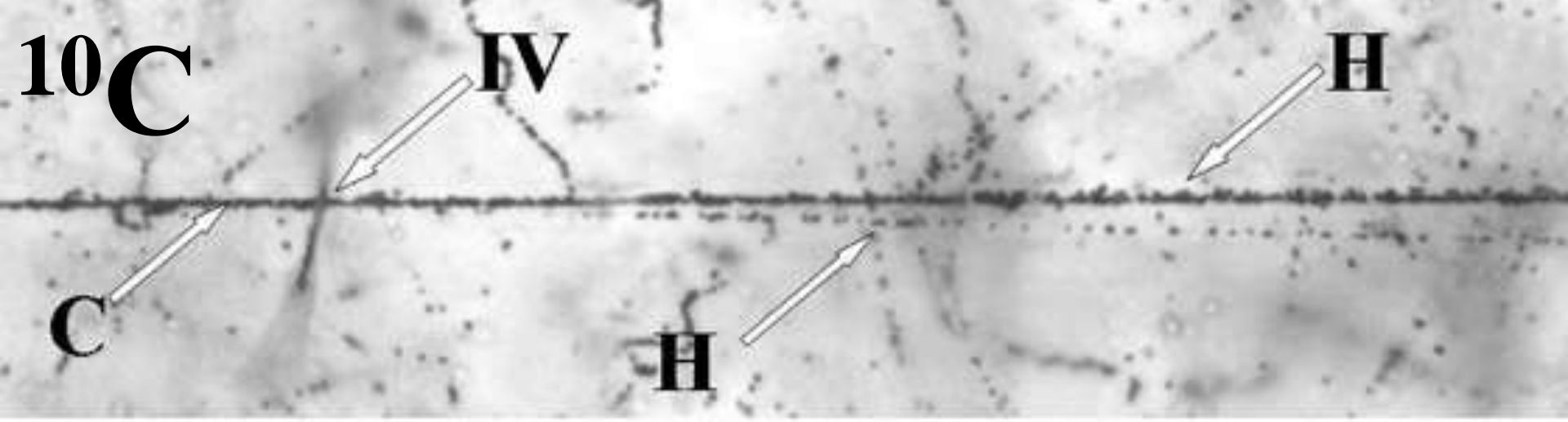


$^6\text{Li}$  7.5%

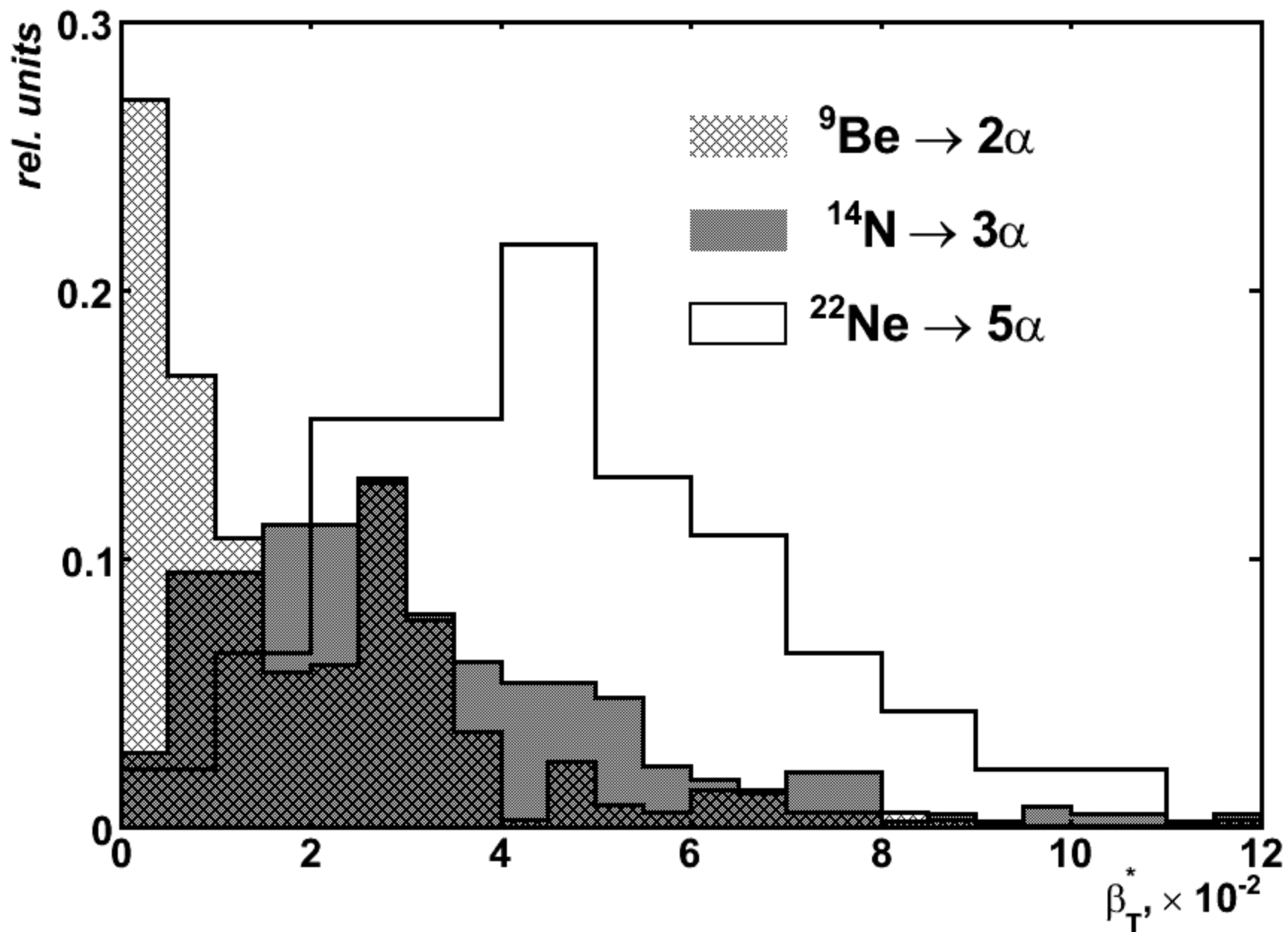


$^7\text{Li}$  92.5%





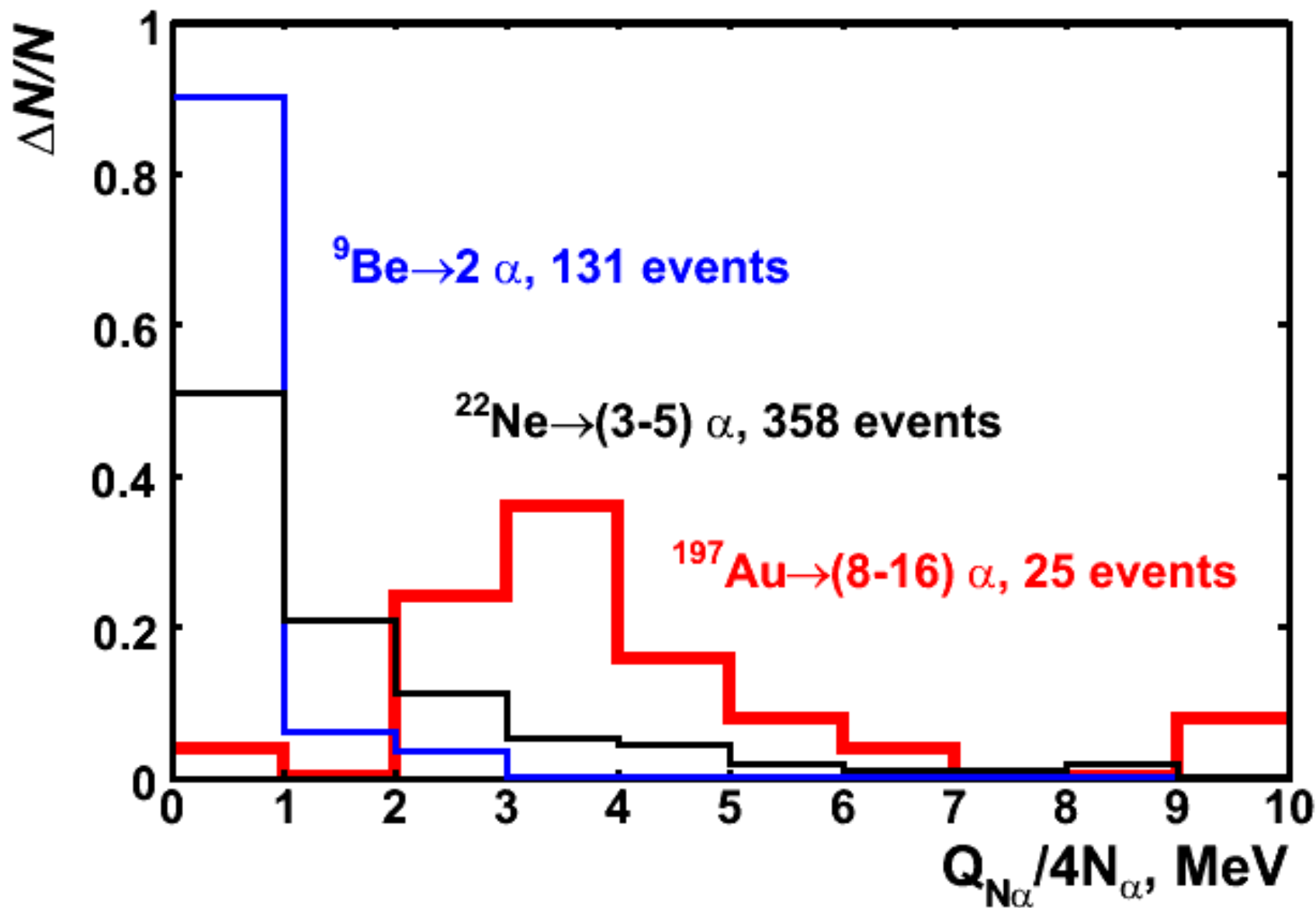


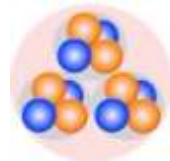
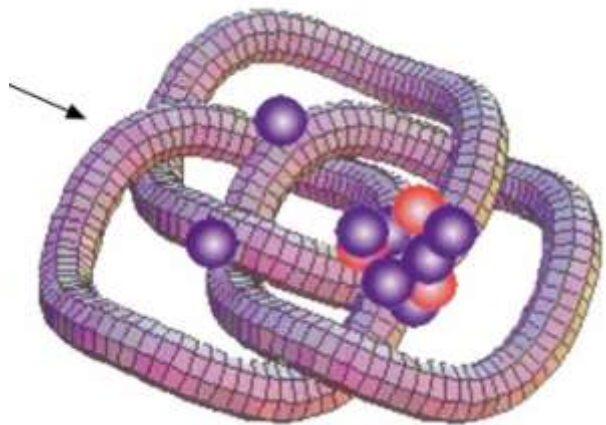


1.2A GeV  $^9\text{Be}$

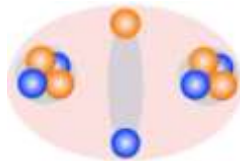
3.22A GeV  $^{22}\text{Ne}$

10.7A GeV  $^{197}\text{Au}$

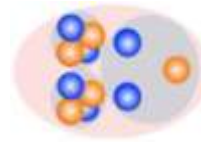




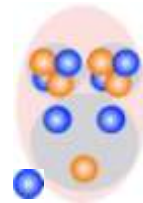
$^{12}\text{B}$  20 ms



$^{10}\text{Be}$  1510000 y



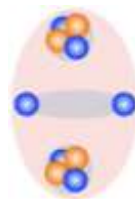
$^{12}\text{Be}$  23 ms



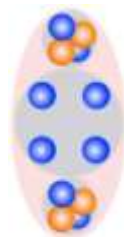
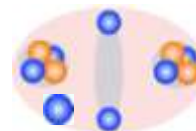
$^{11}\text{Be}$  13.8 s



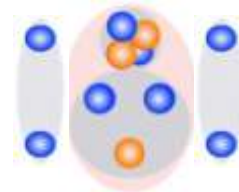
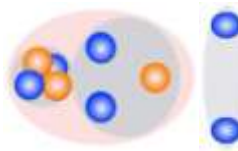
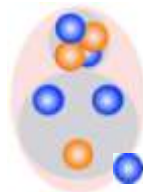
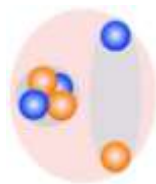
$^8\text{Li}$  838 ms



$^9\text{Li}$  178 ms



$^{11}\text{Li}$  8.5 ms



$0^+$

$0^+$

$^6\text{He}$  807 ms

$^8\text{He}$  119 ms



## Beta Decay of a $C^9$ Nucleus\*

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(Received June 29, 1956)

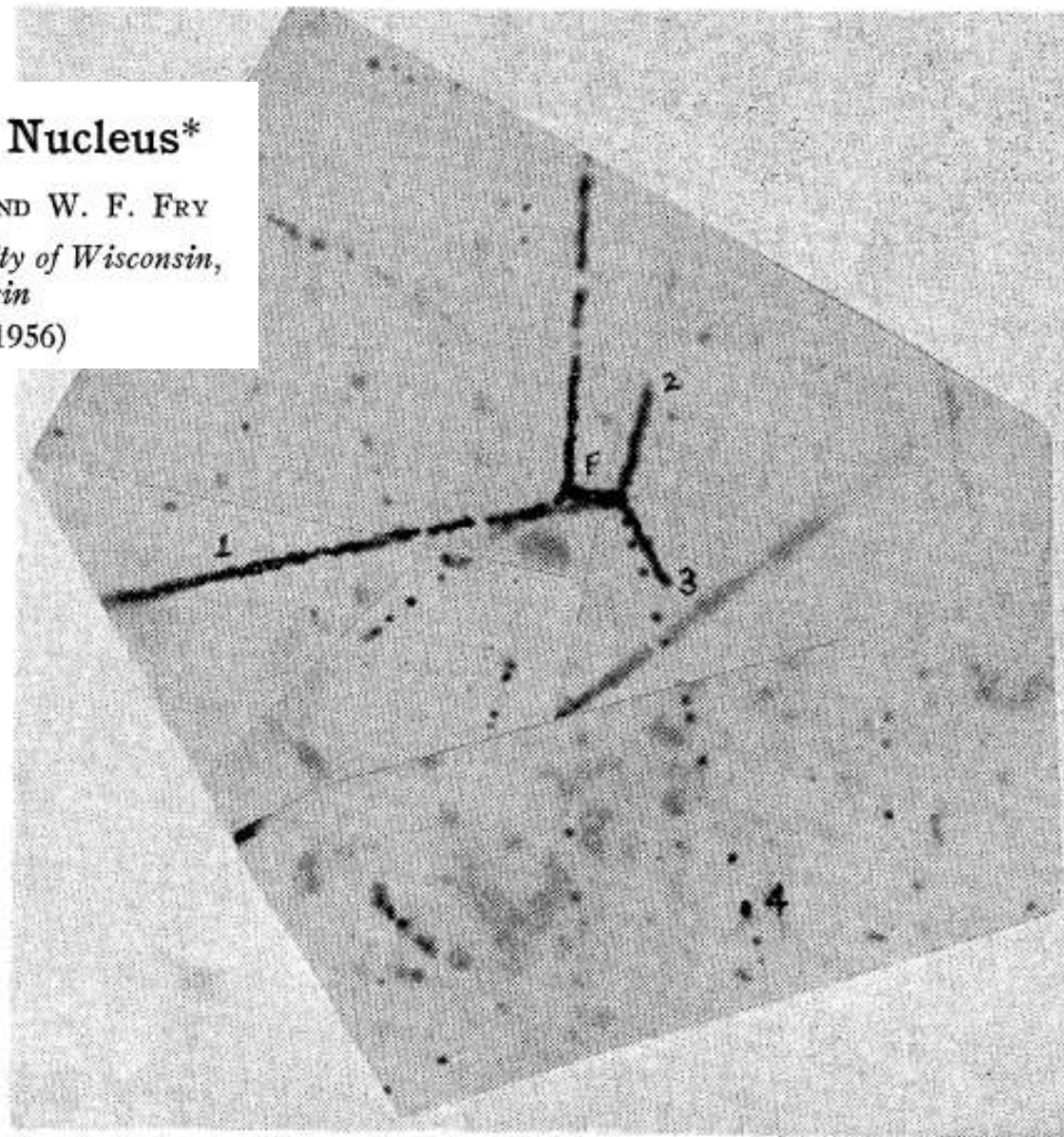
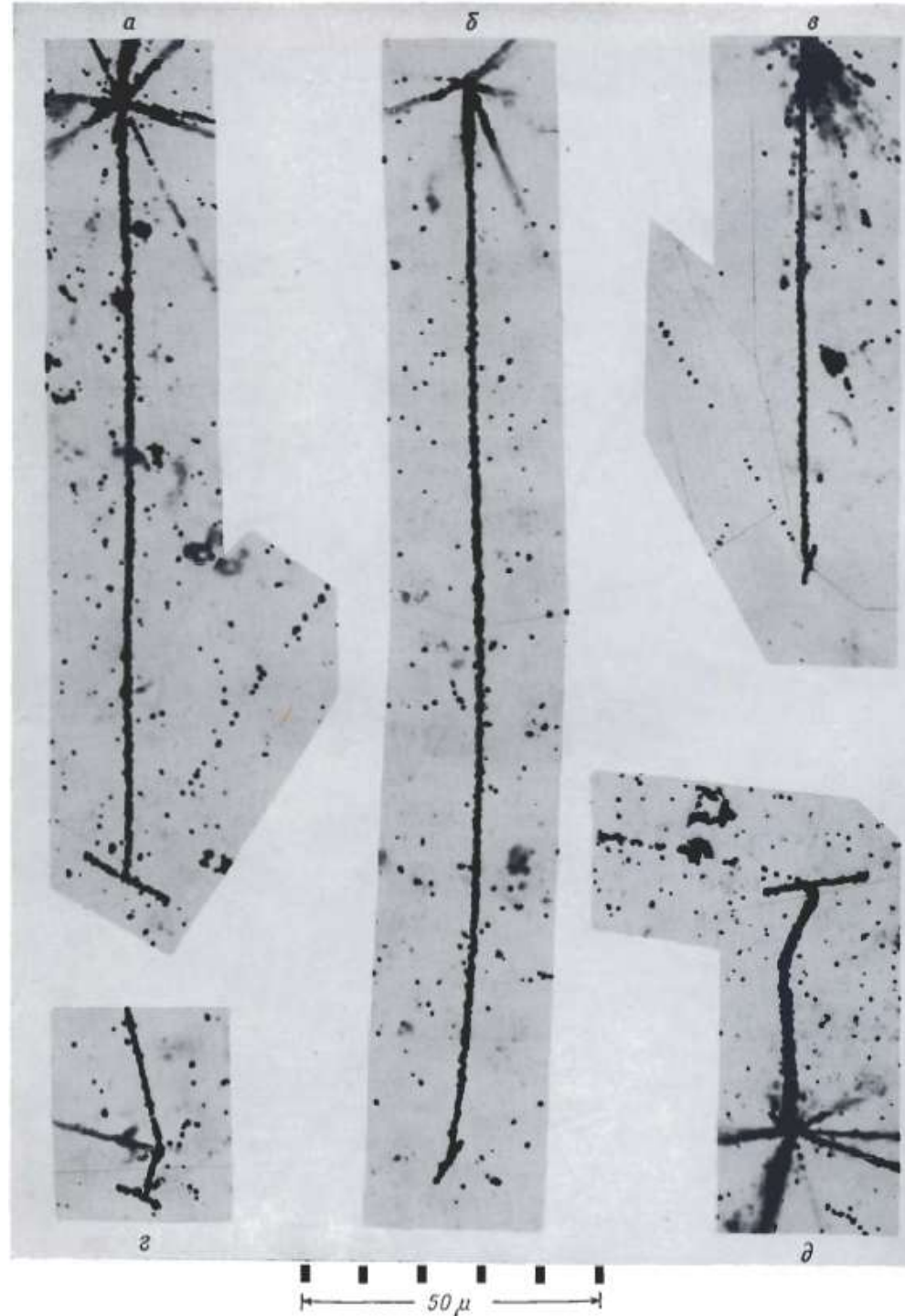


FIG. 1. A photograph of an event interpreted as the beta decay of  $C^9$ . The  $C^9$  nucleus (track  $F$ ) was produced in star ( $A$ ) and disintegrated into a proton, two alpha particles, and a positron (tracks 1, 2, 3, and 4, respectively).

# Hammer tracks in cosmic ray events:

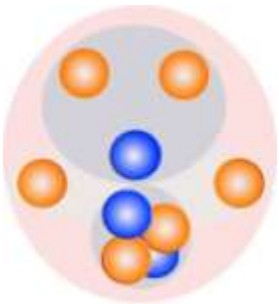
${}^8\text{Be}$  produced in  
 $\beta$ -delayed decay of  
stopped  ${}^8\text{B}$  and  ${}^8\text{Li}$

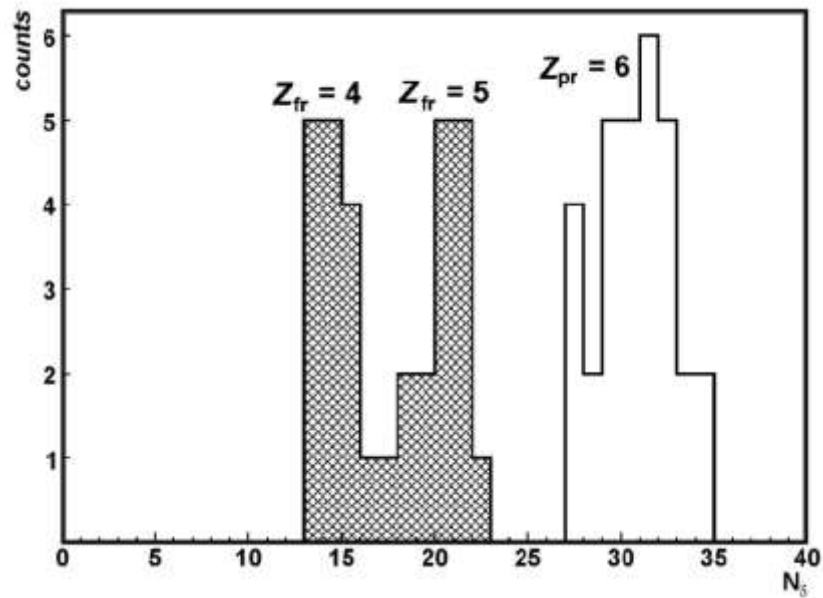


# $^9\text{C}$ exposure

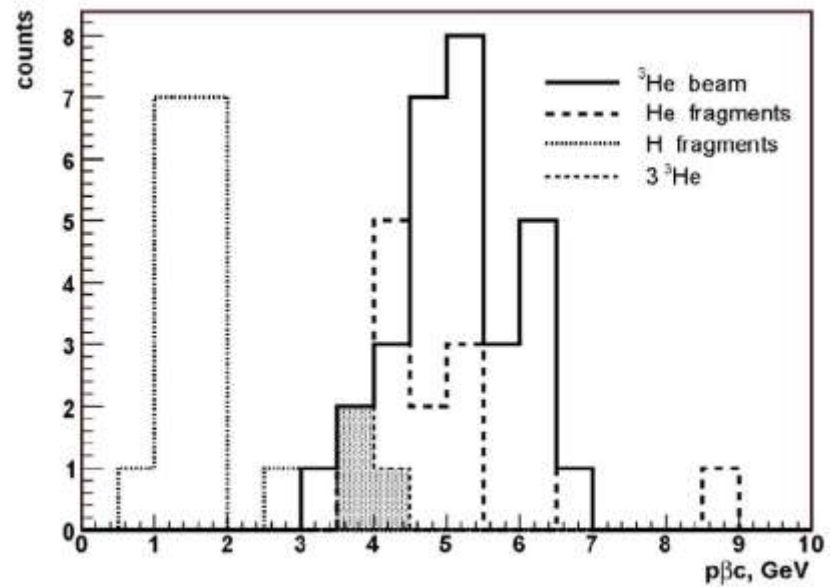
Distribution of 160 "white" stars  $N_{\text{ws}}$  in charge configurations  $\Sigma Z_{\text{fr}}$

$\Sigma Z_{\text{fr}}$	$Z_{\text{fr}}$						$N_{\text{ws}}$
	6	5	4	3	2	1	
7	-	-	-	-	-	7	1
7 ( $Z_{\text{pr}} = 7$ )	-	1	-	-	-	2	1
7	-	-	-	-	1	5	3
7	-	-	-	-	2	3	5
6 ( $Z_{\text{pr}} = 6$ )	-	1	-	-	-	1	15
6 ( $Z_{\text{pr}} = 6$ )	-	-	1	-	-	2	16
6	-	-	-	-	3	-	16
6	-	-	-	1	-	3	2
6	-	-	-	-	1	4	28
6	-	-	-	-	2	2	24
6	-	-	-	-	-	6	6
5 ( $Z_{\text{pr}} = 5$ )	-	-	1	-	-	1	2
5 ( $Z_{\text{pr}} = 5$ )	-	-	-	1	-	2	3
5	-	-	-	-	-	5	2
5	-	-	-	-	1	3	19
5	-	-	-	-	2	1	13
4	-	-	-	-	1	2	5

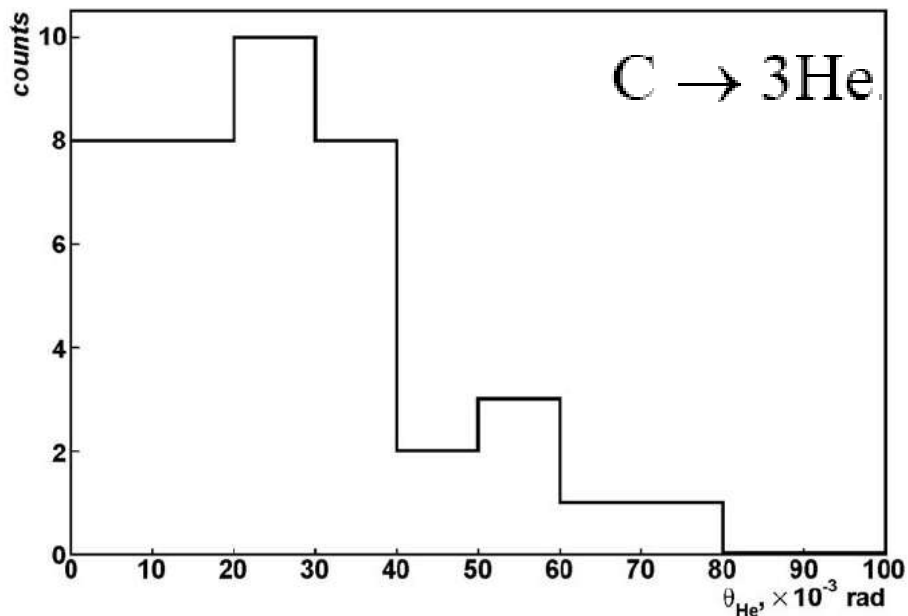




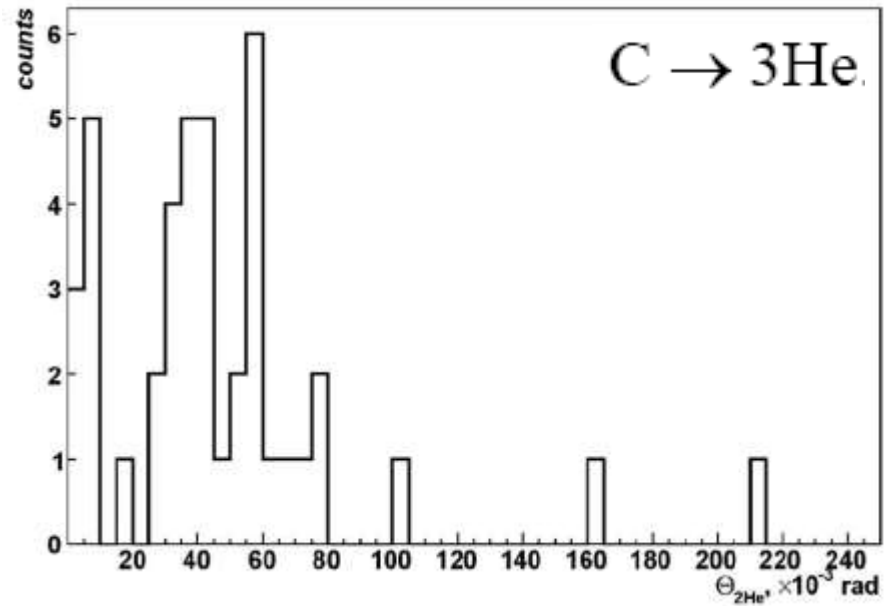
$\delta$ -electrons density for beam particles and relativistic fragments with charges  $Z_{fr} > 2$  from "white" stars  $\sum Z_{fr} = 5 + 1$  and  $4 + 1 + 1$



$p\beta c$  for beam  ${}^3\text{He}$  nuclei, H fragments of the "white" stars  $\sum Z_{fr} = 5 + 1$  and  $4 + 1 + 1$ , He fragments of the "white" stars  $3\text{He}$  and from the  ${}^3\text{He}$  event



Polar angles  $\theta$  for doubly charged fragments in the "white" stars  $\text{C} \rightarrow 3\text{He}$



Opening angles  $\Theta_{2\text{He}}$  between fragments in the "white" stars  $\text{C} \rightarrow 3\text{He}$



## ELEMENTARY PARTICLES AND FIELDS

### Experiment

# Fragmentation Channels of Relativistic ${}^7\text{Be}$ Nuclei in Peripheral Interactions

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V. A. Dronov<sup>1)</sup>, P. I. Zarubin<sup>2)</sup>, I. G. Zarubina<sup>2)</sup>, A. D. Kovalenko<sup>2)</sup>, V. G. Larionova<sup>†1)</sup>,  
A. I. Malakhov<sup>2)</sup>, P. A. Rukoyatkin<sup>2)</sup>, V. V. Rusakova<sup>2)</sup>, S. P. Kharlamov<sup>1)</sup>, and V. N. Fetisov<sup>1)</sup>

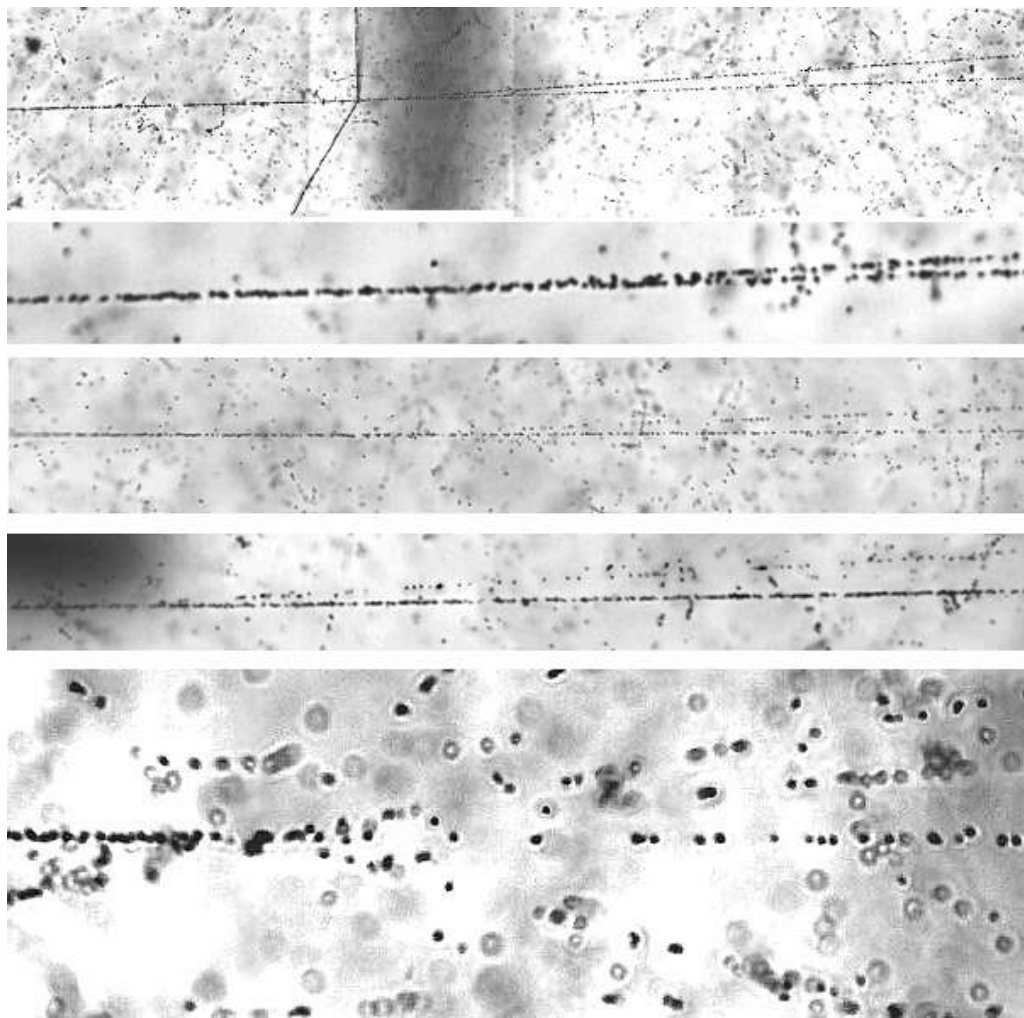
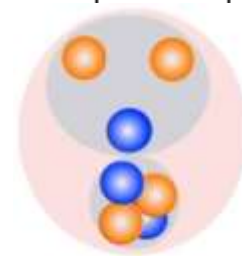
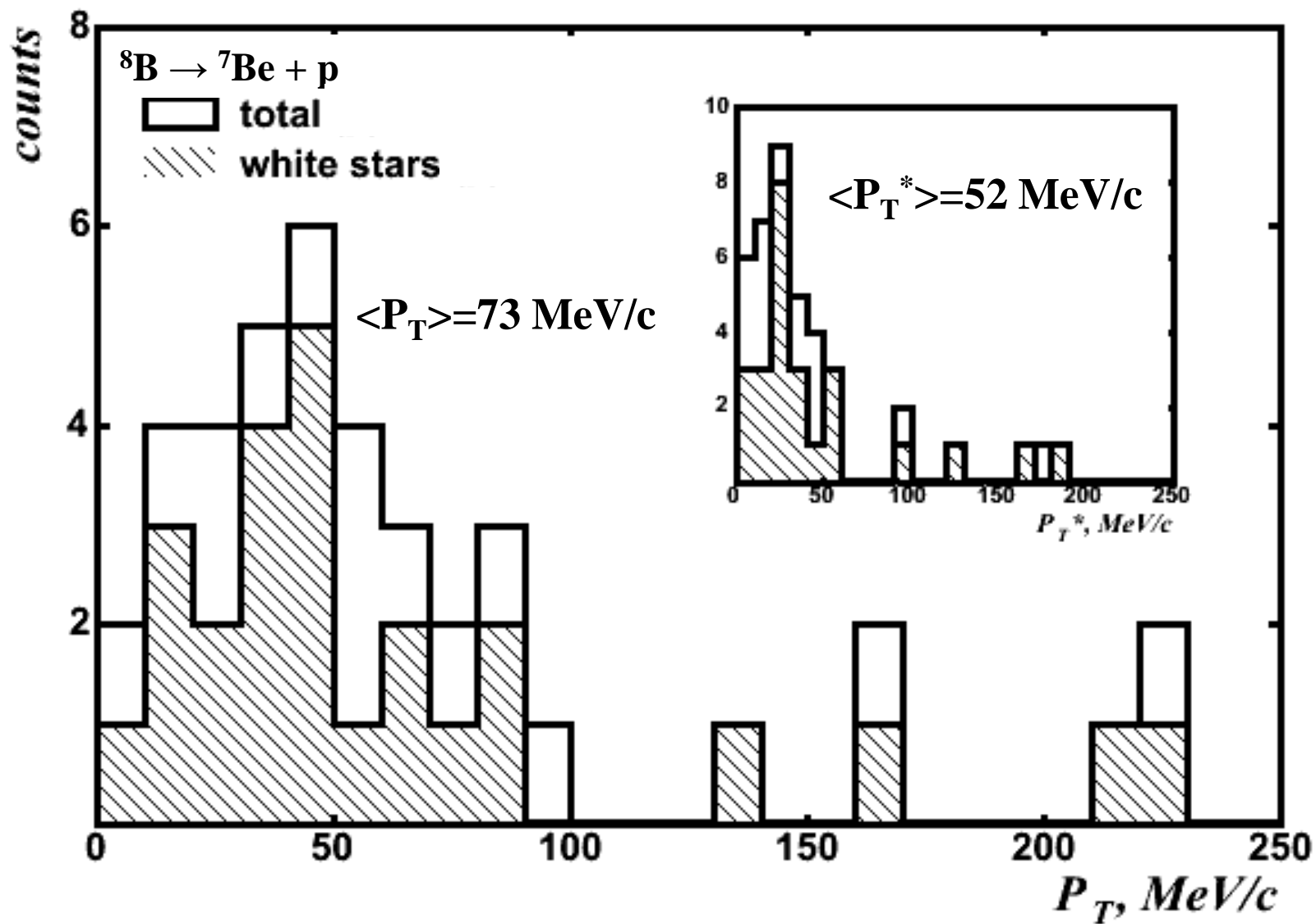


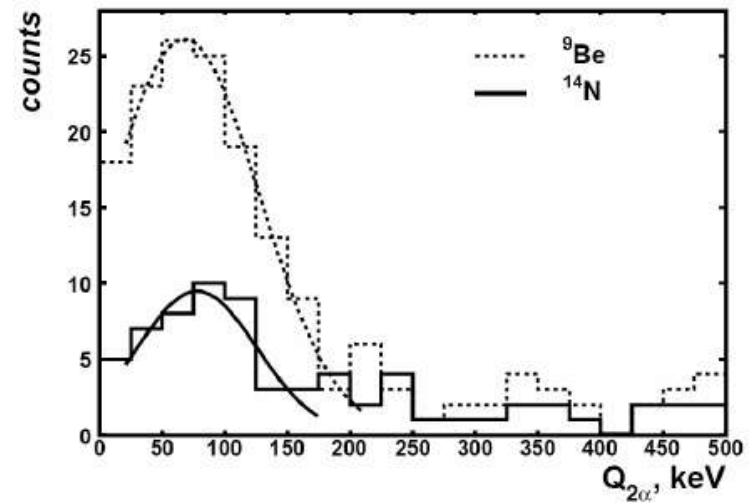
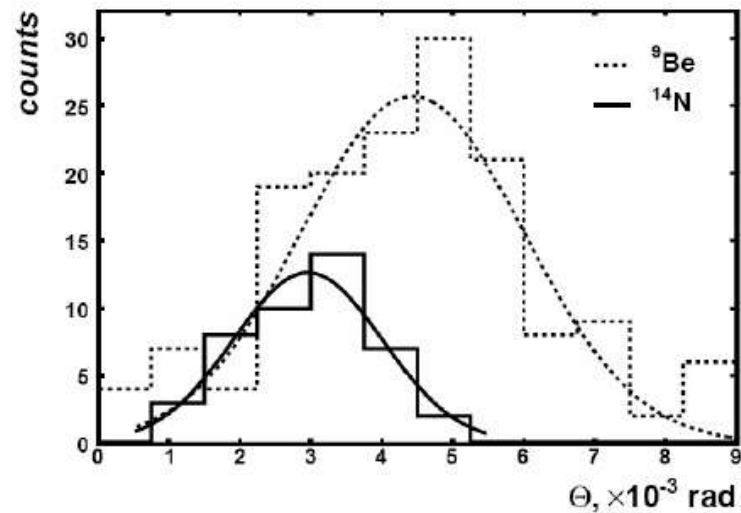
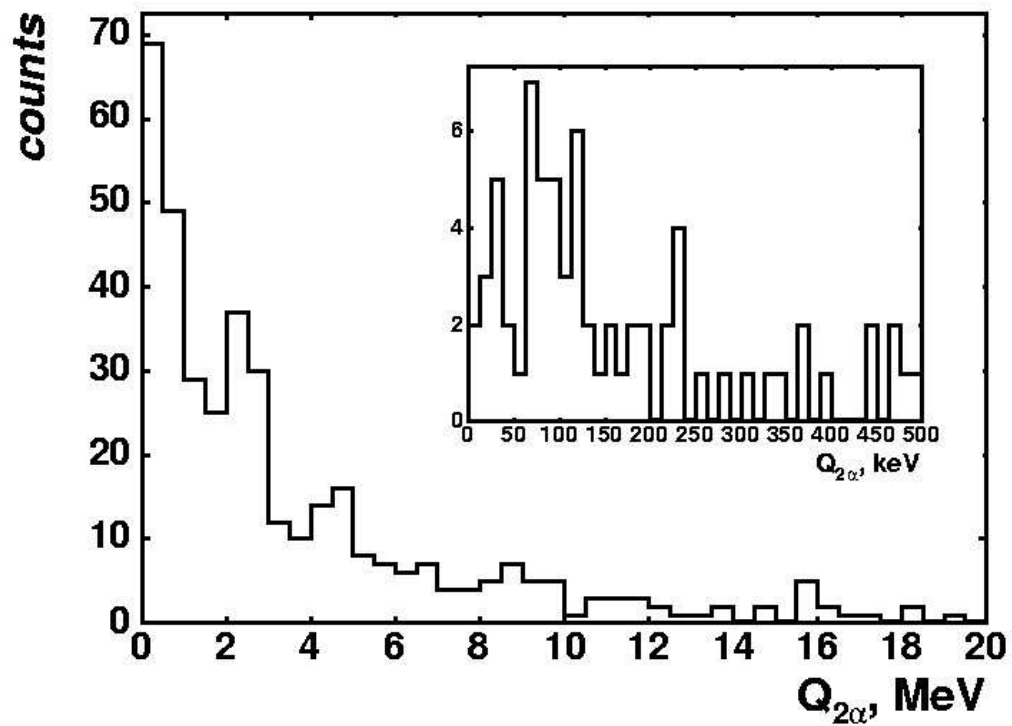
TABLE III:  ${}^7\text{Be}$  fragmentation channel (number of events)**MeV****1.6****6.9****25.3****21.2****5.6**

Channel	2He		He+2H		4H		Li+H		Sum
	$n_b = 0$	$n_b > 0$	$n_b = 0$	$n_b > 0$	$n_b = 0$	$n_b > 0$	$n_b = 0$	$n_b > 0$	
${}^3\text{He}+{}^4\text{He}$	30	11							41
${}^3\text{He}+{}^3\text{He}$	11	7							18
${}^4\text{He}+2\text{p}$			13	9					22
${}^4\text{He}+\text{d}+\text{p}$			10	5					15
${}^3\text{He}+2\text{p}$			9	9					18
${}^3\text{He}+\text{d}+\text{p}$			8	10					18
${}^3\text{He}+2\text{d}$			1						1
${}^3\text{He}+\text{t}+\text{p}$			1						1
$3\text{p}+\text{d}$					2				2
$2\text{p}+2\text{d}$					1				1
${}^6\text{Li}+\text{p}$							9	3	12
Sum	41	18	42	33	2	1	9	3	149





# NUCLOTRON: 2.9A GeV/c $^{14}\text{N}$





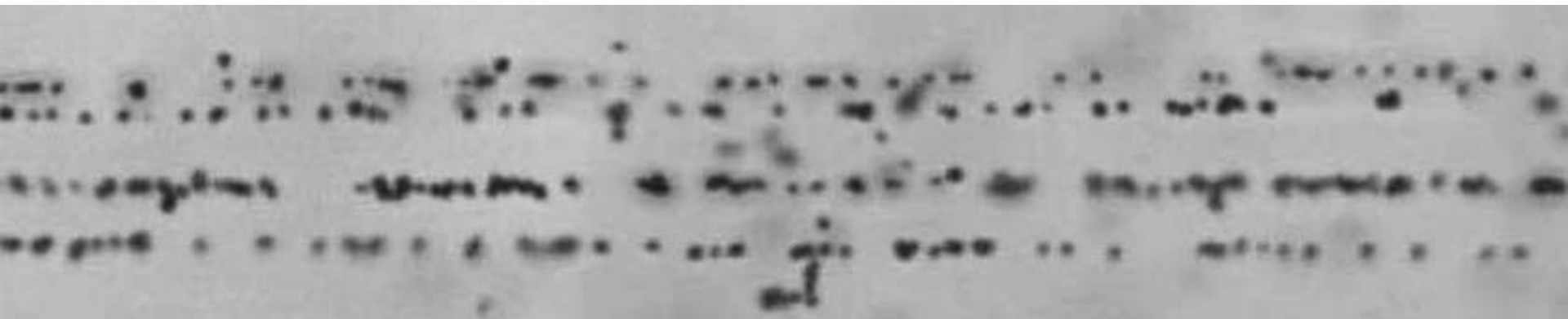
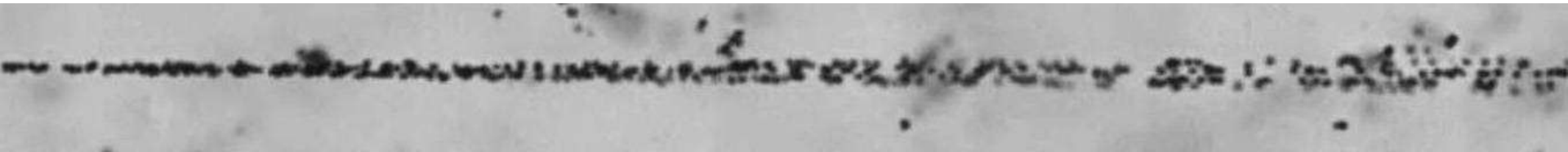
# $T_\alpha : 0.05-2 \text{ MeV}$

The detection of such “ultracold”  $5\alpha$  states is a serious argument in favor of the reality of the phase transition of  $\alpha$  clusterized nuclei to the dilute Bose gas of  $\alpha$ -particles. It gives a special motivation to explore lighter  $n\alpha$  systems produced as potential “building blocks” of the dilute  $\alpha$ -particle Bose gas.

$\alpha$

# $\lambda : 10-1 \text{ fm}$

**3.65A GeV  $^{20}\text{Ne}$  2+2+2+2+2**

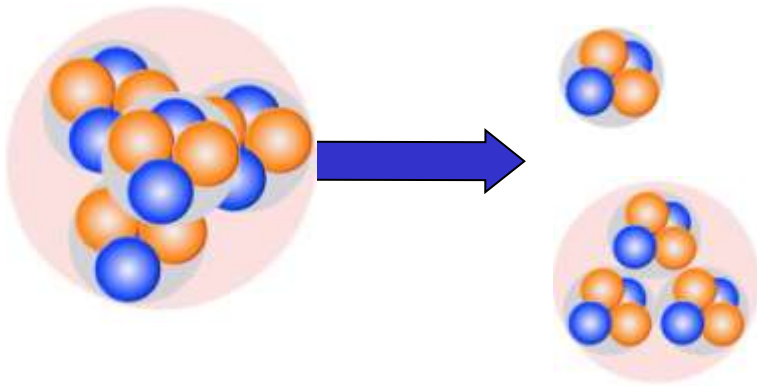


$n_b$	0	0	1	2	3	>3
$n_g$	0	1	0	0	0	0
F + H	26 (19.5)	9 (15.0)	13 (44.8)	2	-	1
O + He	54 (40.6)	19 (31.7)	2 (6.9)	-	1	1
O + 2H	12 (9.0)	7 (11.7)	-	-	-	-
N + He + H	12 (9.0)	7 (11.7)	4 (13.8)	1	-	-
N + 3H	3 (2.3)	3 (5.0)	-	-	-	-
C + 2He	5 (3.8)	3 (5.0)	3 (10.3)	1	-	-
C + 2He + 2H	5 (3.8)	3 (5.0)	3 (10.3)	-	-	-
C + 4H	2 (1.0)	-	-	-	-	-
B + Li + H	1 (0.8)	-	-	-	-	-
B + 2He + H	2 (1.5)	1 (1.7)	-	-	-	-
B + He + 3H	2 (1.5)	1 (1.7)	-	-	-	-
B + 5H	1 (0.8)	-	1 (3.4)	-	-	-
2Be + 2H	-	1 (1.7)	-	-	1	-
Be + Li + 3H	1 (0.8)	-	-	-	-	-
Be + 3He	2 (1.5)	-	-	-	-	-
Be + He + 4H	1 (0.8)	-	-	-	-	-
Li + 3He + H	-	1 (1.7)	-	-	-	-
5He	3 (2.3)	-	1 (3.4)	2	-	1
4He + 2H	1 (0.8)	5 (8.3)	2 (6.9)	-	-	-

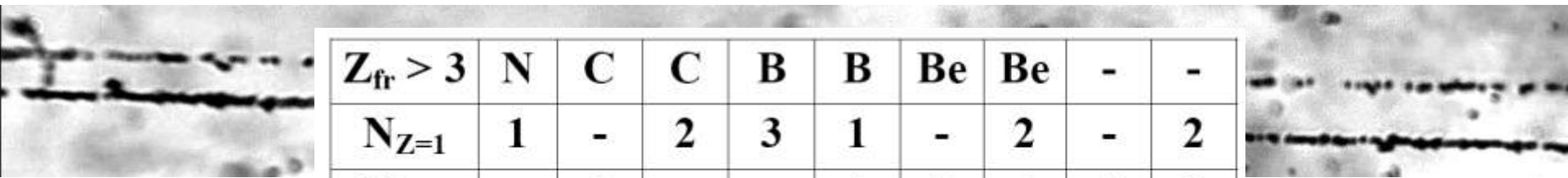
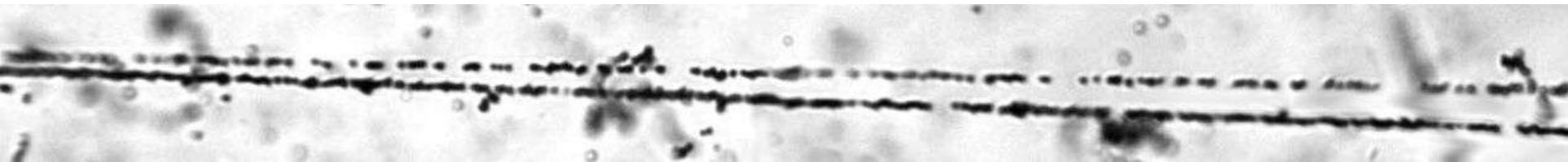
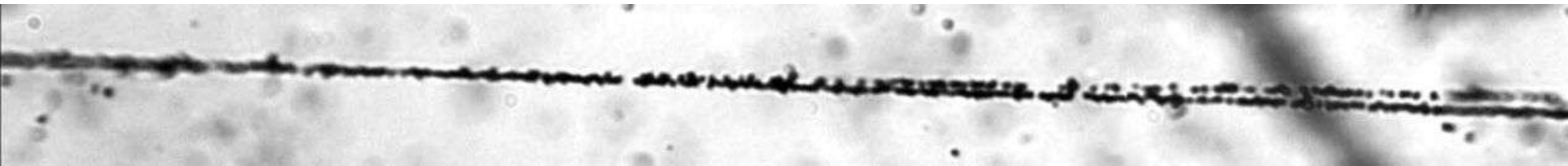
**$^{22}\text{Ne}$  3.22A GeV**

**4100 Inelastic Interactions**





4.5 A GeV/c  $^{16}\text{O}$



$Z_{fr} > 3$	N	C	C	B	B	Be	Be	-	-
$N_{Z=1}$	1	-	2	3	1	-	2	-	2
$N_{Z=2}$	-	1	-	-	1	2	1	4	3
$N_{ev}$	18	21	7	2	10	1	1	9	3