

Гравитационные состояния ультрахолодных квантовых систем

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ПЛАН

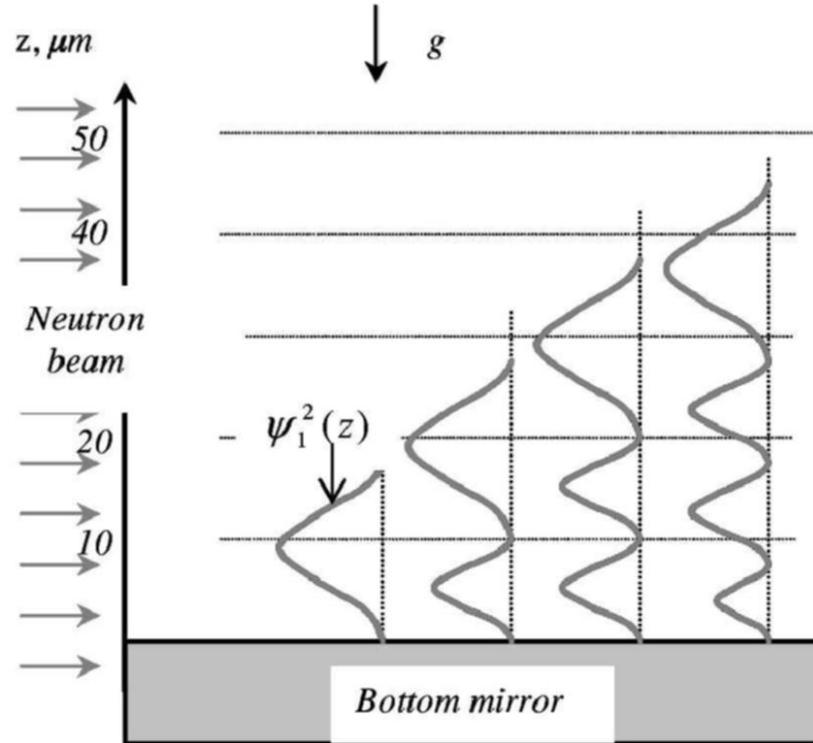
- Квантовые состояния ультрахолодных нейтронов в гравитационном поле Земли-спектрометр ГРАНИТ
- Поиск «нестандартной физики»: ограничения на характеристики аксионного поля из гравитационных экспериментов с УХН
- Гравитационные состояния Антиводорода и измерение гравитационной массы

Гравитационные состояния УХН

$$\left\{ \begin{array}{l} \left[-\frac{\hbar^2}{2M} \frac{\partial^2}{\partial z^2} + Mgz - \varepsilon_n \right] \varphi_n(z) = 0 \\ \varphi_n(0) = 0 \end{array} \right.$$

$$\varphi_n(z) = C_n \text{Ai}(z/l_0 - \lambda_n)$$

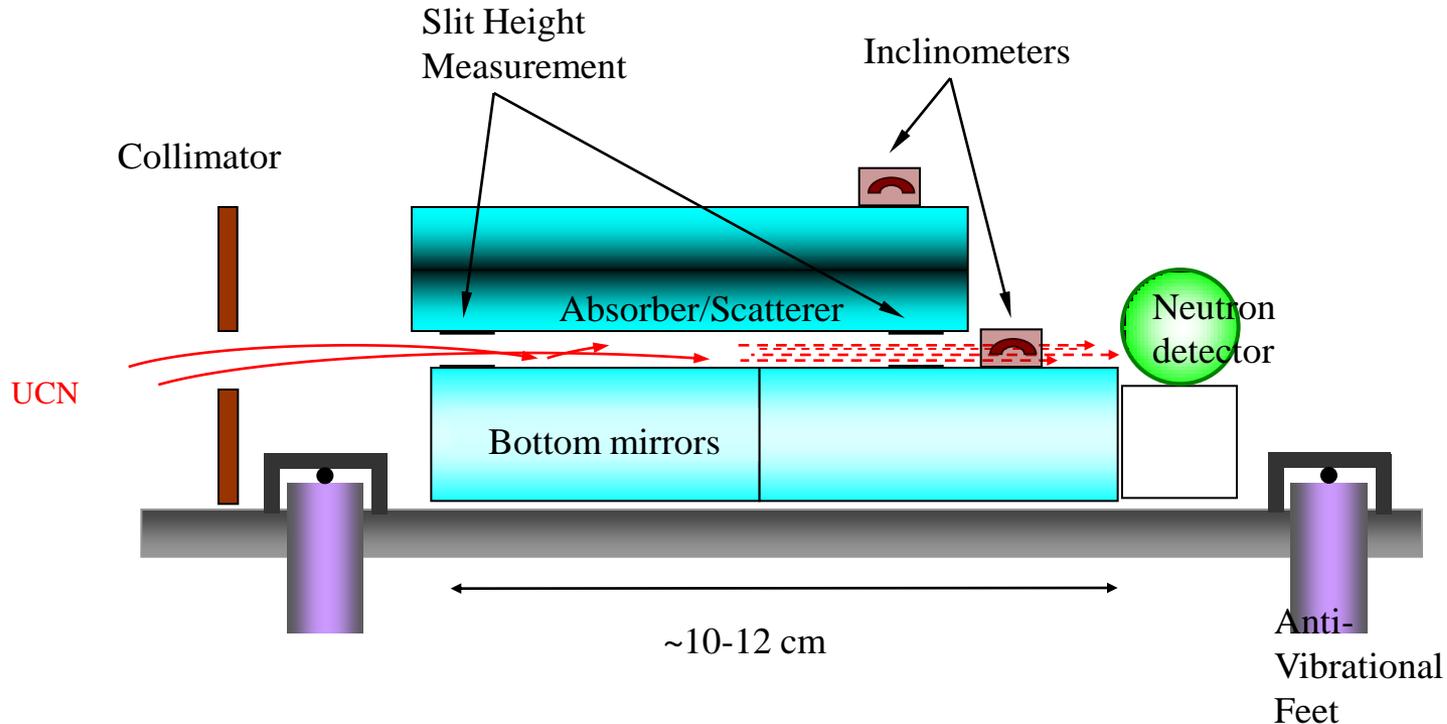
$$\text{Ai}(-\lambda_n) = 0$$



$$\varepsilon_n = \lambda_n \varepsilon \quad \varepsilon = \sqrt[3]{\frac{\hbar^2 M g^2}{2}} = 0.61 \cdot 10^{-12} \text{ eV} \quad l_0 = \sqrt[3]{\frac{\hbar^2}{2M^2 g}} = 5.87 \cdot 10^{-6} \text{ m}$$

Gravitational Bound states – The experiment

Nesvizhevsky V.V. et. al. Nature 415, 297 (2002)



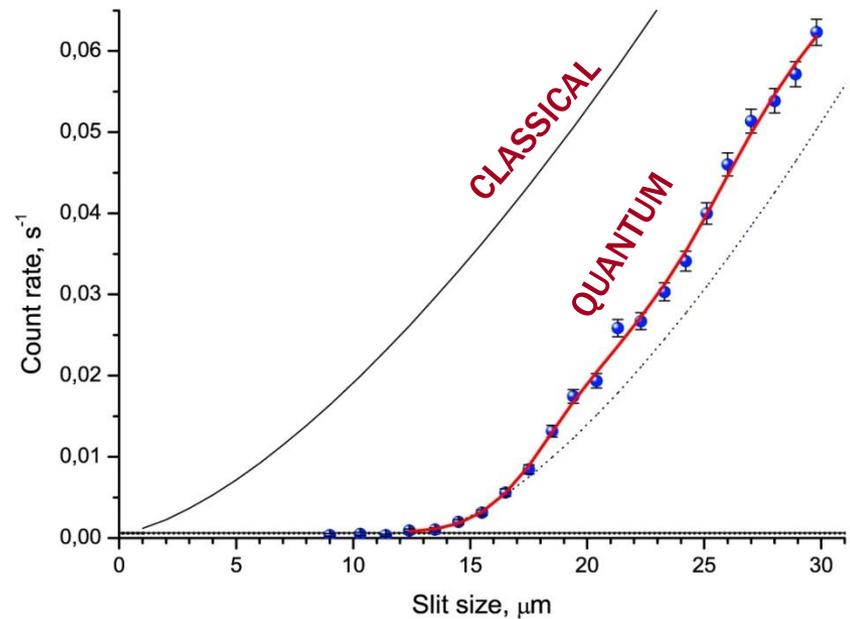
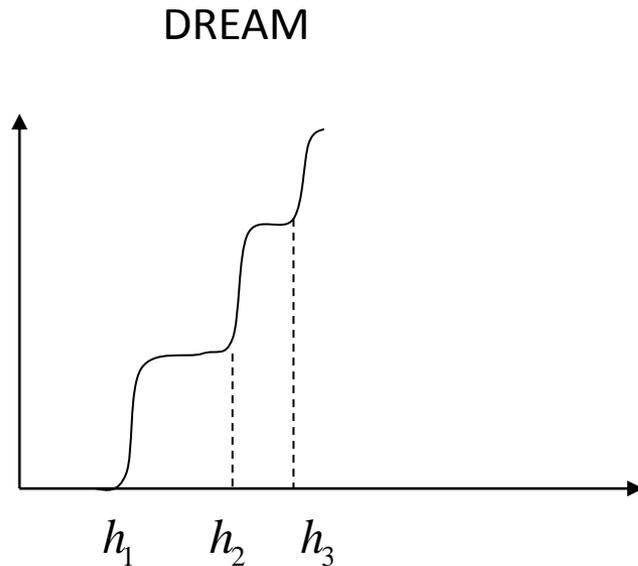
- Count rates at ILL turbine: $\sim 1/s$ to $1/h$
- Effective (vertical) temperature of neutrons is ~ 20 nK
- Background suppression is a factor of $\sim 10^8$ - 10^9
- Parallelism of the bottom mirror and the absorber/scatterer is $\sim 10^{-6}$

$$z_1^{\text{exp}} = 12.2 \pm 1.8_{\text{sys}} \pm 0.7_{\text{stat}} \mu\text{m}$$

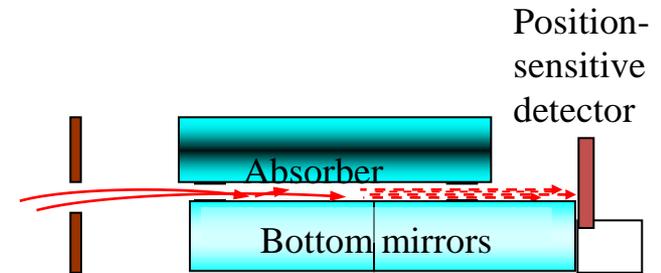
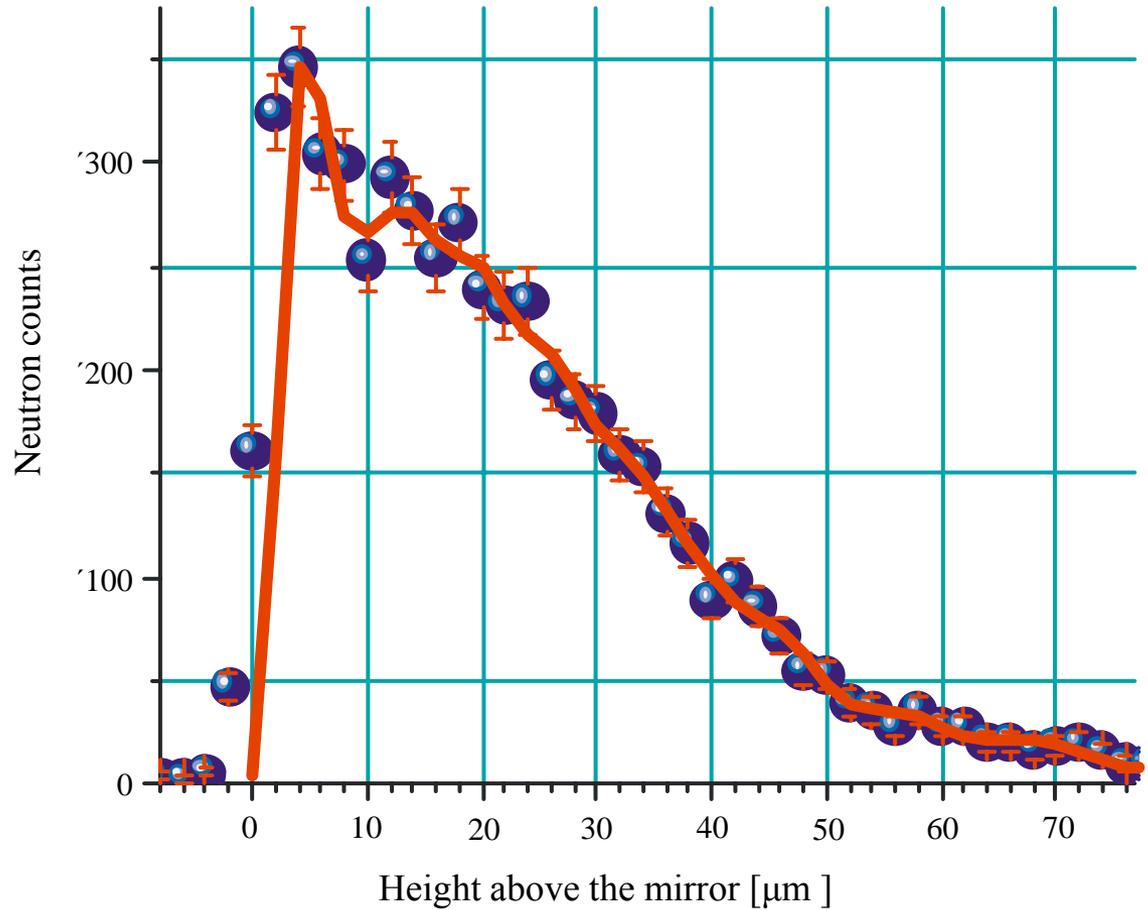
$$z_1^{\text{q.c.}} = \frac{3}{2} \langle 1|z|1 \rangle = 13.7 \mu\text{m}$$

$$z_2^{\text{exp}} = 21.6 \pm 2.2_{\text{sys}} \pm 0.7_{\text{stat}} \mu\text{m}$$

$$z_2^{\text{q.c.}} = \frac{3}{2} \langle 2|z|2 \rangle = 24 \mu\text{m}$$



Results with the Position-Sensitive Detector



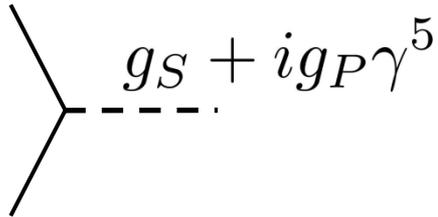
AXION and UCN gravitational states

- **Strong CP-problem-** explanation of extremely small electric dipole moment of neutron
- Pseudo-scalar neutral light bozon

$$10^{-6} < M_a < 10^{-3} eV ; 1.3 \cdot 10^{-4} < \lambda < 1.3 \cdot 10^{-1} m$$

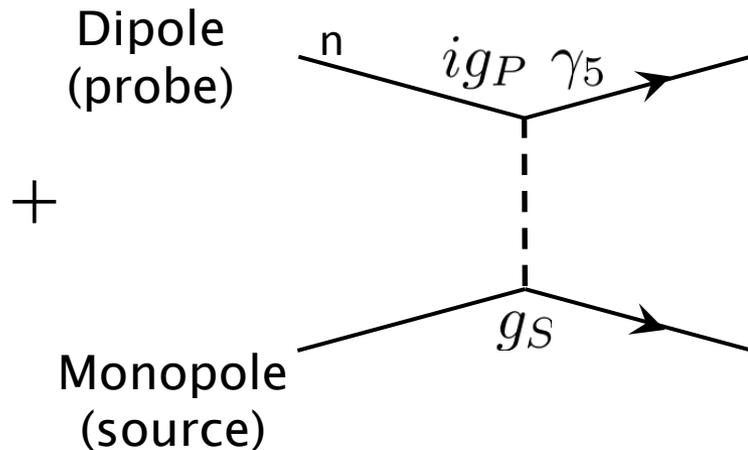
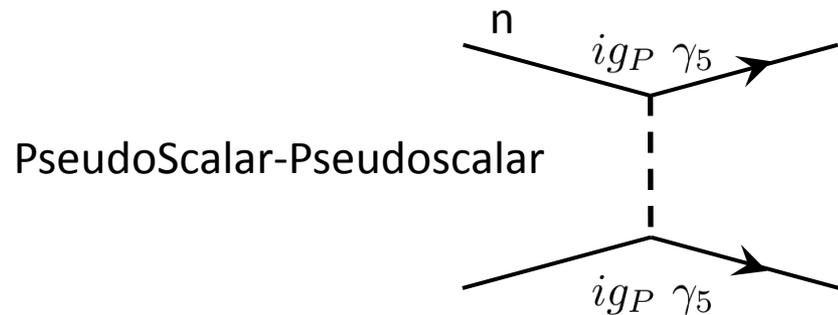
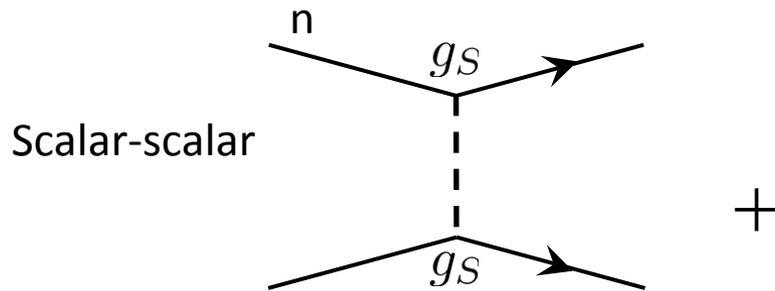
- Axion coupling with electron, photons and **nucleon**

Short-range spin-dependent forces



interactions of range λ
mediated by new light boson of mass M

$$\lambda = \frac{\hbar}{Mc}$$



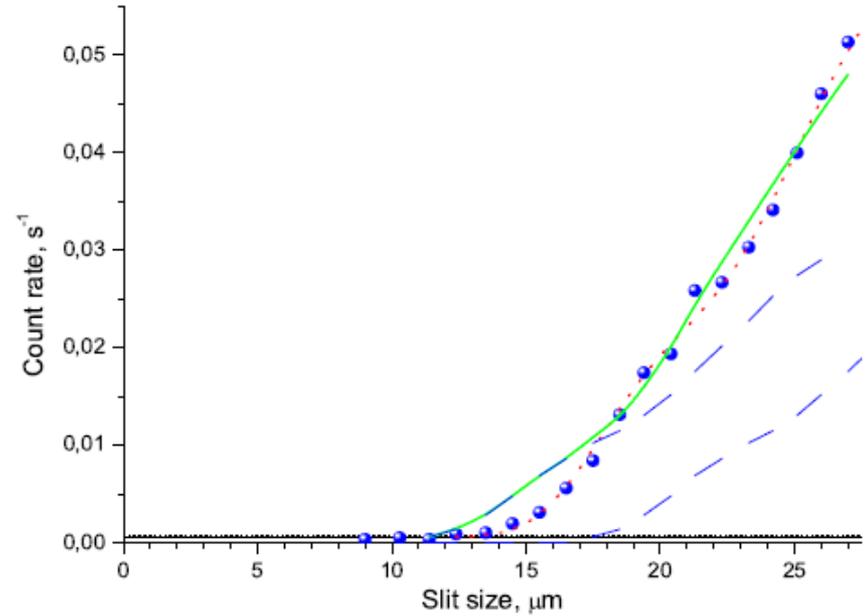
~~P~~ ~~CP~~ Scalar-Pseudoscalar (Axion-like)

$$V_{SP}(\mathbf{r}) = \frac{g_S g_P}{8\pi} \frac{\hbar}{m} \boldsymbol{\sigma} \cdot \mathbf{r} \left(\frac{1}{\lambda} + \frac{1}{r} \right) \frac{e^{-r/\lambda}}{r^2}$$

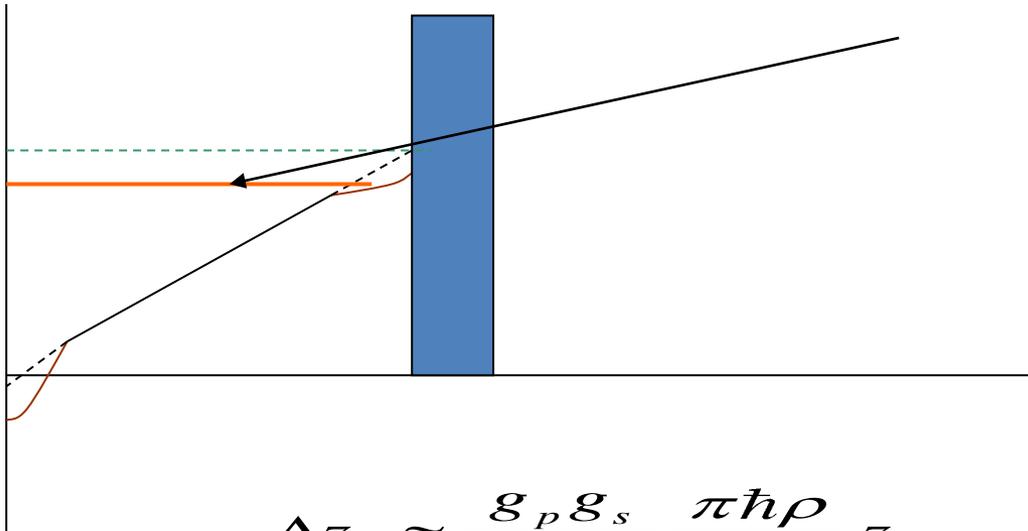
Axion-nucleon interaction:

$$V(\vec{r}) = \frac{\hbar g_p g_s}{8\pi m c} (\vec{\sigma} \vec{n}) \left[\frac{1}{\lambda r} + \frac{1}{r^2} \right] \exp(-r / \lambda)$$

$$V_A(z) = \int_{\text{mirror}} V(x, y, z + z') dz' = \frac{g_p g_s}{4\pi} \frac{\pi \hbar \rho \lambda}{2m^2 c} \exp(-z / \lambda)$$



Level and turning point shift



$$\Delta z_n \approx \frac{g_p g_s}{3\pi m g} \frac{\pi \hbar \rho}{2m^2 c} z_n$$

Effect on Gravitationally Bound States

Integration of 2nd potential over mirror:

$$V(z) = -g_S^N g_P^n \frac{\hbar \rho_m \lambda}{8m_n^2 c} \exp(-z/\lambda) \underbrace{(\sigma_n \cdot \hat{z})}_{\pm 1}$$

Inclusion of absorber:

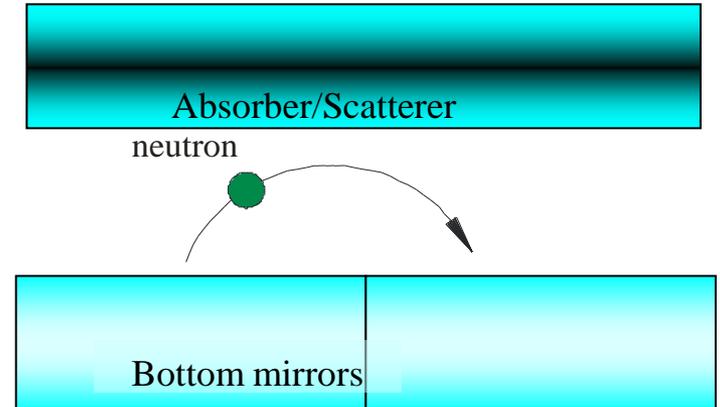
$$W(z) = \pm g_S^N g_P^n \frac{\hbar \rho_m \lambda}{8m_n^2 c} \underbrace{\left[\exp(-z/\lambda) - \exp(-(\Delta h - z)/\lambda) \right]}_{\frac{2z}{\lambda} + \text{const.}}$$

After dropping the invisible constant piece,

$W(z)$ is linear in z

$$g \rightarrow g_{\text{eff}} = g \pm g_S^N g_P^n \frac{2\hbar \rho_m}{8m_n^3 c}$$

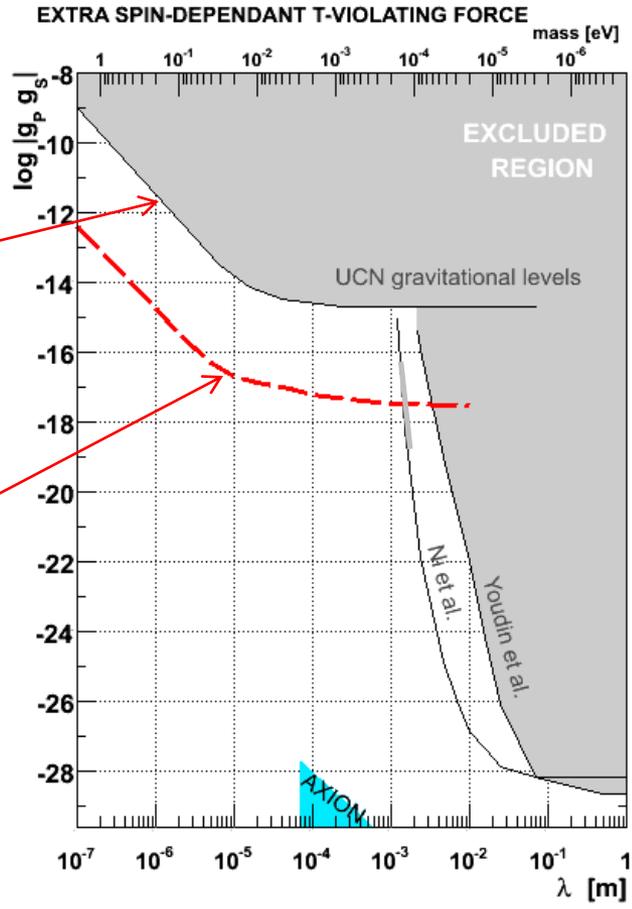
Our limits are calculated from a shift of the turning point by 3 μm .



$$z_1 = 2.34 \sqrt[3]{\frac{\hbar^2}{2m^2 g}} = 13.7 \mu\text{m}$$

$$z_2 = 4.09 \sqrt[3]{\frac{\hbar^2}{2m^2 g}} = 24.0 \mu\text{m}$$

Exclusion plot

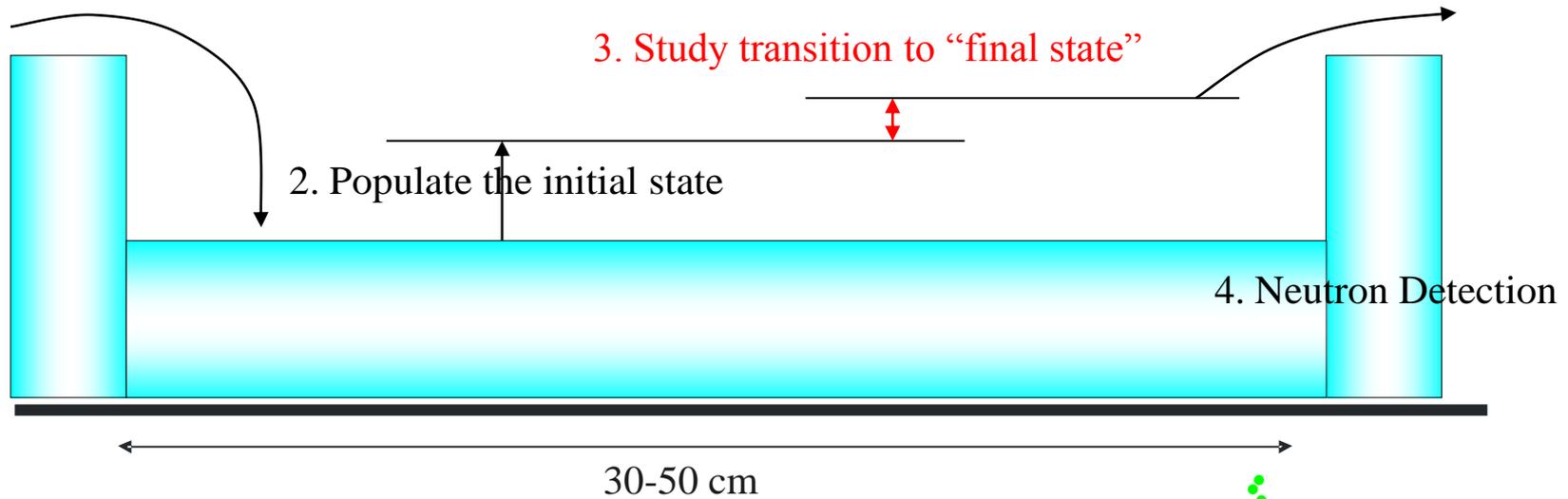


Flow-Through Mode

Trap Mode

GRANIT spectrometer

1. Population of ground state

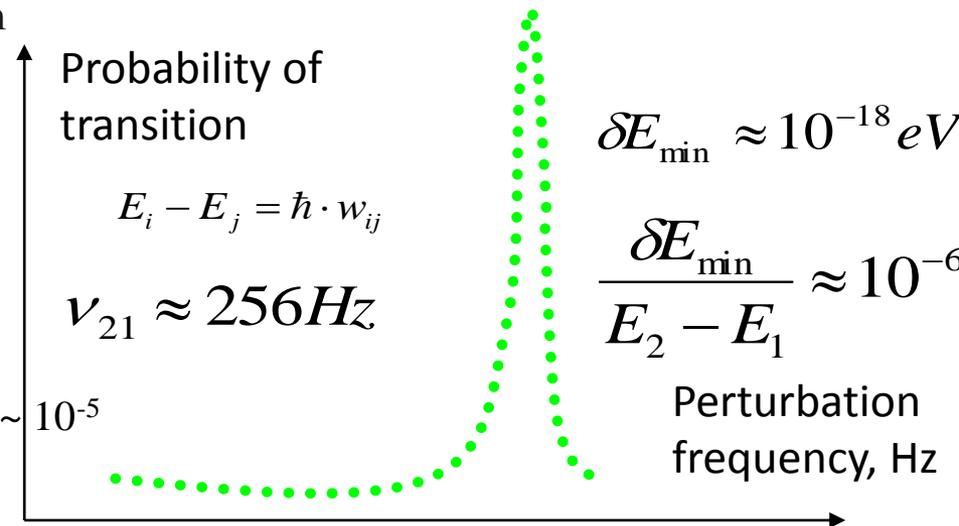


If lifetime is $\tau_n \sim 500$ s, $\frac{\Delta E}{E} \sim \frac{\hbar}{\tau_n E} = 2 \cdot 10^{-6}$

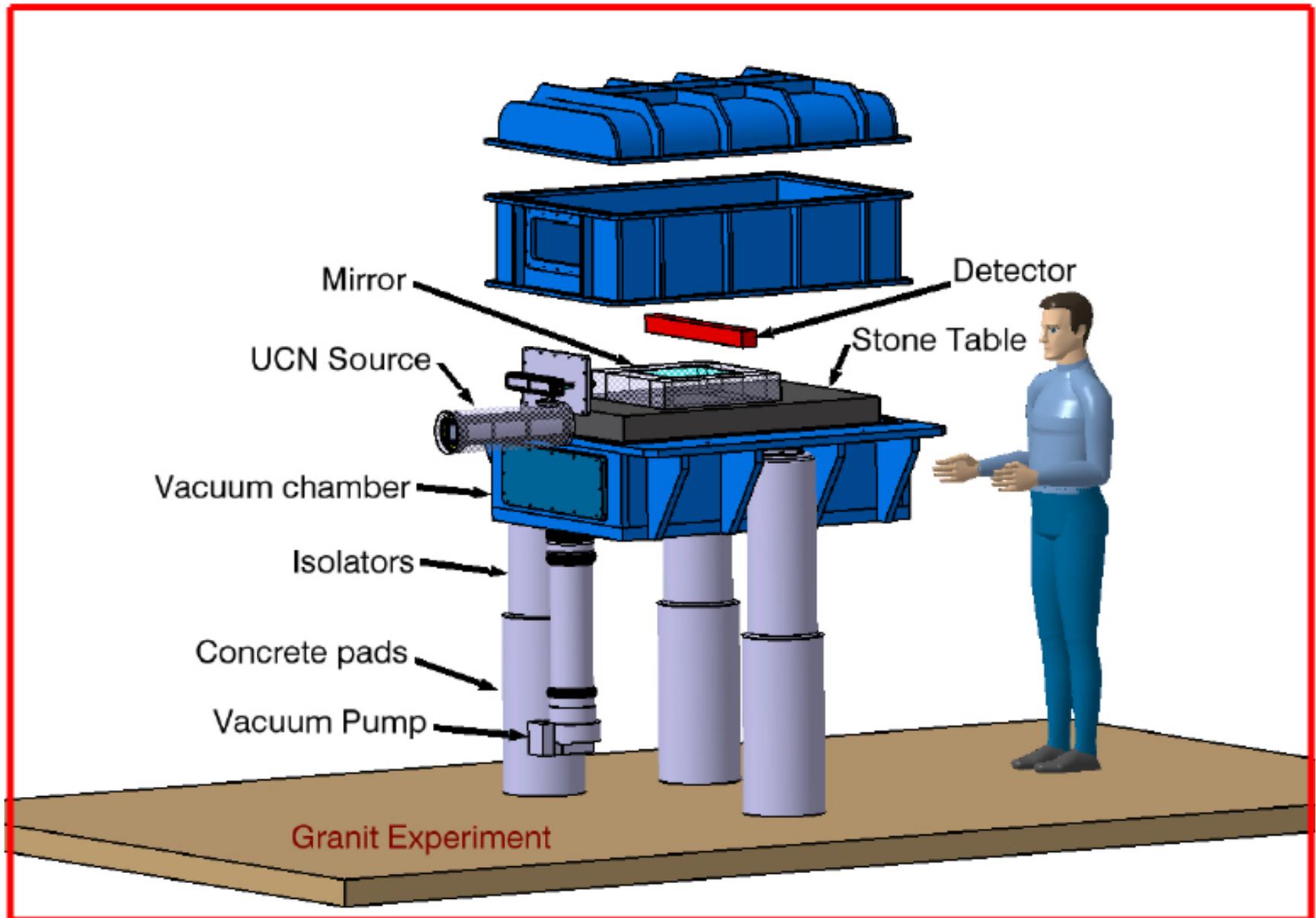
Flatness of bottom mirror: < 100 nm

Accuracy of setting the side walls perpendicular: $\sim 10^{-5}$

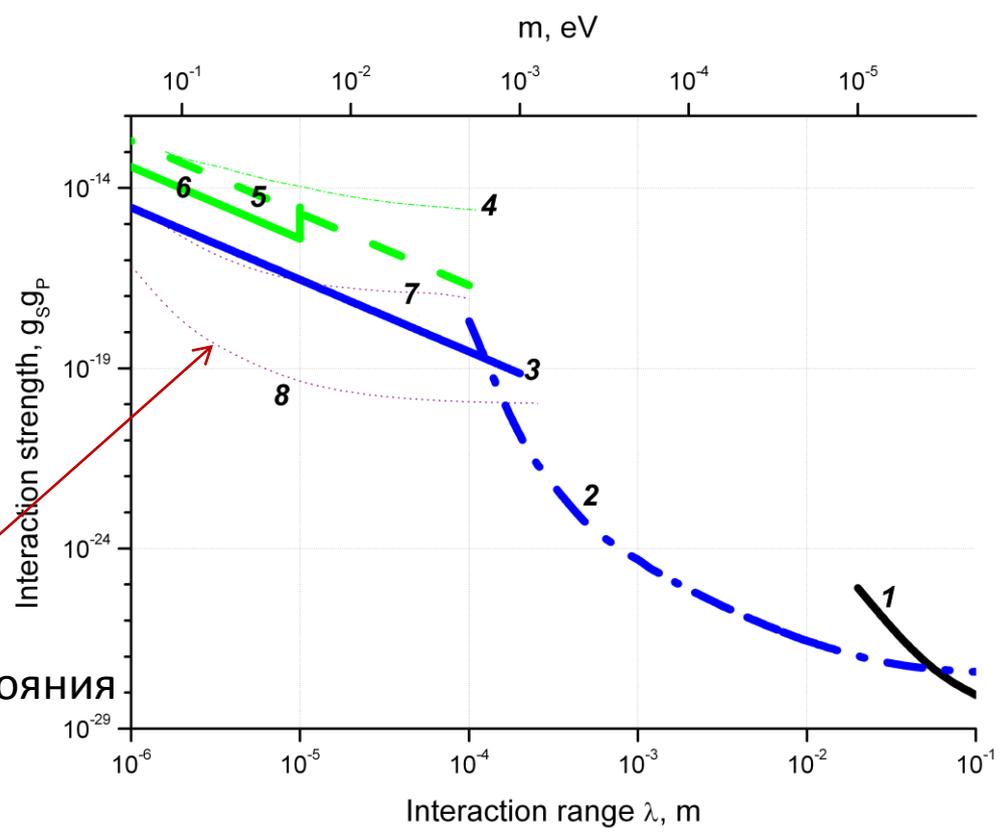
Vibrations, Count Rate, Holes, ...



GRANIT spectrometer

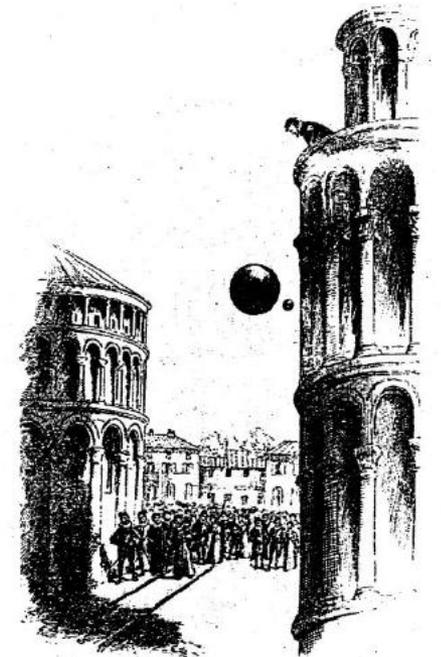


Гравитационные состояния



Гравитационные состояния Антиводорода

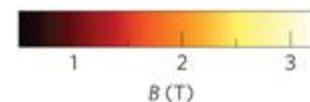
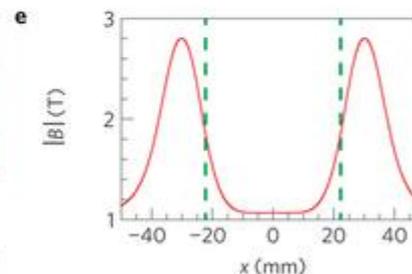
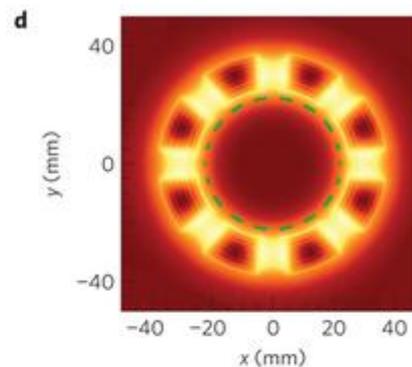
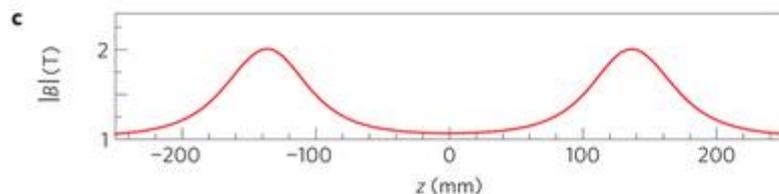
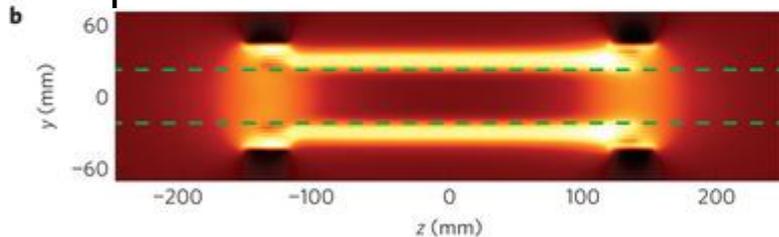
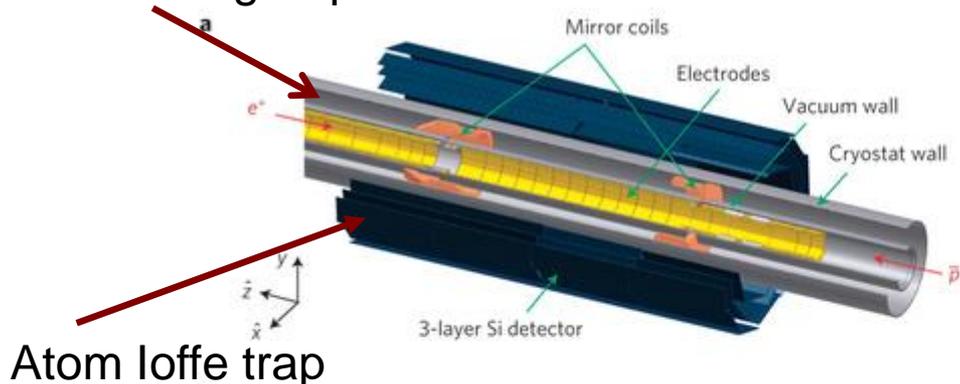
$$m = M?$$



ALPHA collaboration Nature Physics (April, 2011)

309 атомов антиводорода удержаны в течении 1000 с

Penning trap



Глубина ямы для атомов 0.54 К

Gbar : \bar{g} experiment using \bar{H}^+ to get \bar{H} atoms

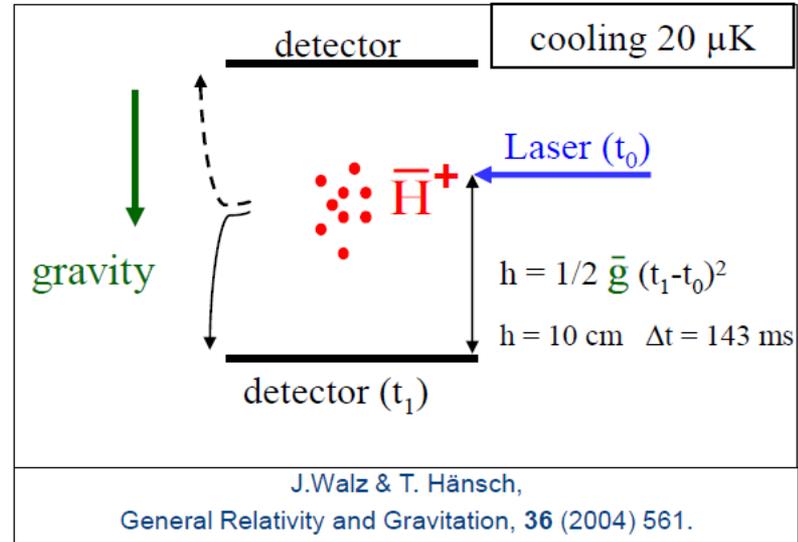
irfu

cea

saclay

- Produce ion \bar{H}^+
- Capture ion \bar{H}^+
- Sympathetic cooling 20 μK
- Photodetachment of e^+
- Time of flight

Error dominated by temperature of \bar{H}^+

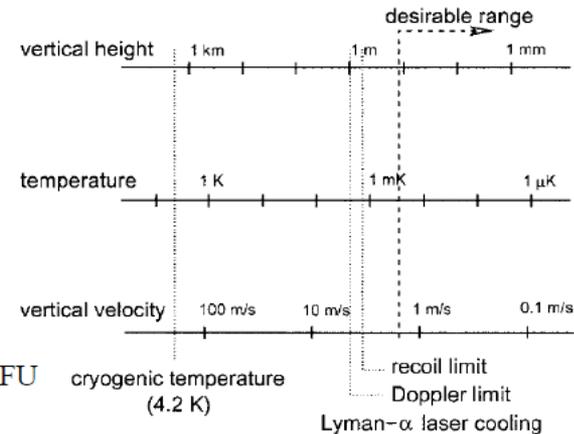


Relative Precision on \bar{g} :

\bar{H}^+ in ion trap	$\Delta g/g$
$5 \cdot 10^5$	0.001
10^4	0.006
10^3	0.02

27/04/2011

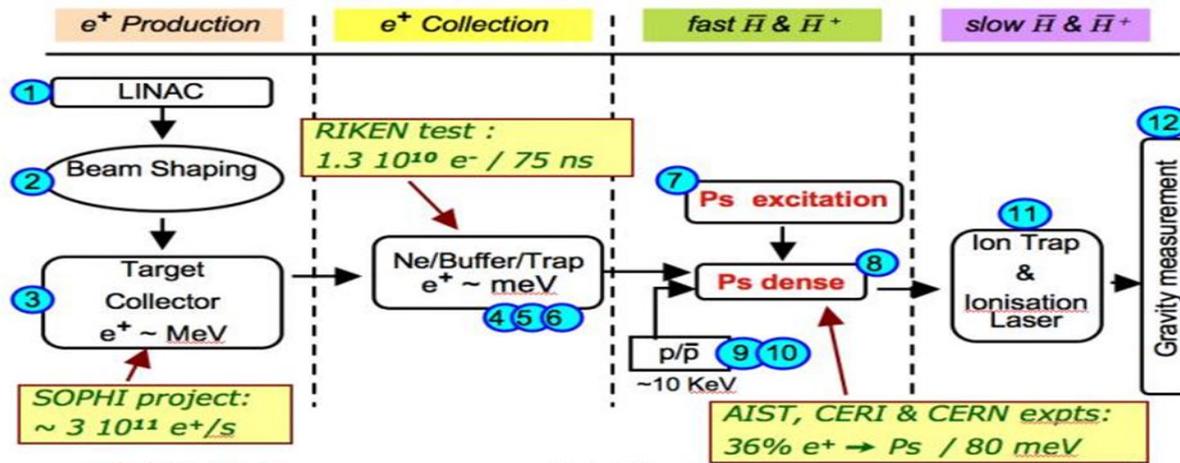
Pascal Debu - CEA/DSM/IRFU



GBar Project CERN approved may2012

гравитационные свойства антиводорода

Прямой тест Принципа Эквивалентности с
антивеществом при $T \sim 1$ nK



2013

2015

ФИАН

Квантовое отражение ультрахолодных антиатомов
Спектроскопия долгоживущих гравитационных состояний

\bar{H} + WALL

~~=~~

ANNIHILATION ?

A. Yu. Voronin, P. Froelich, and B. Zygelman, *Phys. Rev. A* **72**,
062903 (2005).

Quantum reflection=

Reflection from the **fast changing attractive** potential

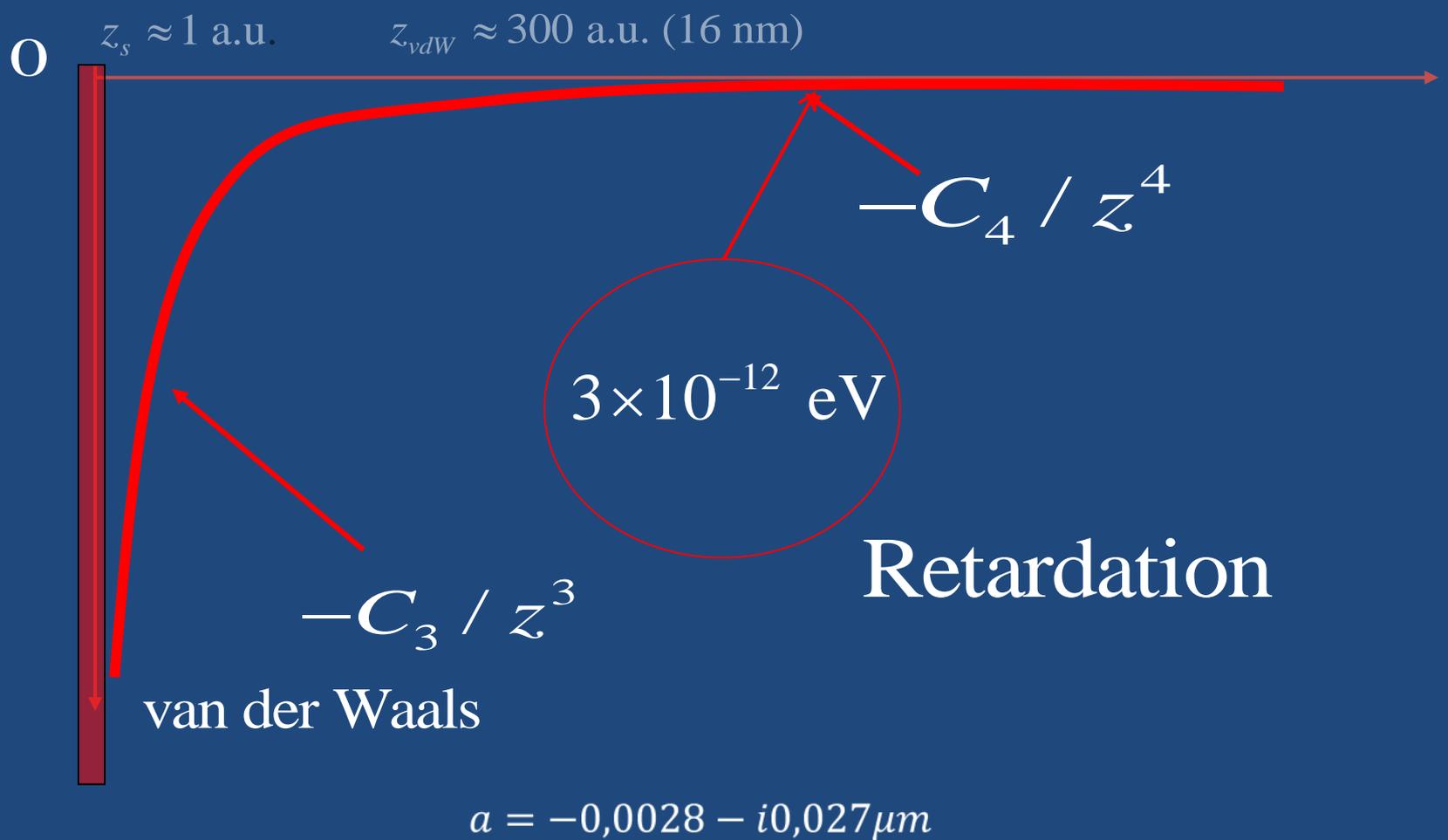
$$\frac{d\lambda_B(z)}{dz} \geq 1; \quad \lambda_B(z) = \frac{2\pi\hbar}{\sqrt{2m(E - V(z))}}$$

$$\frac{d\lambda_B(z)}{dz} = \frac{m\lambda_B^3}{\hbar^2} \frac{dV(z)}{dz}$$

$$\Psi(z \rightarrow \infty) = e^{-ikz} - S e^{ikz}$$

$$\lim_{k \rightarrow 0} S = 1 - 2ika \quad a = \text{Re } a - i |\text{Im } a|$$

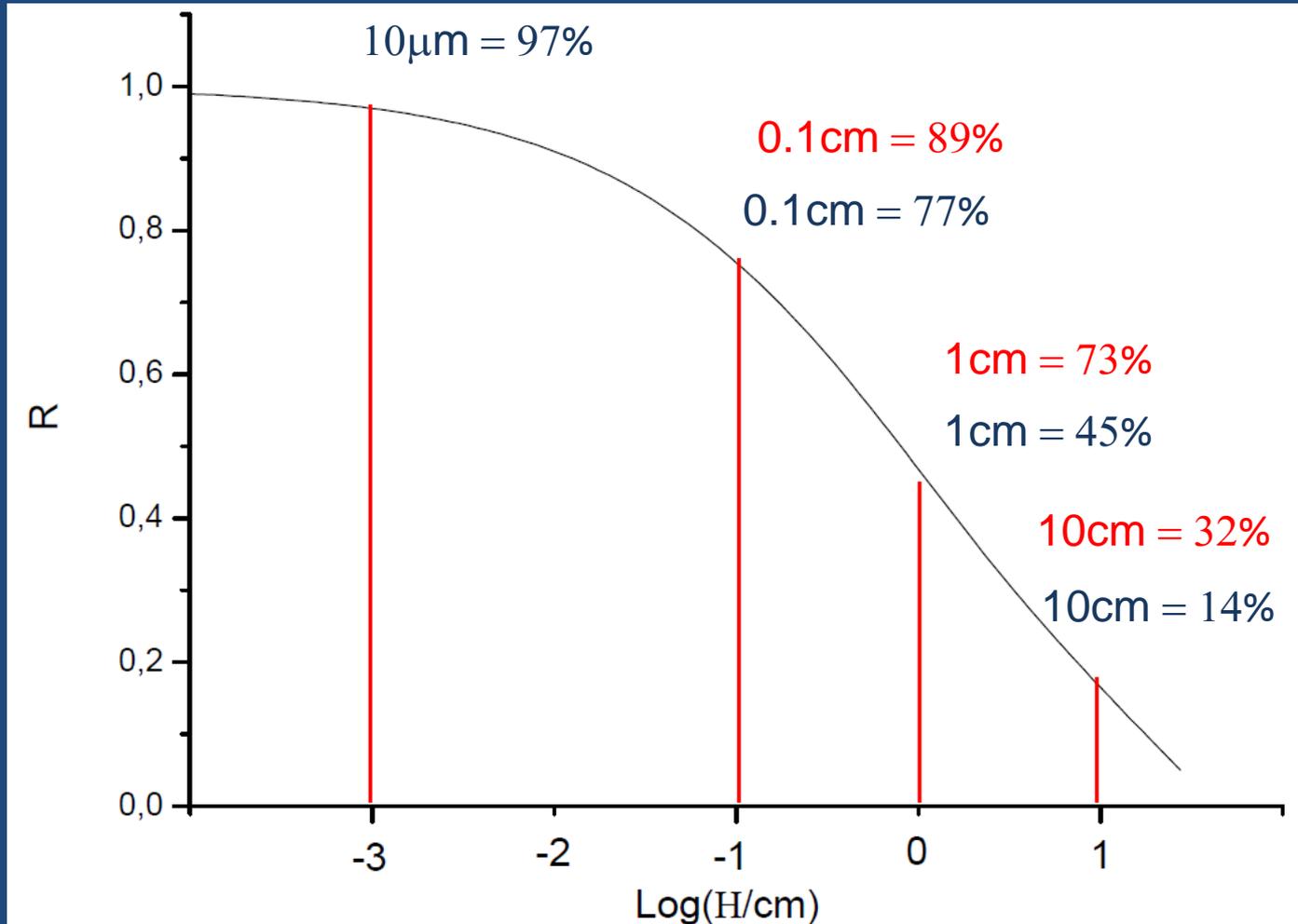
$$\lim_{k \rightarrow 0} R_{\text{qr}} \equiv |S|^2 = 1 - 4k |\text{Im } a| \rightarrow 1; \quad P_{\text{ann}} = 4k |\text{Im } a| \rightarrow 0$$



CASIMIR-POLDER POTENTIAL

$$P_{\text{ann}} = 4k |\text{Im } a| \rightarrow 0$$

From which height can we drop antihydrogen?



Gravitational states

Antihydrogen bouncing on a surface

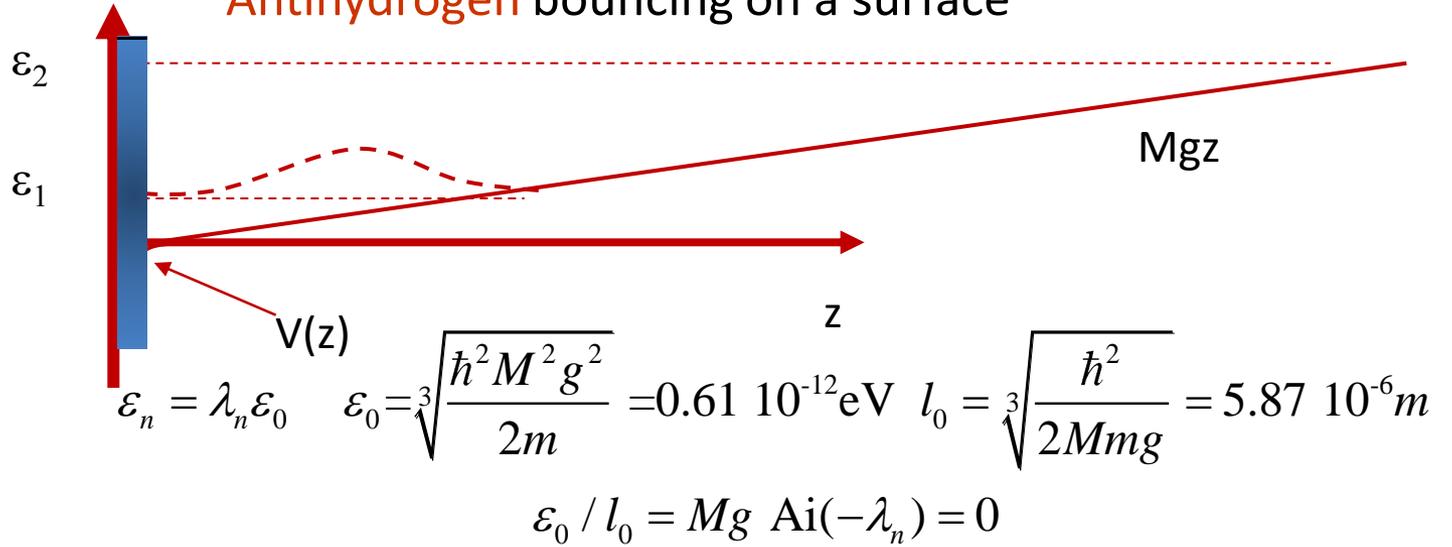
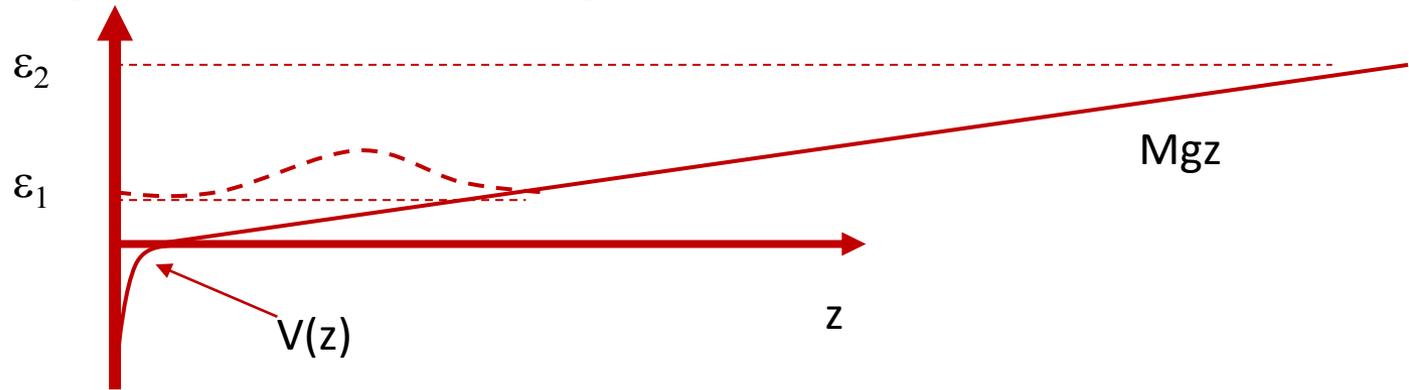


TABLE I. The eigenvalues, gravitational energies, and classical turning points of a quantum bouncer with the mass of (anti)hydrogen in the Earth's gravitational field.

n	λ_n^0	E_n^0 (peV)	z_n^0 (μm)
1	2.338	1.407	13.726
2	4.088	2.461	24.001
3	5.521	3.324	32.414
4	6.787	4.086	39.846
5	7.944	4.782	46.639
6	9.023	5.431	52.974
7	10.040	6.044	58.945

Correction by Casimir-Polder potential + annihilation



$$|a| \approx \sqrt{2mC_4 / \hbar} ; l_0 = \sqrt[3]{\hbar^2 / (2Mmg)} ;$$

$$|a| / l_0 \approx 0.005$$

Correction by Casimir-Polder $V(z)$: $\varepsilon_n = \varepsilon \tilde{\lambda}_n$

$$\tilde{\lambda}_n = \lambda_n + a / l_0$$

$$\varepsilon_n = \varepsilon_0 (\lambda_n + \text{Re } a / l_0) \quad \Gamma = 2\varepsilon_0 |\text{Im } a| / l_0$$

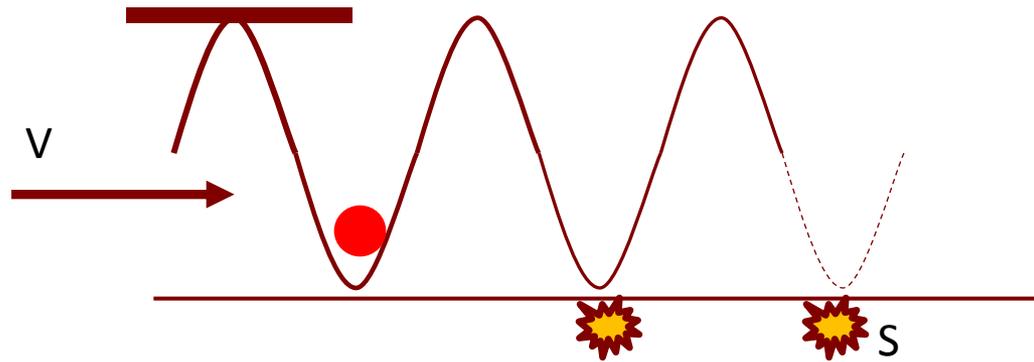
$$\tau = \frac{l_0}{\varepsilon_0} \frac{\hbar}{2|\text{Im } a|} = \frac{\hbar}{2Mg|\text{Im } a|} \approx 0.1s$$

Lifetime is determined by gravitational force Mg and $|\text{Im } a|$.

Numbers: $\varepsilon_1 = 1.43 \text{ peV}$; $\varepsilon_2 = 2.49 \text{ peV}$; $\tau \approx 0.1s$

Antihydrogen “clock”

temporal-energy properties of gravitational states



$$\Phi(z, t) = \varphi_1(z) \exp(-i\varepsilon_1 t - \Gamma t / 2) + \varphi_2(z) \exp(-i\varepsilon_2 t - \Gamma t / 2)$$

$$\varphi_k(z) \sim Ai(z - \lambda_k - a / l_0) \quad \varepsilon_k = \varepsilon_0 \lambda_k \quad \Gamma = 2\varepsilon_0 |\text{Im } a| / l_0$$

$$\langle \varphi_k | \varphi_n \rangle = 2i \frac{\text{Im } a}{l_0} \frac{1}{\lambda_k - \lambda_n + 2i \text{Im } a / l_0} \neq \delta_{kn}$$

$$N_{ann}(t) = -\frac{\partial}{\partial t} \int_0^\infty |\Phi(z, t)|^2 dz = 2\Gamma \exp(-\Gamma t) \left[1 + \cos((\varepsilon_2 - \varepsilon_1)t / \hbar) \right]$$

Bouncing Antihydrogen 2 states

$$\omega_{12} = (\lambda_2 - \lambda_1) \varepsilon_0 = 254.54 \text{ Hz}$$

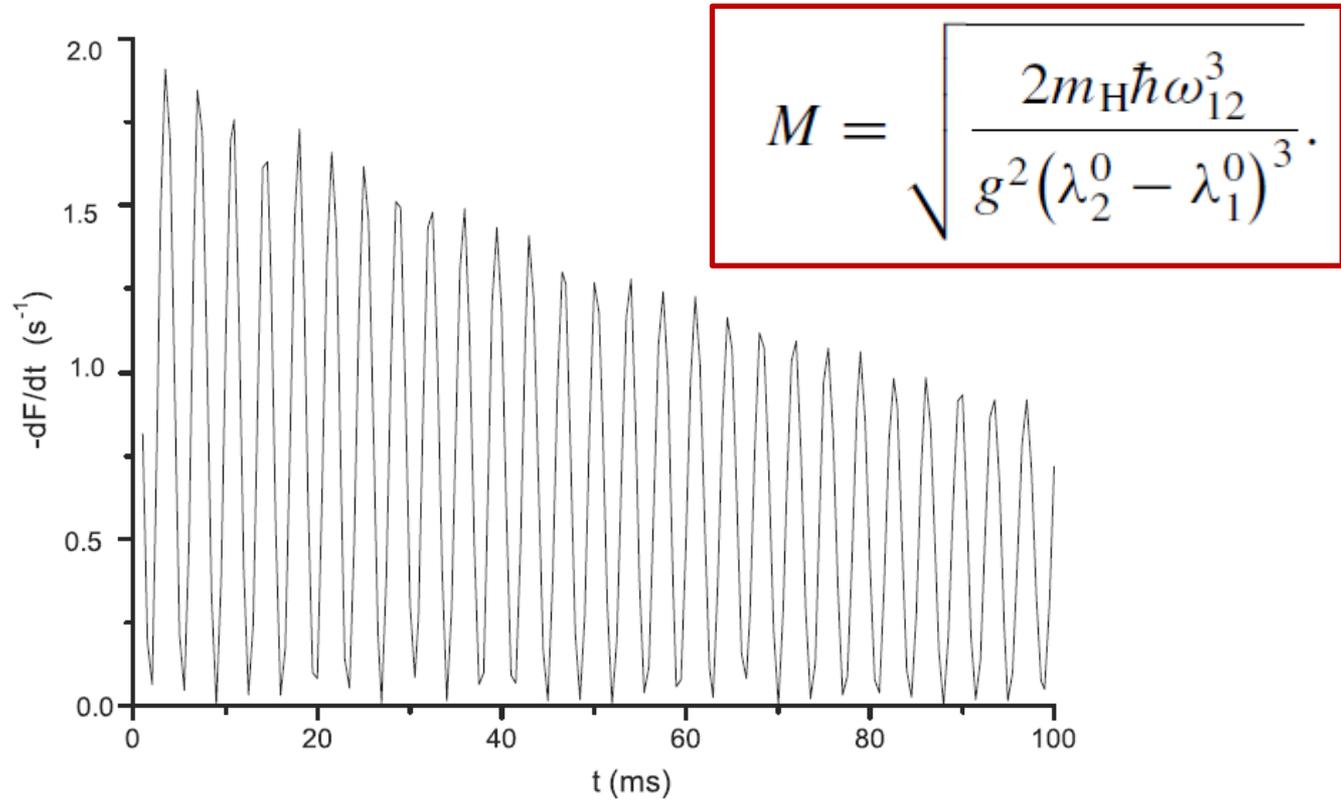


FIG. 2. Evolution of the annihilation rate of $\bar{\text{H}}$ atoms in a superposition of the first and second gravitational states.

RESONANCE SPECTROSCOPY of Gravitational States

Studied by V.Nesvizhevsky et. al in connection with neutrons

E. Kupriyanova MEPHI- magnetic field induced transitions in antihydrogen

Induced resonance transitions :

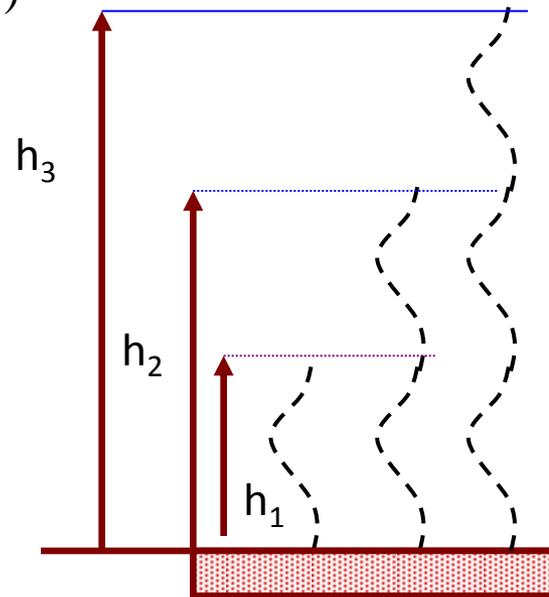
- Gradient magnetic field $\frac{\partial B}{\partial z} = B_0 \cos(\omega t)$
- Oscillating bottom mirror $\frac{\partial h}{\partial t} = a \cos(\omega t)$

$$\omega_{21} = 254.54 \text{ Hz} \quad h_1 = 13.7 \mu\text{m}$$

$$\omega_{31} = 462.83 \text{ Hz} \quad h_2 = 24.0 \mu\text{m}$$

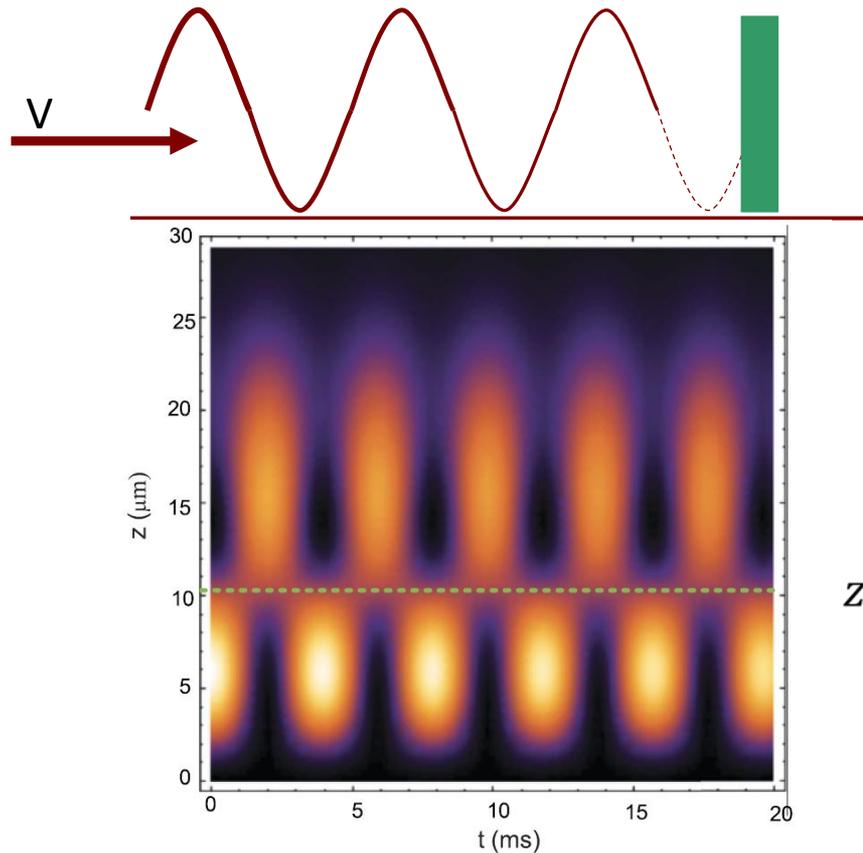
$$\Gamma_{an} \approx 1.6 \text{ Hz} \quad h_3 = 32.4 \mu\text{m}$$

$$M = \sqrt{\frac{2m_H \hbar \omega_{12}^3}{g^2 (\lambda_2^0 - \lambda_1^0)^3}}$$



Quantum ballistic experiment

spatial properties of gravitational states



$$z_{1,2} = (\lambda_2 - \lambda_1) l_0 = 10.27 \mu\text{m}$$

$$M = \frac{\hbar^2 (\lambda_2 - \lambda_1)^3}{2 g m_H z_{1,2}^3}$$

FIG. 4. (Color) The probability density of \tilde{H} in a superposition of the first and second gravitational states, as a function of the height z above the mirror (vertical axis) and the time t (horizontal axis). Dark shade, low probability density; light shade, high probability density. The dashed line indicates the position of the node in the wave function of the second state.

ВЫВОДЫ

- УХН в гравитационном поле- инструмент прецизионных исследований фундаментальных взаимодействий
- Спектроскопия и интерференция гравитационных состояний антиводорода- квантовые измерения гравитационной массы антиводорода